



Cite this: *Mater. Adv.*, 2024,
5, 3904

Received 11th January 2024,
Accepted 5th March 2024

DOI: 10.1039/d4ma00029c

rsc.li/materials-advances

High-throughput screening of stable sulfide semiconductors for solar cell conversion†

Jinjin Yang,^a Zhongxiong Sun,^a Dao-Xin Yao^b and Man-Rong Li^c *^{ac}

The application of chemically stable oxide perovskite solar cells confronts substantial obstacles, such as a large band gap and high internal voltage loss, despite the rapid technical development of device fabrication. In contrast, the design and screening of innovative sulfide materials, which is less explored compared with the oxide counterparts, has become a burgeoning research area. Here, screening for new solar cell absorber materials amongst $\text{RE}_3\text{BB}'\text{S}_7$ ($\text{RE} = \text{La}, \text{Y}$; $\text{B/B}' = \text{Ti-Cu}$ in the 3d row) is performed via high-throughput calculations. The filtration criteria are thermodynamic/kinetic stability, suitable band gap, carrier effective mass, defect tolerance and optical absorption behavior, which are key parameters to determine whether the material is suitable for semiconducting solar cell absorbers or not. Based on the screening criteria, we identified two appealing candidates, $\text{RE}_3\text{CoCrS}_7$ ($\text{RE} = \text{La}, \text{Y}$), with potential application in photoelectrochemical devices. These findings established a solid theoretical foundation for the development and implementation of sulfide solar cells in practical applications.

1. Introduction

Efficient harvesting of solar energy is urgently required to solve the impending energy and environmental crisis. A prerequisite is the screening of materials that have high-performance solar energy conversion.^{1–3} Perovskite materials have demonstrated advantages as light absorbers in solar cells, owing to their outstanding stability in highly oxidizing circumstances, widely tunable band gaps (E_g), high optical absorption rates, long carrier diffusion length, and high carrier mobility. However, two insurmountable problems exist in perovskite oxides: (1) the large E_g for sunlight absorbers due to the low energy of the O 2p orbital contributing to the valence band maximum (VBM), and (2) the anodic VBM with respect to the water redox level resulting in large internal voltage loss in devices.⁴ Fortunately, both aforementioned issues can be resolved by substituting oxygen with sulfur. Because sulphides can possess relatively high conduction band positions suitable for H_2O reduction and better sunlight responses than oxides, which are originated from the higher valence band positions composed of S 3p orbitals. In addition, sulfide perovskites are a more

environmentally friendly option than lead halide analogs, and reduce instability and toxicity commonly associated with halide perovskites. Semiconducting perovskite sulfides have been theoretically predicted and experimentally synthesized, showcasing superior electrical and optical qualities that make them ideal for low-cost tandem solar cells. A number of sulfides, such as AZrS_3 ($\text{A} = \text{Ca}, \text{Sr}, \text{Ba}$), SrTiS_3 , CuTaS_3 , CaSnS_3 and SrSnS_3 , have attracted intense interest in the context of photovoltaic (PV) solar cells.⁵ Recently, Ruddlesden-Popper (R-P) sulfides $\text{A}_3\text{BB}'\text{S}_7$ (also known as $\text{A}_3\text{B}_2\text{S}_7$ if B and B' are the same elements) have theoretically and experimentally received great attention as solar conversion materials.⁶ $\text{A}_3\text{Zr}_2\text{S}_7$ ($\text{A} = \text{Ca}, \text{Ba}$) and $\text{Ca}_3\text{Sn}_2\text{S}_7$, which demonstrate stable ferroelectricity at room temperature with small E_g values (< 2.2 eV), are well suited for significant visible-light absorption. The R-P type $\text{Ca}_3\text{Zr}_2\text{S}_7$ material with low E_g suitable for visible light absorption is anticipated to exhibit stable ferroelectric polarization as a promising ferroelectric photovoltaic material along with high energy conversion efficiency (PCE), where photo-generated electron-hole pairs are effectively separated.⁷ Moreover, the charge recombination rate in $\text{Ca}_3\text{Sn}_2\text{S}_7$ can be experimentally tuned to simulate its superior infrared detection performance, showing small carrier effective mass ($0.04 m_0$), ultrahigh room-temperature carrier mobility ($6.7 \times 10^4 \text{ cm}^2 \text{ V}^{-1} \text{ s}^{-1}$), Fermi velocity ($3 \times 10^5 \text{ m s}^{-1}$), and optical absorption coefficient (10^5 cm^{-1}).^{8,9} Also, the strained ferroelectric $\text{Ba}_3\text{Zr}_2\text{S}_7$ thin film possesses a suitable E_g value of 1.6 eV for PV application, absorbing sunlight in a wide range of visible-light spectra. Thus, $\text{A}_3\text{BB}'\text{S}_7$ sulfides represent a novel and cost-effective material class for solar cells, offering greater stability than oxide counterparts.¹⁰

^a Key Laboratory of Bioinorganic and Synthetic Chemistry of Ministry of Education, School of Chemistry, Sun Yat-Sen University, Guangzhou 510006, P. R. China.
E-mail: limanrong@mail.sysu.edu.cn

^b Guangdong Key Laboratory of Magnetoelectric Physics and Devices, School of Physics, Sun Yat-Sen University, Guangzhou 510275, P. R. China

^c School of Chemistry and Chemical Engineering, Hainan University, Haikou 570228, P. R. China

† Electronic supplementary information (ESI) available. See DOI: <https://doi.org/10.1039/d4ma00029c>



In rare earth (RE) ions, the abundant 4f/5d orbit electrons could generate different intermediate energy states to induce unique photo-physical behavior, suggesting excellent potential in PV based devices. Accordingly, the electronic properties of solar cell materials can be modulated by appropriate RE ions.¹¹ In particular, the incorporation of RE ions has a major effect on the charge carrier mobility, band structure, and absorption/luminescence properties, resulting in modified charge transport dynamics as exemplified in perovskite-related solar cells.¹² More significantly, the up-conversion takes place in materials containing trivalent RE ions by the process of ground- and excited-state absorption of multi-photons from actual energy levels.¹³ In general, when utilizing trivalent RE doped materials as up-converting layers, a single junction connected to an ideal up-conversion layer can achieve a maximum PCE of approximately 47.6% for non-concentrated sunlight and 63.2% for concentrated sunlight. For example, Er/Yb doped M_2O_2S ($M = Gd, Y, La$) are used as a reference to show the enhancement of up-conversion emission efficiency under 1550 nm excitation,^{14,15} and Tb^{3+} - Yb^{3+} co-doped silicate glass can improve the PCE of monocrystalline silicon solar cells by 8.6%.¹⁴ Along these lines, $RE^{3+}_3BB'S_7$ are expected to be excellent light absorbers and enhance the PCE of solar cells, considering their chemical and structural similarities with the reported solar energy conversion materials. Ruled by the charge balance, the valence state of $B-B'$ in $RE_3BB'S_7$ can be either monovalent-tetravalent (I-IV) or bivalent-trivalent (II-III) combinations. So far, the most prevalent cases adopted the monovalent-tetravalent (I-IV, $B = Cu, Ag$; $B' = Si, Ge, Sn$)

combinations in $RE^{3+}_3B^{1+}B'^{4+}S_7$, and mostly crystallized in the La_3SiCuS_7 -type $P6_3$ hexagonal structure, such as Y_3CuGeS_7 and Sm_3CuGeS_7 identified as outstanding infrared nonlinear optical materials. In contrast, the bivalent-trivalent (II-III) combination $RE^{3+}_3B^{2+}B'^{3+}S_7$ has barely been studied as a research object of this work, leaving a vast chemical space to explore new candidates. Structural, magnetism and electrical transport data for $La_3B^{II}Fe^{III}S_7$ ($B = Mn, Fe, Co, Ni$) have been experimentally reported, providing experimental validation of upcoming calculations.¹⁶ $La_3B^{II}Fe^{III}S_7$ crystallizes in the non-centrosymmetric hexagonal space group $P6_3$ with $Z = 2$, where crystal structures consist of parallel chains of face-sharing B-centred octahedra and stacking of Fe-centred tetrahedra with seven-coordinated La atoms in the intervening spaces. Based on previous magnetic and electric tests, $La_3B^{II}Fe^{III}S_7$ ($B = Mn, Fe, Co, Ni$) sulfides are p-type semiconductors and have higher magnetic moments at high-temperatures than those at low temperatures, except for the nickel compound, because of the probable interaction between octahedral site B and tetrahedral site Fe.¹⁶

The performance of solar cell light absorbing materials depends not only on the process of device fabrication, but also on the electrical and optical properties of the original materials. However, it is difficult and costly to measure the microscopic properties of materials experimentally. There is a need for theoretical prescreening to reveal the intrinsic mechanisms and inherent properties of new potential candidates for light absorption layers in solar cells. Recently, data-driven and high-throughput computing (HTC) assisted approaches

