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Unravelling the UV luminescence of Bi-doped LiYGeO₄: a journey from first principles to temperature-dependent photoluminescence

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Materials exhibiting persistent luminescence (PersL) have garnered attention due to their unique ability to emit light for extended periods after the excitation stops. LiYGeO₄ has arisen as a promising host for PersL due to its notable defect abundance. When doped with Bi, it can provide a long-lasting ultraviolet emission, which may find interesting applications in areas such as photodynamic therapy or self-sustained photocatalysis. In the present paper, undoped and Bi-doped LiYGeO₄ (0.5 mol%) samples were synthesized by solid-state reaction. X-ray diffraction confirmed the presence of the LiYGeO₄ phase, alongside residual yttrium germanate phases. Additionally, we calculated its previously undocumented band structure using hybrid density functional theory to provide new theoretical insights. These calculations indicate that LiYGeO₄ has a direct bandgap at the Γ point, further supported by the absorption data. Elemental analyses allowed quantification of the samples, identifying the loss of Li during the synthesis. Room temperature (RT) photoluminescence (PL) showed a strong emission band peaked around 350−360 nm. Despite the extensive research conducted on LiYGeO₄:Bi, the mechanisms underlying the PersL phenomenon remain unclear. To address this, we conducted temperature-dependent PL from 17 K to RT, using a 325 nm photon excitation. We propose that, at RT, the observed emission arises from the overlap of the $^3P_1^{(1)} \rightarrow {}^1S_0$ and $^3P_1^{(2)} \rightarrow {}^1S_0$ Bi³⁺ intraionic transitions, sublevels of the $^3P_1$ state, which is completely split under the local crystal field generated by the ion environment. However, at cryogenic temperatures, the $^3P_0 \rightarrow {}^1S_0$ transition is the dominant one. Moreover, PersL emission was achieved for nearly 7 h, with 250 nm photon excitation for 10 min, arising from a contribution of the overlapped $^3P_1^{(1)} \rightarrow {}^1S_0$ and $^3P_1^{(2)} \rightarrow {}^1S_0$ transitions.

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1. Introduction

In recent years, materials exhibiting the optical phenomenon known as persistent luminescence (PersL) have gained significant attention due to their remarkable properties. The main driving force for such interest is the versatility that the unique ability of these materials to emit light for extended periods without continuous excitation can bring to PersL-based innovative applications. Indeed, PersL phosphors have been explored in multiple applications across different stages of maturity, some more well-established, such as emergency signage, safety indications, luminous paint, or watch dials, others still in the research and development phase, *e.g.* anti-counterfeiting, photocatalysis, data storage, security, sensing, and bio-imaging, among others.[1–7]

Among the numerous PersL emitters explored, considerable research has been undertaken regarding the visible and near-infrared (NIR) spectral regions, however, there is a notable gap in the reported ultraviolet (UV) PersL emitters, despite their

potential in several applications such as photodynamic therapy, 3D-printing technology, disinfection, and persistent photocatalysis.[4,8,9] So far, in the UVA range (315–400 nm), lithium yttrium germanate (LiYGeO$_4$) appears as a promising host candidate, mainly due to its abundance of intrinsic defects, such as vacancies, antisites, and interstitials that can act as trap defects to mediate PersL emission.[10,11] LiYGeO$_4$ has gained significant attention since it is able to provide PersL in different spectral regions depending on the dopant introduced as the emitting centre.[11–17] This matrix is particularly important to provide PersL in the UVA spectral region, especially when doped with trivalent bismuth (Bi^{3+}). Indeed, to the best of the authors' knowledge, LiYGeO$_4$ is the matrix that led to the longest PersL time in this spectral region (nearly 300 h, including photostimulation).[18] Moreover, the selection of this specific host is highlighted by the fact that changing its elements (*e.g.* Li and Ge by Na and Si, respectively) leads to detrimental results either by an undesirable redshift of the emission or by creating traps that are too deep to participate in the PersL mechanism.[10,19] Nevertheless, there are still a great number of open questions regarding the fundamental properties of this material. For instance, Qiao *et al.*[10] and Zhou *et al.*[11] estimated its bandgap through hybrid density functional theory (DFT), having obtained values of 6.84 eV and 6.75 eV, respectively. However, these values are considerably different from the experimental ones. Particularly, Shao *et al.*[16] and Shi *et al.*[18] reported values of 4.51 eV and 4.6 eV, respectively, for the bandgap, obtained through diffuse reflectance and absorption analysis. Such differences are likely related to the methods used to analyse the experimental data, which often do not consider the influence of the defect states that are known to be present near the band edges (the so-called tail states) and that have a significant effect on the bandgap energy determination.

In the particular case of UVA PersL, ions with the s^2 ground state electronic configuration, as is the case with the Bi^{3+}, have been selected as promising emitting centres. This ion has been studied as a luminescent activator for several decades owing to its versatility of emitting in a wide spectral range, from the UV to the IR, depending on the host in which it is incorporated.[20–23] Such characteristics derive from the strong influence that the surrounding environment, *i.e.*, the local crystalline field, has on its energy levels.[24–26] Bismuth is a nontoxic heavy metal that presents valence states varying from −3 to +5. While bismuth anions (Bi^{3-}, Bi^{2-}, Bi^{-}) are unsuitable for luminescence applications since they do not show emission, the cations (Bi^{2+}, Bi^{3+}, Bi^{5+}) are viable for such purposes.[24] Among those, Bi^{3+} is the most stable when incorporated in inorganic hosts,[25] resulting in PersL emission peaking at ∼350 nm when introduced in LiYGeO$_4$, which can last for several hours after stopping the initial photon excitation, typically at 250 nm.[4,10,16,18,27] So far, and based on the findings reported by Shi *et al.*,[18] the optimal doping concentration for achieving the most effective PersL in this material appears to be 0.5 mol% Bi. In fact, Shi *et al.*[18] and Shao *et al.*[16] reported PersL durations of ∼72 h (without any additional stimulation) using this doping concentration. In all the reported cases, the intraionic $^3P_1 \rightarrow {}^1S_0$ transition of Bi^{3+} has been pointed out as the one dominating the emission at room temperature (RT) and responsible for PersL.[4,10,18]

LiYGeO$_4$, an olivine-type material, crystallizes in an orthorhombic phase within the D_{2h}^{16} space group.[16,28,29] Its structure consists of a corner- and edge-sharing of [GeO$_4$] tetrahedra, [LiO$_6$] octahedra, and [YO$_6$] octahedra, connected by oxygen atoms (see Fig. 1). Its cations (Ge^{4+}, Li$^+$, and Y^{3+}) possess, respectively, ionic radii and coordination numbers of $r_{Ge^{4+}}$ = 0.53 Å, CN = 4; r_{Li^+} = 0.76 Å, CN = 6 and r_{Y^+} = 0.9 Å, CN = 6.[16,18] The Li-ion occupies the 4a site (Wyckoff position), characterized by a C_i symmetry with an inversion centre (centrosymmetric), while the Y-ion is located at a 4c site, possessing a C_s symmetry, without an inversion centre.[30] According to Qiao *et al.*,[10] when Bi^{3+} ($r_{Bi^{3+}}$ = 1.03 Å [4,10,16,31]) is introduced into the LiYGeO$_4$ matrix, it is mostly located at the Y^{3+} site due to the similar radii, with the incorporation into the Ge^{4+} or Li$^+$ sites being negligible, as also expected taking into account the charge mismatch. Additionally, in the same work,[10] the authors performed a detailed DFT study regarding possible intrinsic point defects, including vacancies (V$_{Li}$, V$_Y$, V$_{Ge}$, V$_{O1,O2,O3}$), antisites (Li$_Y$, Li$_{Ge}$, Y$_{Li}$, Y$_{Ge}$, Ge$_{Li}$, Ge$_{RE}$), and interstitials (Li$_i$, Y$_i$, and Ge$_i$). The interstitial defects Y$_i$ and Ge$_i$ were discarded from the discussion since no stable geometries were found for them. Regarding the remaining point defects, the most stable spin configuration was obtained. The formation energies of the defects were evaluated considering different positions of the Fermi level (E_F) within the bandgap, but since the materials were synthesized in an oxidizing atmosphere, which corresponds to an E_F slightly lower than the middle of the bandgap, the value of 0.5 eV below the middle position was the main focus. They concluded that, in the case of the LiYGeO$_4$, the most stable point defects (and their respective formation energies) are Y$_{Li}^{2+}$ (−0.9 eV) > Ge$_Y^+$ (−0.57 eV) ≈ Li$_Y^i$ (−0.57 eV) > Ge$_{Li}^{3+}$ (0.32 eV) > V$_{Li}^-$ (0.66 eV).[10]
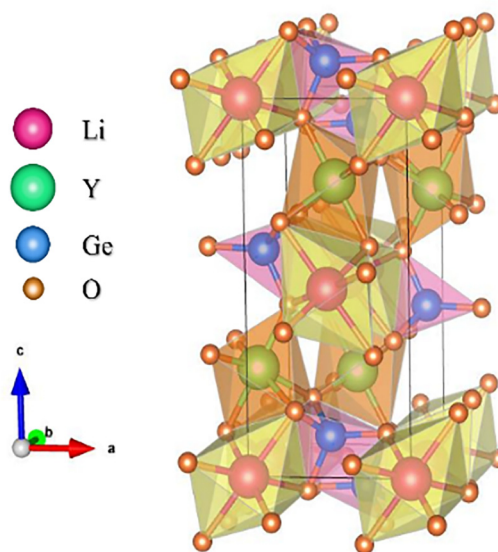

Fig. 1 Schematic representation of the LiYGeO$_4$ unit cell. The model was created using the free program VESTA and the data were adapted from the CIF file of mp-504948 using the information from the 04-002-3479 card.

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In its trivalent charge state, Bi^{3+} presents the electronic configuration [Xe] $4f^{14}$ $5d^{10}$ $6s^2$, with the ground level corresponding to $^1S_0$. When excited, the s^1p^1 electron configuration leads to the formation of an upper ^1P spin-singlet state and a lower ^3P spin-triplet state, with the latter further split by the spin–orbit interaction, resulting in $^3P_0$, $^3P_1$, and $^3P_2$ states in increasing energy. The non-degenerate ^1P level results in the $^1P_1$ state.[4,18,20,21] Upon UV photon excitation, the $^1S_0 \rightarrow {}^3P_0$ transition is strictly forbidden due to spin and parity selection rules, which frequently makes it too weak to be observed, except at cryogenic temperatures ($T < 80$ K).[4,25] However, the $^1S_0 \rightarrow {}^1P_1$ transition is spin-allowed, while the $^1S_0 \rightarrow {}^3P_1$ and $^1S_0 \rightarrow {}^3P_2$ transitions become partially allowed due to spin–orbit coupling and admixture of the wave functions by phonon coupling.[4]

Although already evaluated in other hosts such as Y_2O_3 or La_2O_3 for more than three decades,[24,32–34] to date, no temperature-dependent luminescence studies have been reported on Bi-doped $LiYGeO_4$, which is considered a fairly new material that is gaining momentum lately, essentially due to its PersL properties. Hence, in this study, nominally undoped $LiYGeO_4$ (LYG) and 0.5 mol% Bi-doped $LiYGeO_4$ (LYGBi) samples were produced by solid-state reaction and the thermal, structural, elemental, and optical properties were investigated by thermogravimetry (TGA), X-ray diffraction (XRD), nuclear reaction analysis (NRA), particle-induced X-ray emission (PIXE), Rutherford backscattering spectroscopy (RBS), elastic backscattering spectrometry (EBS), diffuse reflectance (DR), RT steady-state photoluminescence (PL) and PL excitation (PLE), temperature-dependent PL (from 17 K to RT), thermoluminescence (TL) as well as afterglow measurements. Valuable insights were obtained with the temperature-dependent PL, allowing us to identify that the main transition observed at RT arises from the mixture of the transitions from the $^3P_1$ sublevels, which are completely split by the crystal field, to the $^1S_0$ ground state, while at cryogenic temperatures the $^3P_0 \rightarrow {}^1S_0$ transition is the dominant one. These results highlight the importance of a detailed temperature-dependent PL study on the correct assignment of the observed optical transitions, which can have important implications in material optimization and its further applications. Regarding PersL, it was seen that this phenomenon could be attained when the samples were excited with 250 nm (4.96 eV) photons for 10 min, while longer wavelengths were seen to be ineffective for PersL activation. The observed emission can be ascribed to a combination of the $^3P_1^{(1)} \rightarrow {}^1S_0$ and $^3P_1^{(2)} \rightarrow {}^1S_0$ transitions. Additionally, hybrid DFT calculations were performed, providing the band structure determination with bandgap energy that closely matches the value obtained experimentally *via* diffuse reflectance measurements.