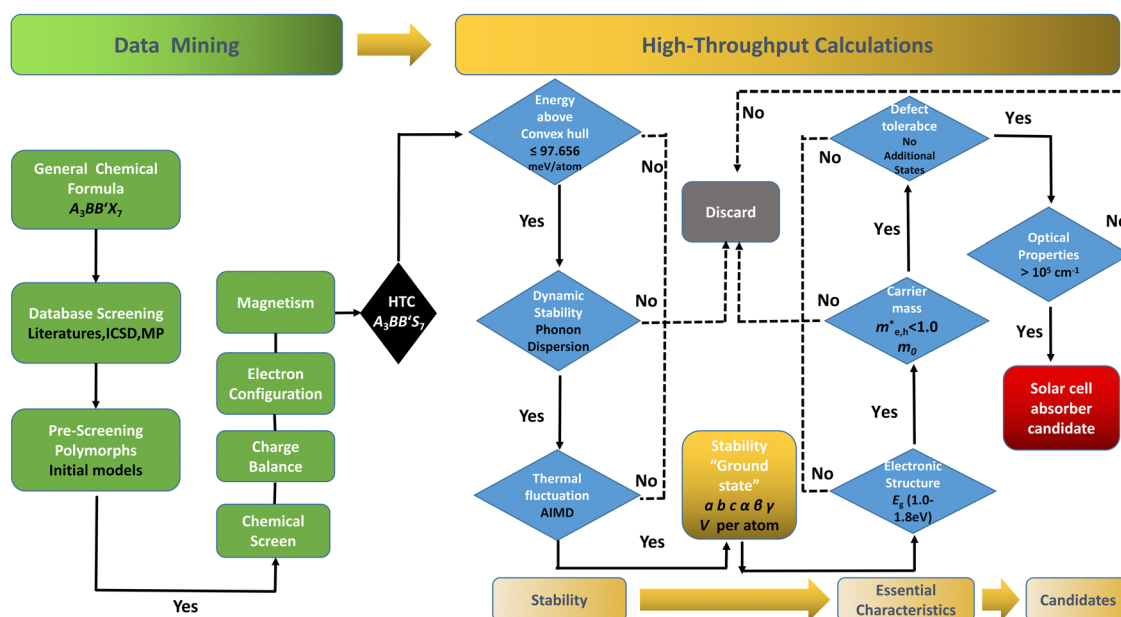


Fig. 1 Workflow of the data mining and HTC procedure to screen solar cell absorber materials. Data mining: data mining for the $A_3BB'S_7$ general formula to find possible polymorphs with confinement(s) by geometric restraints (ionic radius, tolerance factor, distortion, tilting, etc.) and defined characters (magnetism, spin-orbit coupling, charge balance, electron configuration, etc.), and then execute chemical screening to build initial structural models for calculations. HTC: formation energy convex hull screening in Materials Project ($E_{\text{hull}} \leq 97.656$ meV per atom), and the ones with targeted E_g values of 1.0–1.8 eV with the GGA+ U approach remain in the funnel. The mobility of electrons and holes is considered by calculating the effective mass ($m_{e,h}^*$). The defect tolerance is studied to confirm that any potential defect state is not contributing to deep gap recombination sites.

have facilitated significant time and cost savings. The key points are to explore the correlations of band structure, light absorption, carrier mobility and defect tolerance to guide and accelerate the discoveries of new light absorption layer materials. In this work, we present a comprehensive screening of 128 possible candidates in $\text{RE}^{3+}_3\text{B}^{3+}\text{B}'^{2+}\text{S}_7$ (RE = La, Y; B and B' are Ti–Cu in the 3d row of the periodic table of elements) to attentively seek suitable solar cell absorption materials. We perform density functional theory (DFT) based HTC, considering formability of 672 different $\text{RE}^{3+}_3\text{B}^{3+}\text{B}'^{2+}\text{S}_7$ candidates in seven potential polymorphs along three magnetic configurations. The screening workflow is displayed in Fig. 1, including data-mining and HTC. It should be noted that some compounds in the $\text{RE}^{3+}_3\text{B}^{3+}\text{B}'^{2+}\text{S}_7$ system have been experimentally reported,^{17–23} which boosts the examination and validation of suggested approaches through the initial computation test.

Our screening unveils a short list of potential thermodynamically and kinetically stable candidates with suitable E_g for solar cell absorbers, giving two final winners. $\text{RE}_3\text{CoCrS}_7$ (RE = La, Y) with appropriate E_g values of 1.02 (La based) and 1.00 (Y based) eV are predicted to have small carrier effective masses of electrons and holes at the VBM and a conduction band minimum (CBM) causing high mobility and low recombination rate of photo-generated charge carriers, which are necessary for efficient solar cell conversion. The proposed two candidates are identified to be tolerant to vacancies, since the introduction of defects does not give rise to deep gap recombination sites. A strong light absorption coefficient in the visible region is proposed, as evidenced by their calculated optical absorption.

2. Methods

In the frame of DFT, the Vienna *ab initio* simulation package (VASP)^{24,25} adopted projector-augmented wave (PAW)²⁶ in structural relaxation and electronic structure investigation, and generalized gradient approximation (GGA) implemented in the Perdew–Burke–Ernzerhof (PBE) method was employed. The convergence of cut-off energy, k -points and Gaussian broadening was tested to ensure the accuracy of conclusions. The plane wave cut-off energy was 700 eV. The k -point mesh was $9 \times 9 \times 9$ for monoclinic ($C2/m$ and $P2_1/c$) and tetragonal ($P4_2/mnm$ and $I4/mnm$) symmetries, $7 \times 7 \times 9$ for hexagonal ($P6_3$) symmetry, and $5 \times 3 \times 5$ for orthorhombic $Pbam$ and $Cccm$ symmetry, ensuring that the geometric optimization and static electronic calculations are minimized in the constraint of periodic boundary conditions. The energy and force convergence threshold were set as 1.0×10^{-6} eV and 1.0×10^{-2} eV \AA^{-1} , respectively. The exchange correlation energy was treated by GGA-PBE. To lessen the inaccuracy brought by strong on-site Coulomb interactions, electron–electron Coulomb repulsion interactions (U) for 3d transition metal elements were considered in the rotationally invariant form (GGA+ U), which go beyond the conventional local-density approximation. Standard Hubbard U corrections were adopted for Ti (4.0 eV), V (3.25 eV), Cr (3.7 eV), Mn (3.9 eV), Fe (5.3 eV), Co (3.32 eV), Ni (6.2 eV) and Cu (7.0 eV) as

discussed in detail in literature studies and the existing Materials Project (MP) database.²⁷ The given standard set of Hubbard U corrections would be suitably transferrable to magnetic characteristics.^{27,28} The Heyd–Scuseria–Ernzerhof (HSE06) functional was employed to systematically improve the calculation accuracy for the electronic structure. The phonon dispersion was calculated using the Phonopy package with the finite displacement method along the main symmetry directions in a three-dimensional Brillouin zone.²⁹

3. Results and discussion

3.1 Data mining and initial computing models

The distribution of preliminarily screened energy above convex hull (E_{hull}) of formation energy and PBE E_g for 203 known $\text{A}_3\text{B}_2\text{S}_7$ and $\text{A}_3\text{BB}'\text{S}_7$ chemical formulae found in the inorganic crystal structure database (ICSD), Materials Project (MP) database and reported literature as our initial database is plotted in Fig. 2(a). Obviously, compounds with higher E_{hull} values are more unstable and are anticipated to be metallic. To further motivate the choice of structures and chemical combinations considered in the calculations based on the known $\text{A}_3\text{B}_2\text{S}_7$ and $\text{A}_3\text{BB}'\text{S}_7$ from data mining,^{22,30–32} seven possible polymorphs as initial structural models for $\text{A}_3\text{BB}'\text{S}_7$ are shown in Fig. 2(b). The corresponding lattice parameters, perovskite-related tolerance factors, and basic properties of representative compounds in each polymorph are summarized in Table S1 (ESI†). As shown in Fig. 2(b)-i: (1) the monoclinic $C2/m$ layered structure, as found in $\text{Y}_3\text{Y}_2\text{S}_7$ ($A = B = B' = Y$) where the face-sharing pyramidal BS_5 extends along the b -axis forming a linear chain.³³ (2) The monoclinic $P2_1/c$ is exemplified by $\text{Ba}_3\text{Sn}_2\text{S}_7$ (Fig. 2(b)-ii).³⁴ The corner-sharing BS_4 tetrahedra run along the b axis. (3) Orthorhombic $Pbam$ as shown in $\text{Eu}_3\text{Sn}_2\text{S}_7$ (Fig. 2(b)-iii),³⁵ where the corner-shared BS_4 forms a one-dimensional chain along the c axis. (4) The hexagonal $P6_3$ as reported in $\text{La}_3\text{SiCuS}_7$ (see Fig. 2(b)-iv) consists of six-coordinated B cations aligned along the c axis, forming a chain structure by face-sharing BS_6 octahedra.²¹ Hexagonal $P6_3$ is the most frequently identified structure in contrast to the other six rare ones. (5–7) Three perovskite-type structures related to the R–P phase are shown in Fig. 2(b)-v, vi, vii, where the crystal structures are exemplified by three distinct polymorphs of $\text{Ba}_3\text{Zr}_2\text{S}_7$,^{8,36,37} namely orthorhombic $Cccm$ (Fig. 2(b)-v), tetragonal $P4_2/mnm$ (Fig. 2(b)-vi), and $I4/mmm$ (Fig. 2(b)-vii), respectively. These series consist of perovskite structure blocks separated by extra rock-salt layers along the [001] direction. The orthorhombic $Cccm$ is a distorted variant related to the perovskite structure,³⁸ while the tetragonal $I4/mmm$ with an ideal body-centered structure can be regarded as double layers of BS_6 octahedra along the c -axis separated by rock-salt layers. In the slightly distorted tetragonal $P4_2/mnm$ structure, the perovskite slabs and rock-salt layers stagger along the [001] direction. Within the perovskite layers, the octahedra are corner-shared along [110] and $[1-10]$ and tilted along one of the in-plane directions. Here, the above seven structural prototypes are the most likely



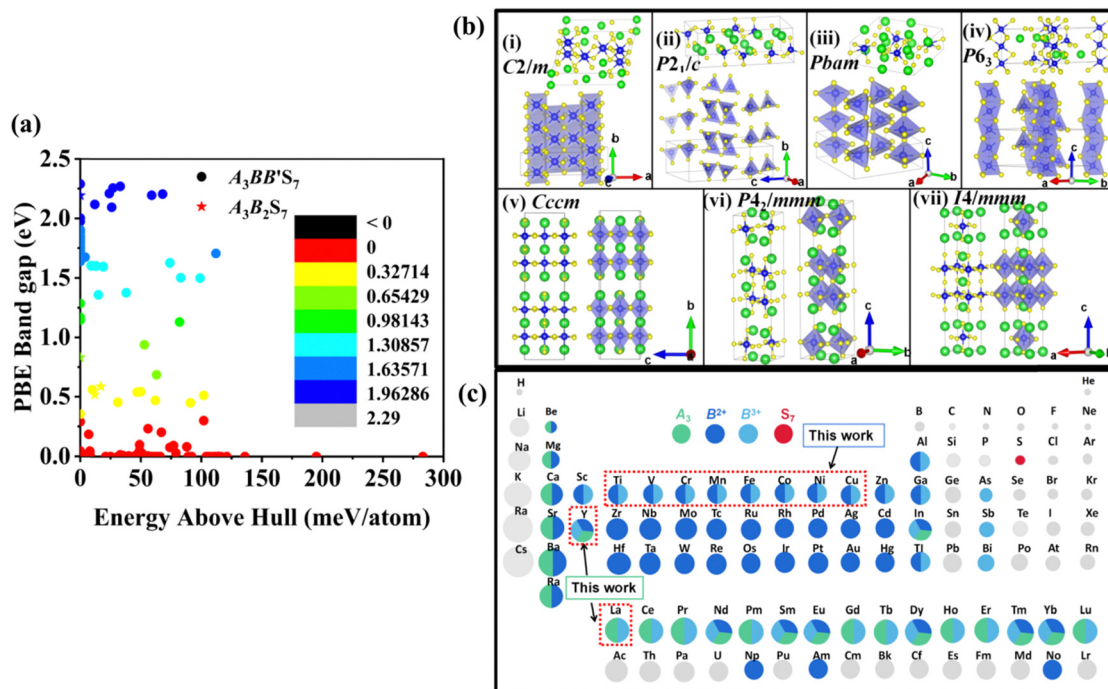


Fig. 2 Data mining and initial computing models. (a) Distribution of calculated energies above E_{hull} and PBE E_g . The scatter plot illustrates the distribution of calculated energies above E_{hull} and E_g values for 203 $A_3B_2S_7$ and $A_3BB'S_7$ compounds. (b) The most common prototypes for the $A_3BB'S_7$ chemical formula by data mining. The unit cell structure and local polyhedral connection are addressed for clarity. (c) Periodic table of the elements considered in $A_3B^{2+}B'^{3+}S_7$ with A, B^{2+} , B'^{3+} and S sites marked in green, dark blue, light blue and red, respectively. If one element is half or one third colored, it means that it can be at two or three sites.

polymorphs for $A_3BB'S_7$, and thus are adopted as the initial structure models in subsequent HTC.

Chemical screening of all possible polymorphs of $(A^{3+})_3(B^{2+})(B'^{3+})(S^{2-})_7$ chalcogenides, where A, B, and B' can be 23, 48, and 30 conceivable elements under a charge balance, respectively, and the combinations of theoretical 33 120 ($23 \times 48 \times 30$) candidates are shown in Fig. 2(c). According to charge neutrality, the sum of cationic oxidation states should be +14 corresponding to the value of −14 from the sulfur anion. The targeted 128 $(RE^{3+})_3(B^{2+})(B'^{3+})(S^{2-})_7$ compounds ($RE = \text{La, Y; B/B}' = \text{Ti-Cu}$ in the 3d row in the periodic table) over the chemical space are restricted by the charge balance and geometrical confinement, as seen in Fig. 1 (data mining). So far, any significant portion of virtual chemical combinations has not been studied yet. Thus, it is urgent to implement further validation for the synthesizable and practical chemical combinations following the workflow in Fig. 1 and Fig. S1 (ESI†).

The contribution of spin-spin interaction (magnetic structure) in 3d transition metals has proven to be critical to calculate the minimum energy for each initial structural model.^{39–41} Determining the magnetic configuration is a highly nontrivial task, and one can only locate the magnetic ground state within the space of a limited number of initial configurations. In this work, we restrict the collinear magnetic structures to feature HTC, since there are numerous possible noncollinear magnetic structure models, making it impossible for any comprehensive calculations. In fact, the energy difference

between the collinear and noncollinear magnetic structures is found to be insufficient to modify the overall results.^{42,43} In $RE_3BB'S_7$ ($RE = \text{La, Y; B/B}' = \text{Ti-Cu}$ in the 3d row), we take the cationic arrangements and selected magnetic 3d B/B' cations into account to match expected correlations. For the R-P phase *Cccm*, *I4/mmm* and *P4₂/mnm*, three-type arrangements of the magnetic moments are considered: ferromagnetic (FM) and two antiferromagnetic (AFM) structures, including FM with all moments ferromagnetically aligned, AFM1 with the FM alignment within BS_6 layers and AFM coupling between layers, AFM2 with entirely AFM both within and between BS_6 layers. For the hexagonal $P6_3$ structure, there are three possible magnetic configurations, including the FM ($B_1 \uparrow B_2 \uparrow B'_1 \uparrow B'_2 \uparrow$), AFM1 ($B_1 \uparrow B_2 \downarrow B'_1 \downarrow B'_2 \uparrow$) and AFM2 ($B_1 \downarrow B_2 \downarrow B'_1 \uparrow B'_2 \uparrow$). For other polymorphs, the FM and multiple AFM magnetic configurations are established. The seven screened structures are endowed with spin configurations as shown in Fig. S2 (ESI†). Partial structural optimization energies of $\text{La}_3\text{B}_2\text{S}_7$ ($B = \text{Ti-Cu}$) are listed in Table S2 (ESI†) to find the lowest energy of the magnetic structure. Since almost all known $A_3BB'S_7$ ($B \neq B'$) compounds crystallize in the hexagonal $P6_3$ structure based on extensive data screening and literature reports,^{17–19,21,30,44} the potential candidates most likely lie in this symmetry. To save computing resources, for the 56 $\text{La}_3\text{BB'S}_7$ ($B \neq B'$) compounds, there are 168 magnetic structures [56×1 (hexagonal $P6_3$) $\times 3$ (three kinds of magnetic configuration)] to be calculated. In the 8 $\text{La}_3\text{BB'S}_7$ ($B = B' = \text{Ti-Cu}$ in the 3d row) cases, 168 (8×7 (seven



structural models) $\times 3$ (three kinds of magnetic configuration)) magnetic structures are examined. So, $\text{La}_3\text{BB}'\text{S}_7$ totally generates 336 possible magnetic structures. A similar situation occurs in 8 $\text{Y}_3\text{BB}'\text{S}_7$ ($\text{B} = \text{B}' = \text{Ti-Cu}$ in the 3d row) cases. Next, the 672 (336 (La-based) + 336 (Y-based)) possible magnetic configurations are passed for HTC based DFT. During optimization, the lattice parameters and atomic positions of these structures are set to be completely relaxed relative to the original structure. All internal atomic coordinates have been relaxed by total energy minimization in our calculations to find the most stable structure for each chemical composition. If the initial structure undergoes limited symmetry-breaking displacement during structural optimization, then in some cases the crystal structure undergoes dramatic changes.