2. Experimental details

2.1 Materials preparation

In the present paper, nominally undoped and 0.5 mol% Bi-doped $LiYGeO_4$ samples were prepared by a high-temperature solid-state reaction. The quantity of Bi doping was selected based on previous studies performed by the authors in which the nominal Bi amount varied (0.1, 0.5, 1 and 5 mol%, not shown), as well as on formerly reported research,[18] which concluded that the maximum PersL emission was achieved at 0.5 mol% of Bi.[18] The raw materials Li_2CO_3 (99%, Panreac), Y_2O_3 (99.99%, Thermo Scientific Chemicals), GeO_2 (99.998%, Thermo Scientific Chemicals) and Bi_2O_3 (99.8%, Alfa Aesar) were weighed according to the stoichiometric ratio to obtain $LiYGeO_4$ and $LiYGeO_4$:Bi (0.5 mol%) samples. The precursors were mixed in an agate ball mill with ethanol for 4 h. Then, the mixtures were left to dry on a hot plate at 90 °C overnight to evaporate all the ethanol. Afterwards, the mixtures were transferred to alumina crucibles and fired at 800 °C for 2 h in an oven (heating rate of 5 °C min^{-1}). After this step, the powders were ground and homogenized in an agate mortar and pressed into pellets with a 13 mm diameter (5-ton uniaxial press, 2 min). These pellets were subjected to a thermal annealing treatment in air at 1000 °C for 40 h in an oven (heating rate of 5 °C min^{-1}). Multiple samples were prepared for each batch, and these batches were repeated in the same conditions to assess the homogeneity and reproducibility of the materials. The analysed properties were found to be consistent, and the data presented in this paper are representative of all the samples.

2.2 Materials characterization

Thermal characterization of the materials was performed in a Hitachi Nexta STA300, placing 15.692 mg in an alumina crucible, under a mix of O_2 and N_2 (air composition) and using a ramp temperature program mode between 24–1180 °C, with a rate of 10 °C min^{-1}.

The crystalline structure of the samples was assessed by XRD, conducted under a PANalytical X'Pert PRO diffractometer operating in the Bragg–Brentano configuration and with CuK$_\alpha$ radiation ($\lambda = 1.54056$ Å). DR was performed in a UVVIS JASCO V-780, with an integrating sphere JASCO ISN-901i, in the wavelength range 190–1600 nm with a bandwidth of 5 nm and a speed of 100 nm min^{-1}.

The elemental and stoichiometric composition of the samples was characterized through ion beam analysis techniques. All the measurements were conducted with the help of the 2.5 MV Van de Graaff accelerator installed at the Laboratory of Accelerators and Radiation Technologies (LATR) of Instituto Superior Técnico (IST) using MeV H$^+$ or He$^+$ beams.[35] Distinct events are simultaneously observed from the particle spectra. The contents of light elements such as O or Li are commonly evaluated through the corresponding elastic backscattering yields *via* EBS. The contents of heavier elements such as Y, Ge or Bi are followed through Rutherford backscattering yields *via* RBS. Simultaneously, the depth profiling and contents of Li may be obtained *via* NRA by following the yield of the $^7Li(p,\alpha)^4He$ nuclear reaction.[36,37] This reaction has a high Q value (17.35 MeV), avoiding the superposition of the NRA yield with those induced by the EBS and RBS events.[38] The collection of particle spectra was performed using a dedicated chamber

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equipped with two unsealed Si PIN diode detectors (Hamamatsu; depletion layer thickness of 300 μm) with an energy resolution of 14 keV. For NRA, the detector was positioned at a scattering angle of −165°, and the nuclear reaction ^{7}Li(p,α)^{4}He was used to determine the Li concentration along the sample's depth, while for EBS and RBS, the detector was placed at +165°; in this case, a higher signal amplification of the detection system was used to enhance the individual analysis of the EBS and RBS yields. The particle spectra were acquired simultaneously, with a total integrated beam charge of 5 μC using a proton beam collimated to a dimension of 0.5 × 0.5 mm^2, with an energy of 2.1 MeV, at which the cross-section of the ^{7}Li(p,α)^{4}He nuclear reaction reaches a maximum of approximately 5 mb sr^{-1}.[36,37] Using the Oxford Microbeams nuclear microprobe beamline for focusing a 2 MeV proton beam down to 5 × 5 μm^2 and scanning a sample area up to 2640 × 2640 μm^2, PIXE spectra were collected with a 30 mm^2 SDD detector. The deconvolution and quantitative analysis with the particle and PIXE spectra were carried out using the NDF[39,40] and GUPIXWIN[41] software, respectively.

Steady-state PL and PLE spectroscopy measurements were conducted at RT using Fluorolog-3 Horiba Scientific setup, equipped with a 450 W Xe lamp coupled to a monochromator as the excitation source, a double additive grating Gemini 180 monochromator (1200 grooves mm^{-1} and 2 × 180 mm) in the excitation side and a triple grating iHR550 spectrometer in the emission (1200 grooves mm^{-1} and 550 mm). The intensity of the PersL with time (afterglow) was carried out in the same equipment, using a kinetics mode. The samples were excited with the same Xe lamp for 10 minutes with 250 nm photons, after which the excitation was stopped and the decay curves were acquired. Additionally, temperature-dependent PL measurements were recorded on a selected sample using as excitation the 325 nm (∼3.81 eV) line of a continuous helium–cadmium (He–Cd) laser (I_0 < 0.6 W cm^{-2}). In this case, the samples were placed on a cold finger He cryostat and the luminescence was dispersed by a SPEX 1704 monochromator (1 m, 1200 grooves mm^{-1}), being detected with a water-cooled Hamamatsu RCA C31034 photomultiplier. The temperature was varied from 17 K up to RT. These experiments ensured optical alignment at a 90-degree geometry between sample irradiation and signal detection. The presented spectra were corrected to the spectral response of the measurement systems. In addition, the Jacobian transformation was considered for the quantitative analysis of the temperature-dependent PL data.

The thermoluminescence (TL) measurements were performed using a Risø thermoluminescence/optically-stimulated luminescence reader (TL/OSL-DA-20), manufactured by DTU Physics, at the Luminescence Dating Laboratory of IST. A ^{90}Sr/^{90}Y β-source (with a nominal activity of 825 MBq and a dose rate in quartz of 0.062 ± 0.001 Gy s^{-1}) was used to excite the samples for 200 s. The T_M–T_{STOP} method was employed to identify the number of traps present in the samples, by monitoring the temperature position of the first peak (T_M) in the glow curve as a function of the pre-heating temperature T_{STOP}, which was varied between 323 and 523 K.[42] After each irradiation and pre-heating, a full TL scan was acquired from RT up to 673 K at a constant heating rate of 2 K s^{-1}.