3.2 HTC

For more predictive and affordable screening results, the chemically screened dynamics and thermodynamically stable compounds are initially calculated in all possible polymorphs over the whole chemical space under the charge balance. Only the stable combinations with targeted suitable E_g remain, and the mobility of an electron (e) and a hole (h) is also taken into account by calculating the carrier effective mass ($m_{\text{e,h}}^*$). Finally, the defect tolerance of surviving compounds is studied to distinguish the potential defect state that does not contribute to deep-gap recombination sites.⁴⁵ The distribution of the

absorption spectrum in the visible region indicates the absorption intensity of potential $\text{RE}_3\text{BB}'\text{S}_7$ ($\text{RE} = \text{La, Y}$; $\text{B/B}' = \text{Ti-Cu}$ in the 3d row) solar cell absorber candidates to achieve the best PCE.

3.2.1 Stability

3.2.1.1 Formability filtration. It is instructive to analyze the energy distance of the most stable polymorph to the energy convex hulls E_{hull} constituted by the stable compounds ($E_{\text{hull}} = 0$ meV per atom) approved in the MP database. Specifically, the thermodynamic stability is evaluated by considering all reactions in which the compound decomposes into stable components. Only the compounds with E_{hull} lower than thresholds are supposed to be synthesizable.⁴⁶ In order to ensure the reported $\text{La}_3\text{BFe}^{\text{III}}\text{S}_7$ ($\text{B} = \text{Mn, Fe, Co, Ni}$) compounds locate on or even below the convex hulls, 97.656 meV per atom for $\text{La}_3\text{FeMnS}_7$, the maximum value of known series in this category, is set as the threshold for screening potential thermodynamically stable or metastable compounds. It should be pointed out that the MP database does not contain all possible decomposition products for some compounds, meaning that the calculated energy above E_{hull} serves only as a lower boundary and actual thermodynamic stability could be lower than predicted. Detailed calculation results of the 128 $\text{RE}_3\text{BB}'\text{S}_7$ in the magnetic polymorph ground state are listed in Tables S3 and S4 (ESI[†]), and the cases including on/below E_{hull} are displayed in Fig. S1 (ESI[†]) (HTC 1). As shown in Fig. 3(a) and (b), the instability considerably increases with the lengthened energy distance to the

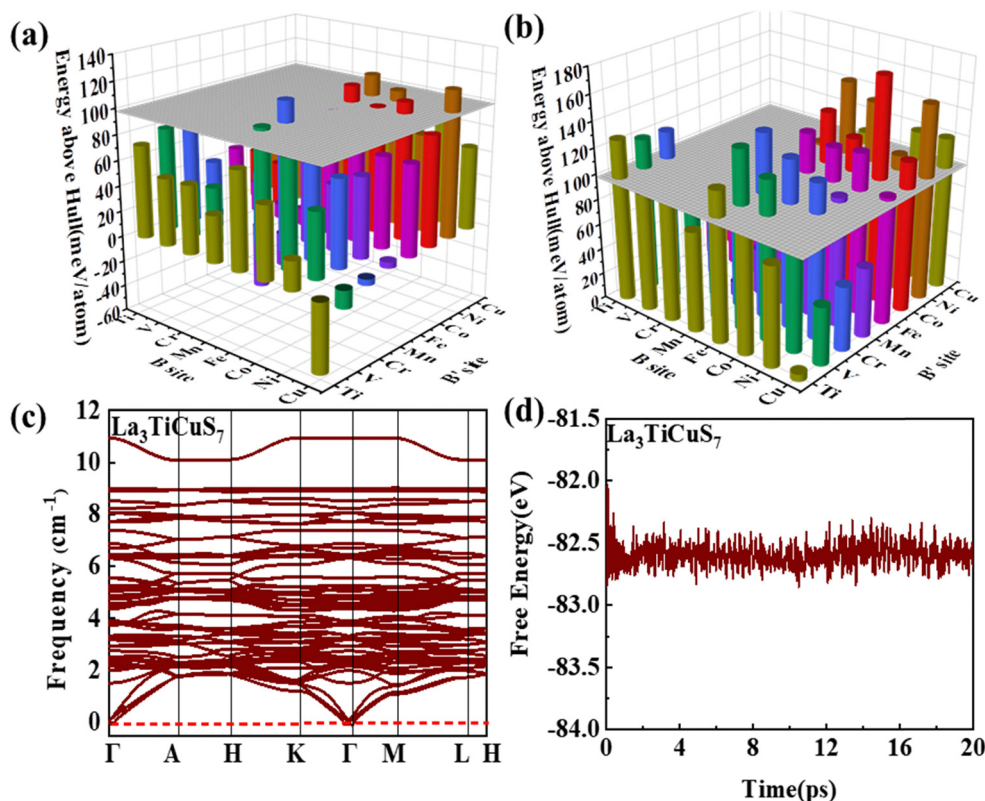


Fig. 3 The stability of the considered combinations. Distribution of the distances to the convex hull of (a) $\text{La}_3\text{BB}'\text{S}_7$ and (b) $\text{Y}_3\text{BB}'\text{S}_7$. (c) Phonon dispersion and variations of representative $\text{La}_3\text{TiCuS}_7$. (d) The total potential energy with respect to the simulation time during AIMD simulations of $\text{La}_3\text{TiCuS}_7$.



convex hull. Employing the boundaries on energy above hull narrows down the search space for synthesizable 94 (128 minus 34) candidates, including 56 La- and 38 Y-based compounds on or below the threshold, indicating that they are thermodynamically stable at 0 K. The E_{hull} benchmark rules out 34 (including 8 La- and 26 Y-based compounds) of the 128 combinations, since they dock above the convex hull and are thermodynamically not synthesizable. Preliminary screening on the electronic structure shows that only 22 (11 La- and 11 Y-based) out of the 94 combinations in the perfect structure are non-zero E_g , as seen in Fig. S1 (ESI†) (HTC 2).

3.2.1.2 Kinetic stability and thermodynamic stability. In condensed-matter physics, phonons are used to describe the simple harmonic vibrations of lattices. Thereby, the phonon spectrum is a good characteristic tool to evaluate the kinetic stability evaluated from finite-difference phonon calculations.²⁹ When the phonon spectrum is all above 0 cm^{-1} , it means that the material has no imaginary frequency and is relatively stable. We subsequently investigate the kinetic stability of these 22 combinations with non-zero E_g . The phonon spectra of $\text{La}_3\text{TiCuS}_7$ without the virtual frequency is exemplified in Fig. 3(c). Partial phonon dispersion of selected stable partial sulfides without an imaginary frequency are shown in Fig. S3 (ESI†). Among the 22 calculated combinations, 18 (8 La-based and 10 Y-based) compounds are found to have no imaginary phonon modes, being dynamical lattice stable at $T = 0$ K. $\text{La}_3\text{MnCoS}_7$, $\text{La}_3\text{V}_2\text{S}_7$, $\text{La}_3\text{Ni}_2\text{S}_7$, and $\text{Y}_3\text{V}_2\text{S}_7$ are discarded because of the strong virtual frequency.

Additionally, for E_{hull} and phonon dispersion calculations it should be noted that these are zero-temperature calculations, and the stability could be altered by finite temperature contributions to the free energy. In other words, the entropy factors in free energy may cause some of the compounds that are anticipated to be unstable at 0 K to turn out to be stable at ambient temperature. The consideration of finite temperature stability can be undertaken as further efforts. Based on *ab initio* molecular dynamics (AIMD) examination,⁴⁷ the energy and temperature in the molecular dynamics simulation process are extracted to prove thermodynamic stability. As the finite energy fluctuates, it is supposed to be relatively stable at 300 K, when the structure is changed without being destroyed. For

example, the AIMD simulation results and variations of the total potential energy for $\text{La}_3\text{TiCuS}_7$ is exemplified in Fig. 3(d), which exhibits a small energy fluctuation and hardly distorted crystal structure. The calculated AIMD results of other partially squeezed stable compounds are shown in Fig. S4 (ESI†). Overall, the 18 (8 La- and 10 Y-based compounds) combinations with finite energy fluctuations and robust structure at 300 K are listed in Table 1 and Fig. S1 (ESI†) (HTC 3). These 18 compounds with desirable stability and non-zero E_g are now passed to the next step to examine the electronic structure with further filtration.