2.3 Modelling and computational details

The spin-polarized first-principles calculations were performed using the VASP code[43] at the HSE06 level of theory,[44] with the fraction of exact exchange in the hybrid functional set to 0.30 to allow the calculated bandgap (6.01 eV) to match the experimental one. The bandgaps calculated with different fractions of exact exchange (α) in the hybrid functional, which were tested in preliminary calculations, are 5.59 eV (α = 0.25) and 6.42 eV (α = 0.35). The range-separated hybrid functional HSE06 was used for all the calculations to maximise the agreement with the experimental results. This functional is known to produce essentially the same results as PBE0,[45] as we verified by calculating the band diagram of LiYGeO$_4$ and finding a band structure that qualitatively resembles the ones given in the literature,[10,11] which capture the most important band features but show an underestimated bandgap. This way, we obtain a more reliable band structure of this material,[46,47] which confirms the standard-DFT results in the literature, and was essential for determining the positions of defect-related levels in the bandgap of LiYGeO$_4$.

The simulations were run on a 1 × 2 × 2 LiYGeO$_4$ supercell, with periodic boundary conditions in all Cartesian directions. The explicitly treated electrons were 2s^1 for Li, 4s^2 4p^6 5s^2 4d^1 for Y, 4s^2 4p^2 for Ge, 2s^2 2p^4 for O, and 5d^{10} 6s^2 6p^3 for Bi, while the remaining core electrons were tackled using the projector augmented-wave method.[48] For all calculations except those of the band structure, the Brillouin zone was sampled using only the k-point Γ, given the extremely high computational cost of hybrid DFT calculations when using the plane-wave formalism.[49] The cutoff value for the plane-wave basis set was 415 eV. The force and energy convergence criteria for structural and electronic relaxation were set to 0.01 eV Å^{-1} and 10^{-5} eV, respectively. The lattice parameters of bulk LiYGeO$_4$ were obtained by optimizing atomic positions, cell shape and cell volume of a unit cell of the material, using the residual minimization method with direct inversion in the iterative subspace, which is ideal for systems with more than a few tens of degrees of freedom that are already close to their ground state structure. In turn, electronic orbitals were optimized with the blocked-Davidson scheme, which is the default in VASP, and their occupancies were determined using a Gaussian smearing with a width of 0.01 eV.

3. Results and discussion

3.1 Band structure calculations

The lattice parameters calculated for LiYGeO$_4$ are a = 11.19 Å, b = 6.34 Å, and c = 5.07 Å. To the extent of our knowledge, the band structure of LiYGeO$_4$ has been calculated with standard GGA functionals only,[10] which are famously known to underestimate the bandgap of materials. Here, using the HSE06 functional, we intend to obtain a more reliable band structure
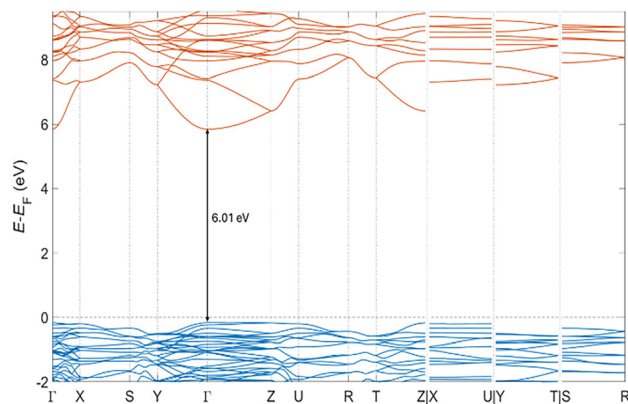
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Fig. 2 Calculated band structure of $LiYGeO_4$ using the HSE06 hybrid functional. The indicated energy is the calculated bandgap value, the blue/red lines indicate the occupied/unoccupied electronic states, and the horizontal dashed line lies at the Fermi energy.

Fig. 3 (a) Band energies at the Γ point of bulk $LiYGeO_4$, or with a Bi_Y^{3+} defect. The density of states of the defective model, projected onto the Bi atom, is also shown. (b) Band energies at the Γ point of models in which a Bi_Y^{3+} has a neighbouring V_{Li}^- or a Y_{Li}^{3+}. In the band energy diagrams, blue (red) lines denote occupied (unoccupied) levels. The dashed line is drawn at the Fermi energy.

of this material, which is shown in Fig. 2. The plot is remarkably similar to the one found by Qiao *et al.*,[10] albeit with a significantly wider bandgap, 6.01 eV, which is almost twice as much as the one calculated at the PBE level. Qualitatively, both methods predict the bandgap to be a direct one at the Γ point, and the top of the valence band to be rather flat, especially along the Γ–Z direction of *k*-points.

The most favourable location of Bi atoms in the $LiYGeO_4$ structure, as well as its most stable charge state, are at a Y site and Bi^{3+}, respectively.[10] The calculations of Qiao *et al.*[10] further showed that the formation energy of Bi_Y is at least 1.0 eV lower than that of the stoichiometric isomer Bi_{Li}–Li_Y. Here, we verified this fact and obtained a difference of 1.24 eV between the formation energies of these two defects. Furthermore, we went a step further and compared Bi_Y with Bi_{Ge}–Ge_Y, having concluded that the latter is thermodynamically disadvantageous by 2.90 eV. Thus, we completely covered the three possibilities of a Bi atom replacing any cation in $LiYGeO_4$.

Upon substitution of a single Y atom by a Bi one, the surrounding structure undergoes nearly no relaxation. In fact, the relaxation energy (the energetic difference between the Bi_Y model after and before atomic position optimization) is a mere 0.25 eV. Furthermore, the average Y–O distance in $LiYGeO_4$ was estimated at 2.27 Å, while the average Bi_Y–O one is 2.35 Å, *i.e.*, only 3% larger. Given the low Bi content of the LYGBi synthesized here (0.5 mol%), this substitution should have a negligible effect on the experimentally determined lattice parameters. Note that, although a Bi_Y defect alone has C_s symmetry (the same as the Y sites in the pristine material), the presence of neighbouring defects that are easily found on $LiYGeO_4$, such as Y_{Li} antisites, or Li vacancies or interstitials, further reduces the local symmetry to C_1. Nonetheless, this should have the same impact on the splitting of energy levels as a C_s environment.

We calculated the band energies at the Γ point of bulk $LiYGeO_4$ and with a Bi_Y defect, which are displayed in Fig. 3(a). Upon replacement of a Y atom by a Bi one in the simulation supercell, a localized state, associated with Bi^{3+} s orbitals, appears within the bandgap of the material, 4.91 eV below
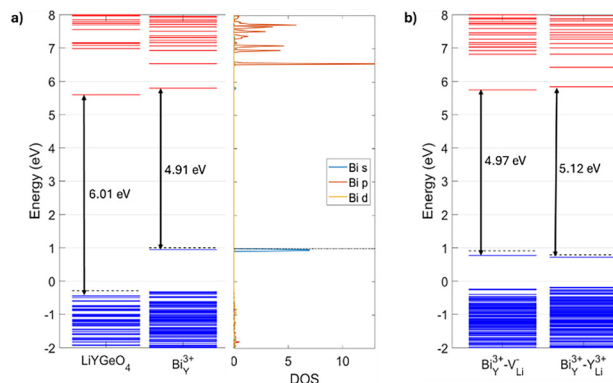
the conduction band minimum (CBM). In other words, the transfer of an electron from the ground state level of Bi^{3+} to the CBM should be associated with an energy of around 5 eV.