3.2.2 Electronic structure. Governed by the well-known Shockley–Queisser limit, the limit efficiency for a single p–n junction based solar-cell semiconductor is related to the E_g value under standard illumination conditions.^{48,49} When E_g is too large, it is unable to absorb sunlight in the visible light band. If the E_g is too small, the open circuit voltage of the cell will decrease and result in faded efficiency. As solar cell absorber materials, the E_g value is required to be within a specific range of 1.0–1.6 eV.^{50,51} Within this E_g range, the materials can equilibrate the number and the energy obtained by absorbed photons. The E_g values of current mainstream solar cell materials are all within this range. For example, the E_g values of Si, GaAs, and $\text{CH}_3\text{NH}_3\text{PbI}_3$ are 1.1, 1.42, and 1.55 eV, respectively. The maximum limit efficiency appears at an E_g value of 1.34 eV with a PCE of 33.7%. Identifying efficient semiconductors with desirable E_g as solar cell absorption materials is currently regarded as an extremely important challenge. Given the inherent limitations of DFT, we expand the E_g screening range to 1.0–1.8 eV to avoid missing potential candidates. It is well known that the methods that have been most commonly used for successfully estimating the E_g values of semiconducting materials up to this point are GGA+U, HSE06, and GW approximation (GWA). However, the HSE06 functional or GWA cannot handle the computationally intensive computations of defect characteristics, which call for the utilization of massive supercells. As a consequence, the electrical and defect characteristics are computed in this work using GGA+U.

In order to screen the target E_g , we examine the electronic characteristics of the 18 (8 La- and 10 Y-based) stable cases to gauge the band structures and electronic density of states

Table 1 Calculated direct or indirect E_g values (eV) and carrier effective masses $m_{\text{e,h}}^*$ (m_0) of selected 18 combinations. The candidates with bold and italicized fonts meet the screening criteria of E_g values in the range of 1.0–1.8 eV and carrier effective masses of less than 1.0 m_0

Combinations	E_g	m_{e}^*	m_{h}^*	Combinations	E_g	m_{e}^*	m_{h}^*
$\text{La}_3\text{TiCoS}_7$	0.83 (direct)	18.40	0.89	Y_3TiCoS_7	0.69 (indirect)	18.65	0.29
$\text{La}_3\text{TiCuS}_7$	1.44 (indirect)	1.33	1.02	Y_3TiCuS_7	1.59 (indirect)	0.79	0.33
La_3VMnS_7	1.43 (direct)	1.89	1.13	Y_3VMnS_7	1.38 (indirect)	2.00	0.41
$\text{La}_3\text{CrCuS}_7$	0.74 (direct)	0.31	0.24	Y_3CrVS_7	0.70 (indirect)	0.56	0.70
$\text{La}_3\text{MnCrS}_7$	1.07 (direct)	1.00	0.55	Y_3CrMnS_7	0.83 (indirect)	1.21	0.31
$\text{La}_3\text{FeMnS}_7$	1.23 (direct)	0.70	2.50	Y_3CrCuS_7	0.68 (indirect)	0.29	0.25
$\text{La}_3\text{FeNiS}_7$	0.62 (indirect)	0.30	0.54	Y_3MnCrS_7	0.90 (direct)	1.07	0.39
$\text{La}_3\text{CoCrS}_7$	1.02 (direct)	0.79	0.20	Y_3MnCoS_7	0.49 (direct)	1.16	0.34
				Y_3FeMnS_7	1.05 (indirect)	0.96	1.79
				Y_3CoCrS_7	1.00 (direct)	0.54	0.42



(DOSs). Nine (5 La- and 4 Y-based, as seen in Fig. S1 (ESI[†]) (HTC 4)) promising compounds with E_g values of 1.0–1.8 eV among the 18 stable compounds are singled out. As listed in Table 1, the 9 candidates in bold are reserved for the next screening process, and the rest are removed due to inappropriate E_g . In most cases, the gaps are indirect, but for some systems the gaps are direct, as seen in Fig. S5 (ESI[†]). The latter ones are particularly relevant for thin photo-absorbers because of the higher absorbance. Specifically, for these 9 compounds, we further calculate the E_g value using the HSE06 functional. In the GGA+*U* method, the E_g values of $\text{La}_3\text{CoCrS}_7$ and Y_3CoCrS_7 are 1.02 and 1.00 eV, close to the 1.75 and 1.69 eV obtained by the HSE06 method, as seen in Fig. 4(a) and (b) and Fig. S6a, b (ESI[†]). In addition, for Y_3TiCuS_7 , the E_g values are 1.56 and 1.93 eV from GGA+*U* and HSE06 methods, respectively, as shown in Fig. 5(a) and (b). GGA+*U* calculations show good agreement with those from HSE06 computations, indicating that GGA+*U* can offer trustworthy insights into this family. The three compounds mentioned above meet the screening criteria of optimal E_g . Next, the distributions of the electron structure are taken into account. A comprehensive and rational explanation has been provided for the analysis of the valence electron state density in 3d transition metal elements. This elucidation serves as a sound basis for understanding the electronic structure and behavior of these elements within the context of the mentioned compound, for which the DOSs in $\text{La}_3\text{CoCrS}_7$ are plotted in Fig. 4(a) and (b) from GGA+*U* and HSE06 methods. This shows that electron distributions in the vicinity of the Fermi level are Co 3d/Cr 3d and S 3p states. The Co 3d/Cr 3d states mainly are more localized in lower energy (−4 to −1 eV), which results in the Co 3d/Cr 3d–S 3p reduced hybridization in a higher energy region (−2 to −0.5 eV). This suggests that $\text{La}_3\text{CoCrS}_7$ has innate

ionic and unique covalent bonds, owing to the high covalency of sulphur atoms. Taking the experimentally synthesized $\text{La}_3\text{FeMnS}_7$ as a validated example, the band structures and DOSs are obtained with the E_g values of 1.17 and 1.73 eV from GGA+*U* and HSE06 methods as shown in Fig. S7 (ESI[†]). It exhibits semiconducting behavior with a direct E_g value of about 1.17 eV, *i.e.*, the CBM and VBM positioned at the *K* point. The presented semiconductor-like properties are in good agreement with the experimental resistivity-temperature curves of $\text{La}_3\text{FeMnS}_7$ as an n-type semiconductor.¹⁶ The AFM configuration is found to be magnetic ground state by comparing the energy of the optimized unit lattice with different magnetic configurations, in line with experimental linear chain antiferromagnetism. In brief, 9 (5 La- and 4 Y-based) compounds, as shown in Fig. S1 (ESI[†]) (HTC 4), meet the E_g screening criteria and are passed to the next step in the screening process, namely, the exploration of carrier transport.

3.2.3 Carrier effective masses. Carrier separation creates current before recombination, resulting in energy loss in devices. Higher mobility carriers have lower recombination probability and can be used for the target screening procedure. Although the carrier mobility (μ) is difficult to calculate, the carrier effective mass is related to carrier mobility ($\propto 1/\mu$).⁵² Obviously, the smaller the carrier effective mass, the faster the carrier mobility. Thus, candidates with higher carrier effective mass are ruled out during the screening process. In the current study, we are particularly interested in materials with low effective electron and hole masses, where m_e^* and m_h^* are less than one m_0 as listed in Table 1. Carrier effective mass is obtained by fitting parabola to the VBM and CBM according to band structures calculated along high symmetry points. For instance, by fitting the dispersion relationship between the VBM and the CBM in the band structure diagrams of Y_3TiCuS_7 , the carrier effective mass of electrons and holes can be obtained, in which the m_e^* values are 0.79 m_0 and 0.33 m_0 for m_h^* . Y_3TiCuS_7 in line with our screening criteria properties is now passed to the next step to consider the influence of defects. The carrier masses of $\text{La}_3\text{TiCuS}_7$ are all larger than 1.0 m_0 , in which the m_e^* values are 1.33 m_0 and 1.02 m_0 for m_h^* , so $\text{La}_3\text{TiCuS}_7$ is removed from the filtering path. From the obtained electronic band structures of experimentally synthesized $\text{La}_3\text{FeMnS}_7$, we found that both conduction band edges exhibit tiny amounts of dispersive components, suggesting heavy carrier effective masses. For $\text{La}_3\text{Fe}^{\text{III}}\text{Mn}^{\text{II}}\text{S}_7$, the electron and hole effective masses are 0.7 and 2.5 m_0 , due to the half-full electronic configuration of the high-spin $d^5\text{-Mn}^{2+}$ in octahedral coordination. Since the d–d transition of this configuration is forbidden in quantum mechanics, $\text{La}_3\text{FeMnS}_7$ is ruled out. Finally, 6 (9 minus 3) compounds ($\text{La}_3\text{TiCuS}_7$, $\text{La}_3\text{MnCrS}_7$, La_3VMnS_7 , $\text{La}_3\text{FeMnS}_7$, Y_3FeMnS_7 and Y_3VMnS_7) are expelled from our final list because of large carrier effective masses. As seen in Fig. S1 (ESI[†]) (HTC 5), only $\text{La}_3\text{CoCrS}_7$, Y_3CoCrS_7 and Y_3TiCuS_7 satisfy the following condition with $m_{e,h}^*$ less than one m_0 and are considered as potentially interesting candidates for PV applications to be the focus of further analysis.

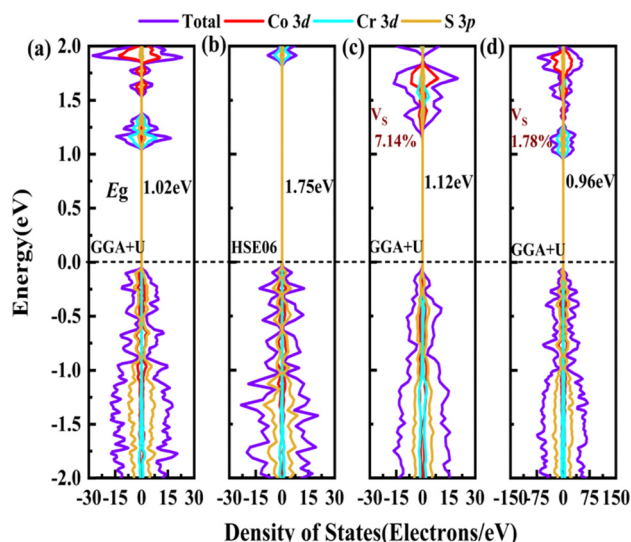


Fig. 4 The electronic structures of $\text{La}_3\text{CoCrS}_7$. The total DOSs for pristine $\text{La}_3\text{CoCrS}_7$ from (a) GGA+*U* and (b) HSE06 methods. The total DOSs for $\text{La}_3\text{CoCrS}_7$ with S vacancy concentrations of (c) 7.14% and (d) 1.78% from the GGA+*U* method. No additional bands are produced. The zero of the energy is set at the Fermi level.



3.2.4 Defect tolerance. The performance of solar cells is highly dependent on the defect characteristics of absorbers. Defects play an important role in semiconductor technologies, such as light-emitting diodes,^{53–55} transistors,^{56–58} and photo-detectors,^{59,60} since defects influence the concentration, mobility and lifetime of electric charge carriers. Finding growth circumstances that favor the desired beneficial defect rather than the undesirable deep level defects is crucial for performance optimization. Actually, defects introduce additional states and restrain the optoelectronic performance. These additional gap states may act as recombination centers and become “traps” for capturing photo-generated carriers (electrons and holes), thus shortening the lifetime of carriers and reducing the charge separation efficiency. Different positions of the additional states have different effects on PV materials. Defect states located deep within the E_g are referred to as defect sensitive, whereas the defect states within the bands or near the band edges are known as defect tolerant. The latter formation of point defects is favorable for PV performance. For instance, the decline in PV energy conversion efficiency with the increasing Ga content in ternary chalcopyrite $\text{Cu}(\text{In,Ga})\text{Se}_2$ alloys is caused by a deep trap.⁶¹ Recombination active defect states at interfaces and in the bulk of $\text{CuIn}_{1-x}\text{Ga}_x\text{Se}_2$ solar cells restrict the open-circuit voltage and fill factor.⁶²

For survived $\text{La}_3\text{CoCrS}_7$, Y_3CoCrS_7 and Y_3TiCuS_7 , we further calculated the effect of point defects on the electronic structure. All possible intrinsic point defects include seven types of vacancies ($V_{\text{La,Y}}$, $V_{\text{Co,Cr,Ti,Cu}}$, V_{S}). The defect tolerances of selected candidates are explored by performing defect calculations on the supercells for seven vacancies in $\text{La}_3\text{CoCrS}_7$, Y_3CoCrS_7 and Y_3TiCuS_7 . The analysis exclusively focuses on neutral defects. It should be stressed that the defect states cannot be expected to be treated accurately at the level of a semi-local exchange–correlation functional as applied here due to the underestimation of the E_g and the delocalization error. The primary goal here is not an in-depth investigation into the precise positions or characteristics of defect levels. Instead, the objective is a qualitative assessment to ascertain whether defect states exist within the E_g . Deep gap states are always caused by an anionic vacancy while the cation vacancy does not seem to introduce deep gap states. Next, we study the presence of a deep lying state in the gap for the S vacancy, showing up as a contribution to the DOSs in the region of the Fermi level. The defect sensitivity for the S vacancy in the three cases is discussed below. Particularly, the vacancy concentration is critical to electron structure properties. In the calculations, S defect concentrations are constructed, namely, 7.14% ($\text{La}_6\text{Co}_2\text{Cr}_2\text{S}_{13}$ (Fig. 4(c)), $\text{Y}_6\text{Co}_2\text{Cr}_2\text{S}_{13}$ (Fig. S6c, ESI†) and $\text{Y}_6\text{Ti}_2\text{Cu}_2\text{S}_{13}$ (Fig. 5(c))), 1.78% ($\text{La}_{24}\text{Co}_8\text{Cr}_8\text{S}_{55}$ (Fig. 4(d)), $\text{Y}_{24}\text{Co}_8\text{Cr}_8\text{S}_{55}$ (Fig. S6d, ESI†) and $\text{Y}_{24}\text{Ti}_8\text{Cu}_8\text{S}_{55}$ (Fig. 5(d))), respectively. It turns out that the introduction of the S vacancy in the cases of $\text{La}_3\text{CoCrS}_7$ and Y_3CoCrS_7 does not give rise to additional states inside the E_g . Defect states close to the band edges and a slight change in E_g are recognized with 1.12 and 1.05 eV for 7.14% (0.96 and 0.85 eV for 1.78%) defect concentrations, respectively. So, $\text{La}_3\text{CoCrS}_7$ and Y_3CoCrS_7 demonstrate

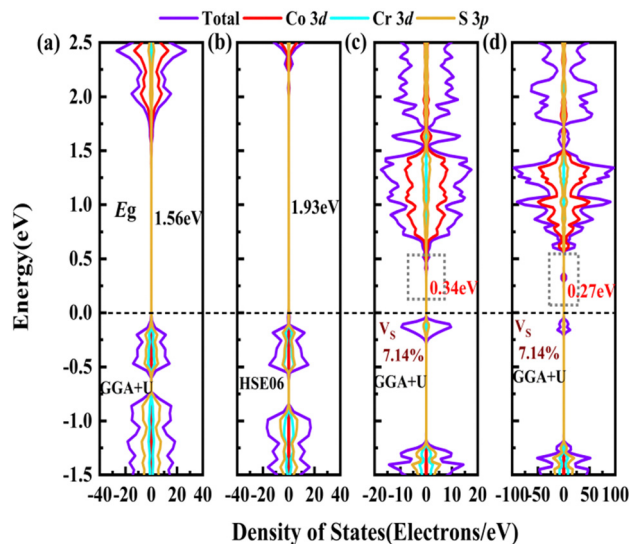


Fig. 5 The electronic structures of Y_3TiCuS_7 . The total DOSs for pristine Y_3TiCuS_7 from (a) GGA+U and (b) HSE06 methods. The total DOSs for Y_3TiCuS_7 with S vacancy concentrations of (c) 7.14% and (d) 1.78% from the GGA+U method. Additional bands are generated to reduce the E_g value. The zero of the energy is set at the Fermi level.

defect tolerance. However, deep states in Y_3TiCuS_7 are found within the E_g , regardless of the defect concentration being 7.14% or 1.78%, so it exhibits defect sensitivity with obvious decreased E_g , about 0.34 and 0.27 eV at S vacancy concentrations of 7.14% and 1.78%. Hence, Y_3TiCuS_7 is discarded from our proposed list of materials due to additional states from the sulfur vacancy. Finally, 2 out of the 3 candidates, $\text{RE}_3\text{CoCrS}_7$ ($\text{RE} = \text{La}, \text{Y}$) are equipped with vacancy tolerance as shown in Fig. S1 (ESI†) (HTC 6), and screened out as potential materials for solar energy conversion applications.