Given the high probability of occurrence of Li vacancies (V_{Li}) and Y_{Li} antisites, it should be common to find a Bi_Y near such a defect. In fact, given that (i) Li vacancies are very common, (ii) the most stable position of a Bi atom is at a Y site, and (iii) in pristine $LiYGeO_4$ the defect with the lowest formation energy is Y_{Li}, it stands to reason that a combination of these three defects, yielding Bi_Y–Y_{Li}, should be observable. Such a combination can be regarded as a region of the material in which a dopant Bi atom occupies a Y site, having pushed the Y atom towards a Li-vacant site. For this reason, we performed calculations using models with a combination of Bi_Y and a neighbouring V_{Li} or Y_{Li}. According to the results, both are stable – the reconstruction of the lattice around the defective area is minimal and accompanied by a small relaxation energy, of a few tenths of an eV, as in the case of a simple Bi_Y. In addition, the formation energies of Bi_Y–V_{Li} and Bi_Y–Y_{Li} are lower than that of Bi_Y. In particular, the formation energy of Bi_Y near a Li vacancy is lower by around 0.8 eV than in the absence of such a vacancy. This implies that Bi_Y is thermodynamically easier to form when the material has Li vacancies or, equivalently, that Bi_Y are attracted to regions of the material where there are Li vacancies. The most stable charge states of the two composite defects are Bi_Y^{3+}–V_{Li}^- and Bi_Y^{3+}–V_{Li}^{3+}. The respective band energies at the Γ point are illustrated in Fig. 3(b) and appear to be very similar to those of Bi_Y. Finally, according to our theoretical calculations, the energetic difference between having two Bi_Y replacing two nearest Y atoms and having two infinitely separated Bi_Y is a mere 1 meV. This implies that Bi_Y defects practically do not interact (*i.e.*, they do not attract or repel each other), and therefore can be distributed in any way in the $LiYGeO_4$ lattice.

3.2 Thermal, structural and elemental analysis

Thermal analysis was used to assess the behaviour of the precursor powders in the mixture with the desired stoichiometry
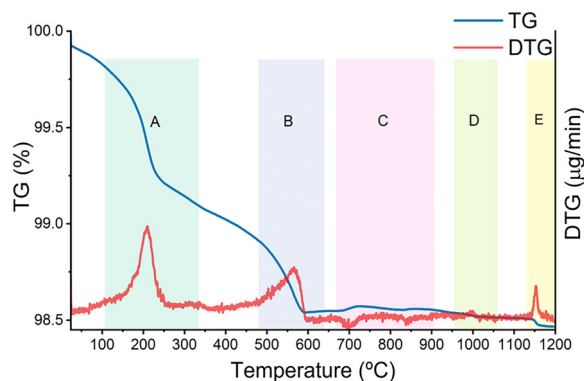
Fig. 4 Thermogravimetric and corresponding derivative curves for the LiYGeO$_4$ synthesis, starting from the precursor's powders in the used stoichiometry.

with temperature variation, as the TG and DTG curves shown in Fig. 4.
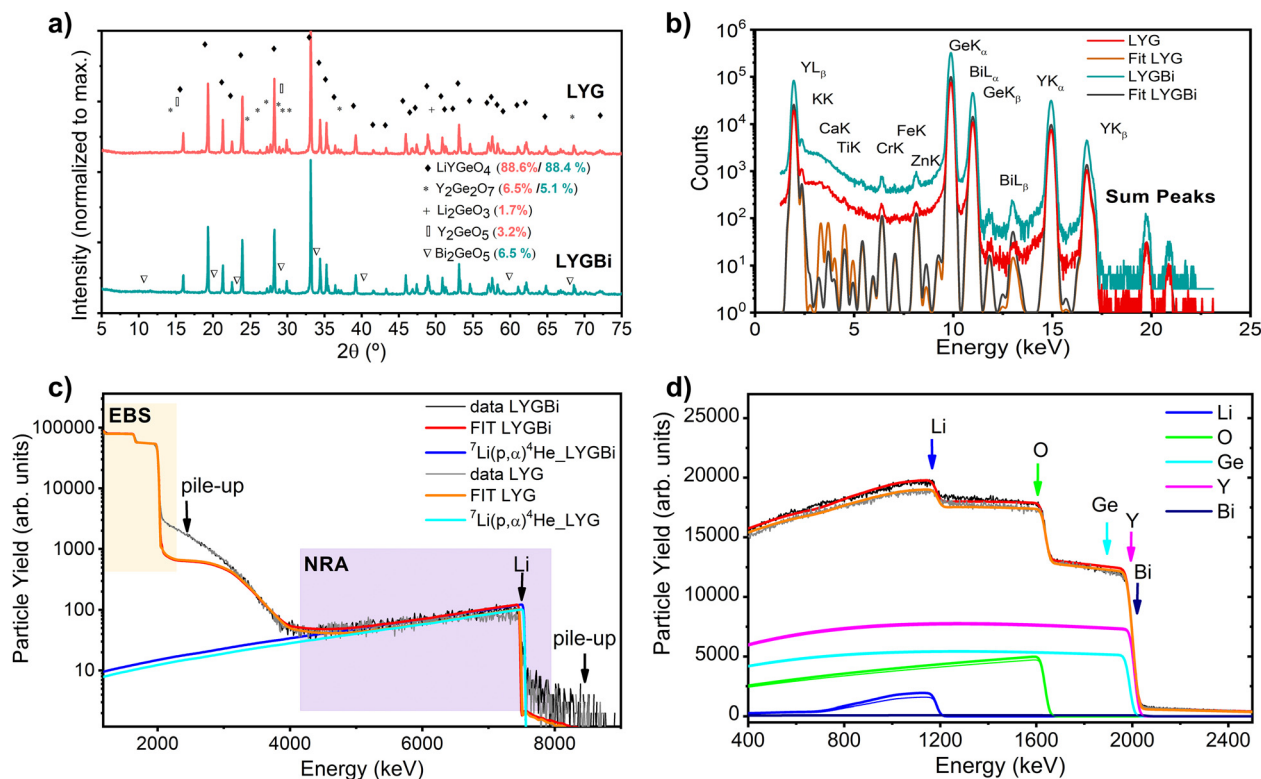
The TG curve clearly shows four major weight loss steps within areas A, B, D and E, and very small weight gains are identified within the temperature range of area C. The first weight loss of around 0.4% (area A) can be associated with the evaporation of residual water as well as CO$_2$ release.[50] The second major weight loss of 0.3% (area B) is related to the decomposition of Li$_2$CO$_3$ and conversion to LiO$_2$, followed by further rearrangement between the precursors/products (area C) since weight gain is observed within area C.[51] During this reorganization of the overall crystalline host, evaporation of lithium and/or of the remaining cations may take place to some extent.[51,52] The variation in weight observed around 1000 °C results from the crystallization of the final LiYGeO$_4$ (area D), culminating in a small weight loss (area E), possibly due to the further sublimation of lithium.[52,53]

XRD was performed to assess the phase composition of the prepared samples, and the results are depicted in Fig. 5(a). Table 1 presents the phase composition of the LYG and LYGBi samples. As can be seen by inspecting both the figure and the table, the produced samples exhibit a polyphasic nature. Nevertheless, the LiYGeO$_4$ (ICDD card 04-002-3479) crystalline phase is the predominant one ($\sim$88%), and the estimated lattice parameters are a = 11.1359 Å, b = 6.3305 Å and c = 5.0628 Å for LYG, and a = 11.1346 Å, b = 6.3315 Å and c = 5.0622 Å for LYGBi, well in line with previously reported ones,[11,53] and notably close to those resulting from our calculations. The tetragonal Y$_2$Ge$_2$O$_7$ (ICDD card 04-001-6776) phase is also present in small amounts in both samples. Besides, sample LYG also exhibits small percentages of Y$_2$GeO$_5$ (ICDD card 00-023-1484) and Li$_2$GeO$_3$ (ICDD card 00-029-0806) phases, while sample LYGBi shows evidence of the Bi$_2$GeO$_5$ (ICDD card 04-011-8707) phase, whose presence is most likely the result of an inhomogeneity in the distribution of Bi during the mixture of the precursors. Note that the Y$_2$Ge$_2$O$_7$, Y$_2$GeO$_5$ and Li$_2$GeO$_3$ phases are related to the LiYGeO$_4$ instability and the breakdown of the stoichiometry at high temperatures, especially because of the sublimation of lithium mentioned above.[14,30]

The Wyckoff positions of the ions in the LiYGeO$_4$ phase were also confirmed using XRD data, and these positions are compatible with the previous description, supporting that Li is located at the 4a site and Y occupies the 4c site. By comparing the peak position of the graphs in Fig. 5(a), a slight shift ($\sim$0.0079°) is observed. This shift is likely attributed to the incorporation of Bi into the lattice, as this ion is expected to occupy the position of the Y ion and has a slightly larger radius.

The stoichiometric analysis of the LYG and LYGBi samples was carried out using the simultaneous analysis of the particle yields. Since the samples are composed of light and heavy elements, different approaches were carried out. The quantification of the individual contributions of the heavier elements, namely Y, Ge and Bi, was assessed by a PIXE analysis, while the quantification of light elements was assessed through the combination of EBS and NRA. As shown in Fig. 5(b), from the PIXE spectra, it is possible to confirm the presence of Y, Ge, and Bi in both samples. Using the GUPIXWIN software code,[41] the concentration of these elements was accurately estimated (see Table 2). Besides the host and dopant elements, small quantities of other undesired elements (Fe and Zn) were identified, with percentages lower than 0.01%, so their information was not included in Table 2 but can be seen in Fig. 5(b). It is worth noting that the presence of Bi in the undoped sample is related to trace contamination of the raw Y$_2$O$_3$ and Li$_2$CO$_3$ precursors, as reported in the corresponding Certificate of Analysis provided by the supplier (analysis method: mass spectrometry). The same is valid for the remaining impurities that were identified, most of them originating from trace contaminants of the used raw materials. The relative concentrations of Y, Ge and Bi estimated from the fit of PIXE spectra were subsequently used to define the fitting model for EBS and NRA spectra in order to estimate the concentration of all elements using the NDF software. In defining the fitting model, contaminations of Cr, Fe, K, and Ca, identified in the PIXE spectra, were not considered due to low concentrations. It is also noteworthy that the individual EBS and NRA spectra fittings, represented by the red and orange lines in Fig. 5(c) and (d), were performed simultaneously. These fits, in addition to allowing the evaluation of stoichiometry, clearly suggest that the samples exhibit high homogeneity in depth. EBS events enable identifying and quantifying the light elements by analysing the barriers associated with corresponding backscattered particles.[37,54,55] In addition, RBS yields enable a similar analysis involving heavier elements.[56] However, as can be seen in Fig. 5(c), in the case of the complex matrices involving different light and heavy elements, it is challenging to accurately identify and quantify the elemental concentration of each element due to the overlap of the barriers, as in the case of Bi, Y and Ge, or due to their low signal, as in the case Li. In particular, for the Li, to address this limitation, complementary NRA analysis of the ^{7}Li(p,α)^{4}He nuclear reaction was performed. As can be seen in Fig. 5(c), the high Q-value of the reaction[38] enables the clear separation of the induced alpha particle yield from those associated with the proton yields elastically backscattered by heavier elements such as Y, Ge, Bi, and by light elements such as O and Li. Despite the much lower

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Fig. 5 (a) X-ray diffraction patterns of the produced samples. The data are vertically shifted for clarity. (b) PIXE data for both LYG and LYGBi samples. For the LYG sample, the data demonstrate the existence of Bi contamination. The data are shifted vertically for clarity. (c) Particle spectra collected with a 2.1 MeV H$^+$ beam. The data correspond to the EBS with the NRA spectra obtained with the detector located at $-165°$ and their respective fits (red and orange lines). In addition to identifying the different contributions, the input associated with the signal resulting from the overlap of two backscattered particles hitting the detector simultaneously, referred to as pile-up, is also identified. (d) EBS spectra, and the different contributions estimated from the fit using the NDF software. The latter two spectra were obtained with a similar detector positioned at 165° and with higher signal amplification to achieve higher resolution in this region, thereby enabling more precise identification of the different barriers.