3.2.5 Optical absorption coefficient. Compared with the recently developed lead halide perovskites, sulfide devices have the potential to utilize even thinner absorbers in order to achieve high PCE.^{11,63} Employing absorbers with exceedingly high optical absorption offers significant advantages to solar cell performance, since photo-generated electrons and holes travel short distances before being collected. By doing so, the linked open circuit voltages will increase and significantly reduce non-radiative recombination. To determine the optical absorption coefficient in the visible range of A_3CoCrS_7 ($\text{A} = \text{La}, \text{Y}$), we calculate the optical absorption coefficient as a function of wavelength, following methodology described in the theoretical methods section. The calculated spectra of survived A_3CoCrS_7 ($\text{A} = \text{La}, \text{Y}$) with the PBE method are shown in Fig. 6(a) and (b). The spectra exhibit roughly similar behavior represented by high-intensity absorption features appearing in the visible range with broad edges extending to 450 and 580 nm, which the theoretical optical absorption coefficient can reach 10^5 cm^{-1} in the visible-light region. This value is much higher than those for perovskite oxides ($\sim 10^3\text{--}10^4 \text{ cm}^{-1}$), and even widely used silicon ($\sim 10^4 \text{ cm}^{-1}$). The [001] direction exhibits a notably higher optical density between 430 and



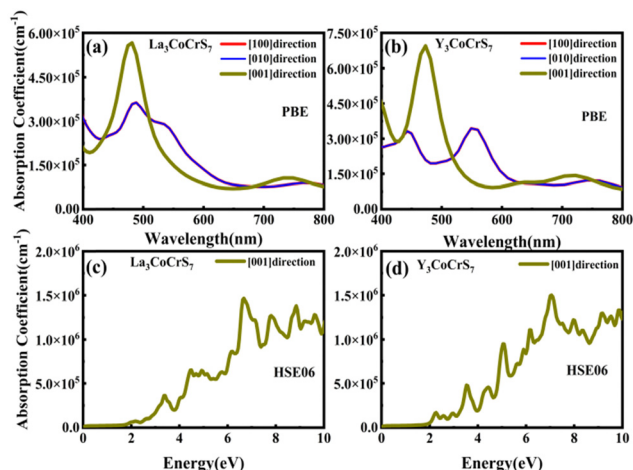


Fig. 6 The optical absorption spectra in the visible range of (a) and (c) $\text{La}_3\text{CoCrS}_7$ and (b) and (d) Y_3CoCrS_7 with the PBE and HSE06 methods.

500 nm compared to the [100] or [010] directions, emphasizing the importance of anisotropy and making it easier for electron-hole transfer to travel in one direction than the others, which is suitable for high-absorptivity solar energy absorbing coating. The absorption edge is an important feature of the semiconductor absorption spectrum, which corresponds to the lowest energy value of the electron transition from the VBM to the CBM. The calculated optical absorption coefficient as a function of energy in incident light is shown in Fig. 6(c) and (d) with the HSE06 method. The absorption edges of A_3CoCrS_7 ($\text{A} = \text{La}, \text{Y}$) locate at the positions of 1.75 and 1.69 eV, which is exactly the semiconductor E_g value calculated from the electronic structure.

4. Conclusions

In accordance with the prerequisite for excellent solar cell absorber materials, a systematic screening procedure, considering stability (energy above E_{hull} , dynamic and thermodynamic), E_g values in the range of 1.0–1.8 eV, carrier effective masses of less than $1.0 m_0$, reasonable defect sensitivity, and high optical absorption coefficient, is conducted for the $\text{RE}_3\text{BB}'\text{S}_7$ ($\text{RE} = \text{La}, \text{Y}$; $\text{B/B}' = \text{Ti-Cu}$ in the 3d row) family using the first principle calculations based on the GGA+ U and HSE06 approximations. The initial structural models established by comprehensive data mining deliver seven possible polymorphs, on which collinear magnetic configurations are applied to examine the ground state. Finally, high-throughput calculations confined by the screening criteria for solar cell materials leave the synthesizable $\text{RE}_3\text{CoCrS}_7$ (hexagonal, $P6_3$ (no. 173), $\text{RE} = \text{La}, \text{Y}$) as potential light absorbers for solar energy conversion. Our work provides a reasonable prediction of magnetic, electrical and optical properties to guide the experimental synthesis of the $\text{A}_3\text{BB}'\text{S}_7$ family, and sheds light on the essential conditions required for stable sulfide solar cell conversion materials to offer valuable insights into their potential advantages in sulfide photovoltaic devices. Furthermore, the

methodology applied in this work is expected to accelerate the discovery of novel light absorption layer materials in potential sulfide systems via the proposed high-throughput computing and screening approach.

Author contributions

Jinjin Yang collected publications and completed the framework of the manuscript. Zhongxiong Sun, Dao-Xin Yao, and Man-Rong Li interpreted the results, revised the manuscript, and approved the final version.

Conflicts of interest

The authors declare no competing financial interest.

Acknowledgements

This work was financially supported by the National Science Foundation of China (grant no. 22090041), the Program for Guangdong Introducing Innovative and Entrepreneurial Teams (grant no. 2017ZT07C069), and the Guangdong Basic and Applied Basic Research Foundation (grant no. 2022B1515120014). The authors thank the National Supercomputer Center in Guangzhou.

Notes and references

- 1 L. Liu, Y. Xie, J. S. Tse and Y. M. Ma, *Mater. Adv.*, 2023, **4**, 570–577.
- 2 D. Cardenas-Morcoso, M. García-Tecedor, T. Merdzhanova, V. Smirnov, F. Finger, B. Kaiser, W. Jaegermann and S. Giménez, *Mater. Adv.*, 2020, **1**, 1202–1211.
- 3 K. Timmo, K. Muska, M. Pilvet, M. Altosaar, V. Mikli, M. Danilson, R. Kaupmees, J. Krustok, M. Grossberg-Kuusk and M. Kauk-Kuusik, *Mater. Adv.*, 2023, **4**, 4509–4519.
- 4 K. Zhang and L. Guo, *Catal. Sci. Technol.*, 2013, **3**, 1672–1690.
- 5 H. I. Eya, E. Ntsoenzok and N. Y. Dzade, *Mater.*, 2020, **13**, 978.
- 6 N. Meftahi, M. Klymenko, A. J. Christofferson, U. Bach, D. A. Winkler and S. P. Russo, *npj Comput. Mater.*, 2020, **6**, 166.
- 7 H. Wang, G. Gou and J. Li, *Nano Energy*, 2016, **22**, 507–513.
- 8 S. Niu, D. Sarkar, K. W. Williams, Y. Zhou, Y. Li, E. Bianco, H. Huyan, S. B. Cronin, M. E. McConney, R. Haiges, R. Jaramillo, D. J. Singh, W. A. Tisdale, R. Kapadia and J. Ravichandran, *Chem. Mater.*, 2018, **30**, 4882–4886.
- 9 J. Du and J. Shi, *Adv. Mater.*, 2019, **31**, 1905643.
- 10 N. Bozdog, H. Koc, S. Simsek, A. M. Mamedov and E. Ozbay, *Ferroelectrics*, 2019, **544**, 1–10.
- 11 O. Shahid, S. Yadav, D. Maity, M. Deepa, M. K. Niranjana and J. Prakash, *New J. Chem.*, 2023, **47**, 5378–5389.
- 12 F. Zhao, Y. Liang, J. B. Lee and S. J. Hwang, *Mater. Sci. Eng., B*, 2019, **248**, 114404.