Table 1 Crystalline systems and phase composition of the prepared samples obtained from the XRD data

Sample	Composition	Crystal system	%
LYG	LiYGeO$_4$	Orthorhombic	88.6
	Y$_2$Ge$_2$O$_7$	Tetragonal	6.5
	Y$_2$GeO$_5$	Monoclinic	3.2
	Li$_2$GeO$_3$	Orthorhombic	1.7
LYGBi	LiYGeO$_4$	Orthorhombic	88.4
	Y$_2$Ge$_2$O$_7$	Tetragonal	5.1
	Bi$_2$GeO$_5$	Orthorhombic	6.5

Table 2 Concentration of the various elements estimated from the fits of the individual EBS and NRA spectra performed simultaneously using the NDF software and from the fit of the PIXE spectra using the GUPIXEWIN software

Sample	Elements	wt%	Nominal wt%
LYG	Li	1.63 ± 0.46	2.97
	Y	38.75 ± 0.88	38.08
	Ge	32.91 ± 0.75	31.10
	O	26.64 ± 0.76	27.40
	Bi	0.066 ± 0.002	—
LYGBi	Li	2.40 ± 0.19	2.97
	Y	37.75 ± 0.37	38.08
	Ge	32.00 ± 0.31	31.10
	O	27.65 ± 0.32	27.40
	Bi	0.213 ± 0.002	0.45

sensitivity of EBS for Li quantification, a correspondence is obtained from the analysis of both Li-EBS and Li-NRA yields.

Furthermore, Fig. 5(b) suggests that the Bi content in the intentionally doped sample differs from the nominal one (0.213 wt% *vs.* ∼0.447 wt%). The lower content is likely related to the loss of this element during the synthesis process.[57] As can be seen, a significant amount of Li is also lost during the synthesis process, especially in the case of sample LYG, resulting in Li vacancies in the structure, as already anticipated by other works, although no final quantification was performed.[14,58] The remaining cations display wt% close to the nominal ones, with negligible differences that can be

accounted for by the error of the measurement. Nevertheless, as pointed out by the thermal analysis, reorganization of the crystalline network takes place during the synthesis, and some losses may have occurred during the process.

Please note that the values presented in Table 2 represent average values obtained from the probed samples. Consistent with the XRD data, which confirmed the presence of the

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Bi_2GeO_5 phase, the Bi distribution in the LYGBi sample is non-uniform. Analysis of various random spots revealed significant variations in Bi content, while the Ge and Y concentrations remained relatively constant throughout the sample.

3.3 Optical assessment

3.3.1 RT diffuse reflectance, PL and PLE. The optical characterization of the LYG and LYGBi samples was first carried out at RT by diffuse reflectance, PL and PLE, as displayed in Fig. 6(a). The diffuse reflectance spectra are depicted in absorption mode after conversion using the Kubelka–Munk model.[59,60] The PLE spectra were obtained by monitoring the maximum of the emission band at 358 nm ($\sim$3.46 eV), and the PL spectra were acquired under 250 nm (4.96 eV) and 300 nm ($\sim$4.13 eV) photon excitation. The optical results were found to be very similar for the two samples. Considering the absorption spectra, a strong and steep decrease in intensity is observed at shorter wavelengths (from shorter to longer wavelengths), particularly in the LYG sample, likely related to bandgap absorption. Yet, it is important to note that the measuring equipment, which makes use of an integrating sphere, only provides reliable results from 200 nm onward. Below this threshold, the data are not trustworthy, making the accurate determination of the bandgap energy more complicated. In the literature, the Tauc method is usually used to estimate the bandgap of semiconductors. However, since the present materials are known to have a large concentration of defects, this methodology may not be the most adequate, as the presence of defects will lead to the emergence of tail states near the band edges, resulting in the underestimation of the correct value. As a result, the actual bandgap value may be higher than the one estimated by this method. Still, for the sake of comparison, this approximation was used in the present work (see Fig. 6(b)). Using this approach, we obtained a value of approximately 6.05 eV, considerably higher than the previously reported values in the literature (4.5–4.9 eV).[15,16,18] Nevertheless, comparing the spectral shape of the absorption/reflectivity curves published in the literature with the ones obtained here, the resemblances are significant. For instance, the absorption spectrum presented by Shi *et al.*[18] also presents its maximum near $\sim$200 nm, displaying, however, a more pronounced tail of states, which led to an underestimation of the bandgap value. Indeed, the experimental value assessed in the present work is in line with the one estimated by our first-principles calculations (6.01 eV), as well as the one reported by Qiao *et al.*[10] and Zhou *et al.*[11] Additionally, the steep slope of the absorption suggests that this material should be of a direct bandgap, consistent with the hybrid DFT results discussed above.
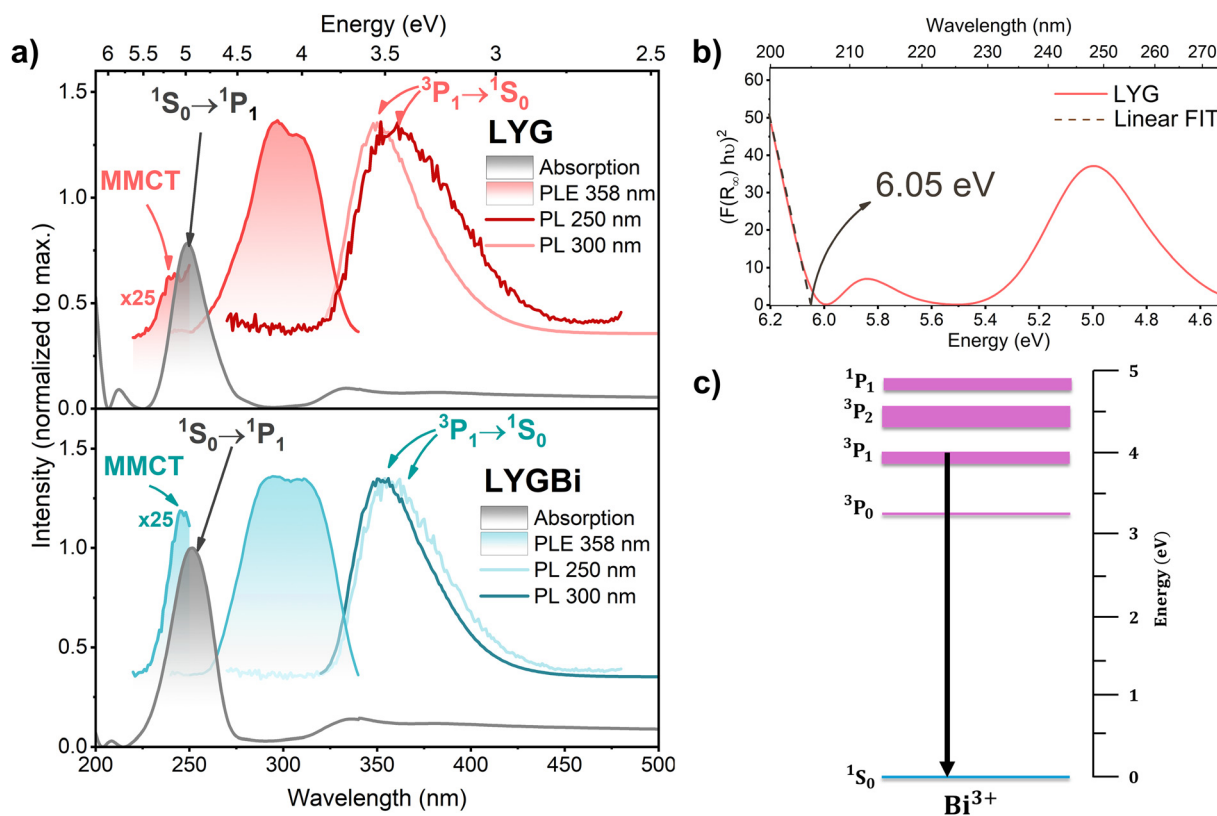