- 13 L. Sun, H. Dong, P. Zhang and C. Yan, *Annu. Rev. Phys. Chem.*, 2015, **66**, 619–642.
- 14 M. Pokhrel, G. A. Kumar and D. K. Sardar, *MRS Online Proc. Libr.*, 2012, **1471**, 56–61.
- 15 G. A. Kumar, M. Pokhrel and D. K. Sardar, *Mater. Lett.*, 2012, **68**, 395–398.
- 16 K. S. Nanjundaswamy and J. Gopalakrishnan, *J. Solid State Chem.*, 1983, **49**, 51–58.
- 17 M. Daszkiewicz, L. D. Gulay, A. Pietraszko and V. Y. Shemet, *J. Solid State Chem.*, 2007, **180**, 2053–2060.
- 18 L. D. Gulay, O. S. Lychmanyuk, I. D. Oleksyuk, M. Daszkiewicz, J. Stępień-Damm and A. Pietraszko, *J. Alloys Compd.*, 2007, **431**, 185–190.
- 19 L. D. Gulay, O. S. Lychmanyuk, I. D. Oleksyuk and A. Pietraszko, *J. Alloys Compd.*, 2006, **422**, 203–207.
- 20 O. M. Strok, M. Daszkiewicz, L. D. Gulay and D. Kaczorowski, *J. Alloys Compd.*, 2010, **493**, 47–49.
- 21 W. Yin, W. Wang, L. Kang, Z. Lin, K. Feng, Y. Shi, W. Hao, J. Yao and Y. Wu, *J. Solid State Chem.*, 2013, **202**, 269–275.
- 22 T. Murugesan, S. Ramesh, J. Gopalakrishnan and C. N. R. Rao, *J. Solid State Chem.*, 1981, **38**, 165–172.
- 23 J. He, Z. Wang, X. Zhang, Y. Cheng, Y. Gong, X. Lai, C. Zheng, J. Lin and F. Huang, *RSC Adv.*, 2015, **5**, 52629–52635.
- 24 G. Kresse and J. Hafner, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 1995, **47**, 558–561.
- 25 G. Kresse and J. Hafner, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 1994, **49**, 14251–14269.
- 26 P. E. Blöchl, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 1994, **50**, 17953–17979.
- 27 M. K. Horton, J. H. Montoya, M. Liu and K. A. Persson, *npj Comput. Mater.*, 2019, **5**, 64.
- 28 A. Jain, G. Hautier, S. P. Ong, C. J. Moore, C. C. Fischer, K. A. Persson and G. Ceder, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 2011, **84**, 045115.
- 29 A. Togo and I. Tanaka, *Scr. Mater.*, 2015, **108**, 1–5.
- 30 L. D. Gulay, I. D. Oleksyuk, M. Wolczyr and J. Stępień-Damm, *Z. Anorg. Allg. Chem.*, 2005, **631**, 1919–1923.
- 31 X. Q. Liu, J. Wu, X. Shi, H. J. Zhao, H. Y. Zhou, R. Qiu, W. Zhang and X. M. Chen, *Appl. Phys. Lett.*, 2015, **106**, 202903.
- 32 B. W. Rudyk and S. S. Stoyko, *J. Solid State Chem.*, 2013, **208**, 78–85.
- 33 L. D. Gulay, V. Y. Shemet and I. D. Oleksyuk, *J. Alloys Compd.*, 2005, **394**, 250–254.
- 34 A. C. Ludvigsen, Z. Lan and I. E. Castelli, *Materials*, 2022, **15**, 309.
- 35 S. Jaulmes and M. Julien-Pouzol, *Acta Crystallogr.*, 1977, **33**, 3898–3899.
- 36 M. Saeki, Y. Yajima and M. Onoda, *J. Solid State Chem.*, 1991, **92**, 286–294.
- 37 B. Chen, B. W. Eichhorn and W. Wong-Ng, *Acta Crystallogr.*, 1994, **50**, 161–164.
- 38 Y. Han, J. Xu, Y. Liang, X. Chen, M. Jia, J. Zhang and Z. Shi, *Chem. Eng. J.*, 2023, **473**, 145351.
- 39 M. Li, E. E. McCabe, P. W. Stephens, M. Croft, L. Collins, S. V. Kalinin, Z. Deng, M. Retuerto, A. Sen Gupta, H. Padmanabhan, V. Gopalan, C. P. Grams, J. Hemberger, F. Orlandi, P. Manuel, W. Li, C. Jin, D. Walker and M. Greenblatt, *Nat. Commun.*, 2017, **8**, 2037.
- 40 G. Cai, M. Greenblatt and M. Li, *Chem. Mater.*, 2017, **29**, 5447–5457.
- 41 M. Li, M. Croft, P. W. Stephens, M. Ye, D. Vanderbilt, M. Retuerto, Z. Deng, C. P. Grams, J. Hemberger, J. Hadermann, W. Li, C. Jin, F. O. Saouma, J. I. Jang, H. Akamatsu, V. Gopalan, D. Walker and M. Greenblatt, *Adv. Mater.*, 2015, **27**, 2177–2181.
- 42 Y. Han, M. Wu, C. Gui, C. Zhu, Z. Sun, M. Zhao, A. A. Savina, A. M. Abakumov, B. Wang, F. Huang, L. He, J. Chen, Q. Huang, M. Croft, S. N. Ehrlich, S. A. Khalid, Z. Deng, C. Jin, C. P. Grams, J. Hemberger, X. Wang, J. Hong, U. Adem, M. Ye, S. Dong and M. Li, *npj Quantum Mater.*, 2020, **5**, 92.
- 43 S. Zhao, J. Yang, Y. Han, M. Wu, M. Croft, P. W. Stephens, D. Walker, M. Greenblatt and M. Li, *Chem. Mater.*, 2022, **34**, 1930–1936.
- 44 L. Gao, X. Wu, J. Xu, X. Tian, B. Zhang and K. Wu, *J. Alloys Compd.*, 2021, **900**, 16353.
- 45 K. Kuhar, A. Crovetto, M. Pandey, K. S. Thygesen, B. Seger, P. C. Vesborg, O. Hansen, I. Chorkendorff and K. W. Jacobsen, *Energy Environ. Sci.*, 2017, **10**, 2579–2593.
- 46 W. Sun, S. T. Dacek, S. P. Ong, G. Hautier, A. Jain, W. D. Richards, A. C. Gamst, K. A. Persson and G. Ceder, *Sci. Adv.*, 2016, **2**, e1600225.
- 47 G. J. Martyna, M. L. Klein and M. E. Tuckerman, *J. Chem. Phys.*, 1992, **97**, 2635–2643.
- 48 J. Guillemoles, T. Kirchartz, D. Cahen and U. Rau, *Nat. Photonics*, 2019, **13**, 501–505.
- 49 J. Feng, X. Wang, J. Li, H. Liang, W. Wen, E. Alvianto, C. Qiu, R. Su and Y. Hou, *Nat. Commun.*, 2023, **14**, 5392.
- 50 A. Srivastava, P. Sarkar, S. K. Tripathy, T. R. Lenka, P. S. Menon, F. Lin and A. G. Aberle, *Sol. Energy*, 2020, **209**, 206–213.
- 51 M. S. Kumar, S. P. Madhusudanan and S. K. Batabyal, *Sol. Energy Mater. Sol. Cells*, 2018, **185**, 287–299.
- 52 S. Haldar, V. K. Dixit, G. Vashisht, S. K. Khamari, S. Porwal, T. K. Sharma and S. M. Oak, *Sci. Rep.*, 2017, **7**, 4905.
- 53 Y. Kim, S. Kim, A. Kakekhani, J. Park, J. Park, Y. Lee, H. Xu, S. Nagane, R. B. Wexler, D. Kim, S. Jo, L. Martínez-Sarti, P. Tan, A. Sadhanala, G. S. Park, Y. Kim, B. Hu, H. J. Bolink, S. Yoo, R. H. Friend, A. M. Rappe and T. Lee, *Nat. Photonics*, 2021, **15**, 148–155.
- 54 T. Hu, D. Li, Q. Shan, Y. Dong, H. Xiang, W. C. Choy and H. Zeng, *ACS Mater. Lett.*, 2021, **3**, 1702–1728.
- 55 J. M. Azpiroz, E. Mosconi, J. Bisquert and F. De Angelis, *Energy Environ. Sci.*, 2015, **8**, 2118–2127.
- 56 Y. Liu, H. Liu, Y. Chu, Y. Cui, T. Hayasaka, V. Dasaka, L. Nguyen and L. Lin, *Adv. Mater. Interfaces*, 2018, **5**, 1701640.
- 57 L. L. Gao, X. Zhang, H. Yu, M. Hong, X. Wei, Z. Chen, Q. Zhang, Q. Liao, Z. Zhang and Y. Zhang, *ACS Appl. Mater. Interfaces*, 2023, **15**, 38603–38611.



- 58 Y. Song, H. Zhou, X. Cai, Y. Liu, P. Yang, G. Zhang, Y. Zhang, M. Lan and S. Wei, *ACS Appl. Mater. Interfaces*, 2020, **12**, 29993–29998.
- 59 J. Y. Li, G. Zhang, Z. Zhang, J. Li, Z. Uddin, Y. Zheng, Y. Shao, Y. Yuan and B. Yang, *ACS Appl. Mater. Interfaces*, 2021, **13**, 56358–56365.
- 60 D. Wu, J. Guo, C. Wang, X. Ren, Y. Chen, P. Lin, L. Zeng, Z. Shi, X. J. Li, C. Shan and J. Jie, *ACS Nano*, 2021, **15**, 10119–10129.
- 61 L. L. Kerr, S. S. Li, S. W. Johnston, T. J. Anderson, O. D. Crisalle, W. K. Kim, J. Abushama and R. N. Noufi, *Solid State Electrochem.*, 2004, **48**, 1579–1586.
- 62 J. Chen, X. Chang, J. Guo, Q. Gao, X. Zhang, C. Liu, X. Yang, X. Zhou, B. Chen, F. Li, J. Wang, X. Yan, D. Song, H. Li, B. S. Flavel, S. Wang and J. Chen, *Research*, 2023, **6**, 0084.
- 63 D. O. Scanlon and A. Walsh, *Acta Crystallogr.*, 2015, **71**, 702–706.