Fig. 6 (a) Absorption (grey filled lines), photoluminescence excitation (blue and red full-filled lines, monitored at 358 nm ($\sim$3.46 eV)) and photoluminescence (solid lines, acquired under 250 nm (4.96 eV) and 300 nm ($\sim$4.13 eV) photon excitation) spectra obtained for the studied samples. The spectra were shifted vertically for the sake of clarity. (b) Bandgap energy estimation from the Tauc plot method. The linear part of the plot is extrapolated to the energy axis. (c) Energy level diagram of the Bi^{3+} ion, the vertical arrow represents the $^3P_1 \rightarrow {}^1S_0$ transition of the Bi^{3+} ion. Note that the presence of the Bi^{3+}-related transitions in the LYG sample is related to the presence of Bi contaminants in the raw precursors.

Besides this absorption, both samples exhibit a lower intensity absorption band at $\sim$212 nm ($\sim$5.85 eV), followed by a strong asymmetric absorption band peaked around 250 nm ($\sim$4.96 eV). The latter was seen to widen and increase its intensity with increasing dopant concentration, as confirmed here by the high intensity of this band in the LYGBi sample, which suggests a correlation with the presence of the Bi^{3+} ion. By comparing the present data with the energy positions estimated by Qiao et al.[10] for the most common defects present in this compound, the low-intensity band at $\sim$5.85 eV ($\sim$0.68 eV below the CBM) is compatible with the energy position estimated for Ge_{Li} antisites, while the strong absorption band is most likely related to the $^1S_0 \rightarrow {}^1P_1$ transition of the Bi^{3+} ions.[24] As stated above, when considering the Russell–Saunders (LS) coupling scheme, only the transitions from the $^1S_0$ ground state to the $^1P_1$ excited state are allowed by the electric dipole. However, the strong spin–orbit coupling leads to interaction between the $^3P_1$ and $^1P_1$ levels, resulting in mixed states. Thus, the $\Delta S = 0$ rule is relaxed, and the $^1S_0 \rightarrow {}^3P_1$ becomes partially allowed. Likewise, the interaction with the lattice phonons also makes the $^1S_0 \rightarrow {}^3P_2$ transition possible.[24] As so, and according to the diagram in Fig. 6(c), the two strongly asymmetric bands centred at $\sim$310 nm and 296 nm ($\sim$4 eV and $\sim$4.19 eV, respectively) observed in the PLE spectra are probably associated with the $^1S_0 \rightarrow {}^3P_1$ and $^1S_0 \rightarrow {}^3P_2$ transitions, exhibiting an energy separation around $\sim$100–200 meV (considering the peak position of the maximum of the excitation bands and their similar full widths at half maximum). The broadening of the bands can be associated with the vibronic levels that originate from the interaction with the lattice phonons.[24] Moreover, one must take into account that, when introduced in the Y sites, the Bi ions will be placed at a C_s or C_1 symmetry, as mentioned in Section 3.1, in which the three-fold degenerate $^3P_1$ and $^1P_1$ and the five-fold degenerate $^3P_2$ J-levels will be completely split.[24]

Additionally, a closer inspection of the PLE spectra also reveals the presence of an absorption band at 245 nm ($\sim$5.06 eV), which can be assigned to a metal-to-metal charge transfer (MMCT) transition. Indeed, in the case of Bi ions, besides the $6s^2 \rightarrow 6s^16p^1$ interconfigurational transitions, other luminescence signals are commonly identified in different matrices, as is the case of this MMCT transition.[24,25,61] As reported by other authors, above the $^1S_0 \rightarrow {}^1P_1$ band, but below the absorption of the host, a further band is generally encountered and attributed to an MMCT state. This band generally occurs at fairly low energy and may overlap or replace the absorption bands associated with the Bi^{3+} ion intraionic transitions.[24] The MMCT state results from the interaction between the Bi^{3+} levels and the electronic levels of host metal cations having d^0 or d^{10} configuration, which, in the present case, corresponds to Ge^{4+} ([Ar] $3d^{10}$).[25,61] Boutinaud[61] reported a quantitative model for estimating the position of the MMCT energies of Bi^{3+} ions in different oxide hosts based on the electronegativity of the metal cations (M^{n+}), the coordination number of the Bi^{3+} and the host M^{n+}, as well as the shortest interatomic distances between them.[24,61] Note that such a

model is only indicative for oxides. Therefore, the electronegativities of the host cations strongly influence the energies of charge transfer excitations.[24,61] Considering Ge^{4+} as the metal cation involved in this transition, this could be represented as $Bi^{3+}(6s^2)/Ge^{4+}(d^{10}) \rightarrow Bi^{4+}(6s^1)/Ge^{3+}(s^1)$.[61] Qiao et al.[10] reported that in the case of $LiYGeO_4$, this band should occur around 225–234 nm, which fairly matches the value obtained for the PLE band at 245 nm ($\sim$5.06 eV). Such an MMCT transition corresponds to the transfer of an electron from the ground level ($^1S_0$) of the Bi^{3+} ion to the CBM of the host.[24,61] Assuming that this is indeed the present case, this band allows us to locate the $^1S_0$ level within the bandgap of the material reported here. The value of 5.06 eV is compatible with the diagrams in Fig. 3 and with the energy estimated by Qiao et al.,[10] which reported that the Bi^{3+} ground level was $\sim$5 eV below the CBM. According to Boutinaud's models,[61,62] when this transition is present at energies higher than that of the $^1S_0 \rightarrow {}^3P_1$ transition, excitation through this band can be followed by emission either from the MMCT or via the $^3P_{1,0} \rightarrow {}^1S_0$, after transfer of the absorbed energy. Such findings may also explain why the PersL reported here (see Section 3.3.4) is mostly promoted when the samples are excited with photons with wavelengths $\sim$230–250 nm and not when excited directly to the ion excited levels, as was also reported in previous works,[18] indicating that the PersL mechanism must be mediated by the conduction band of the host.
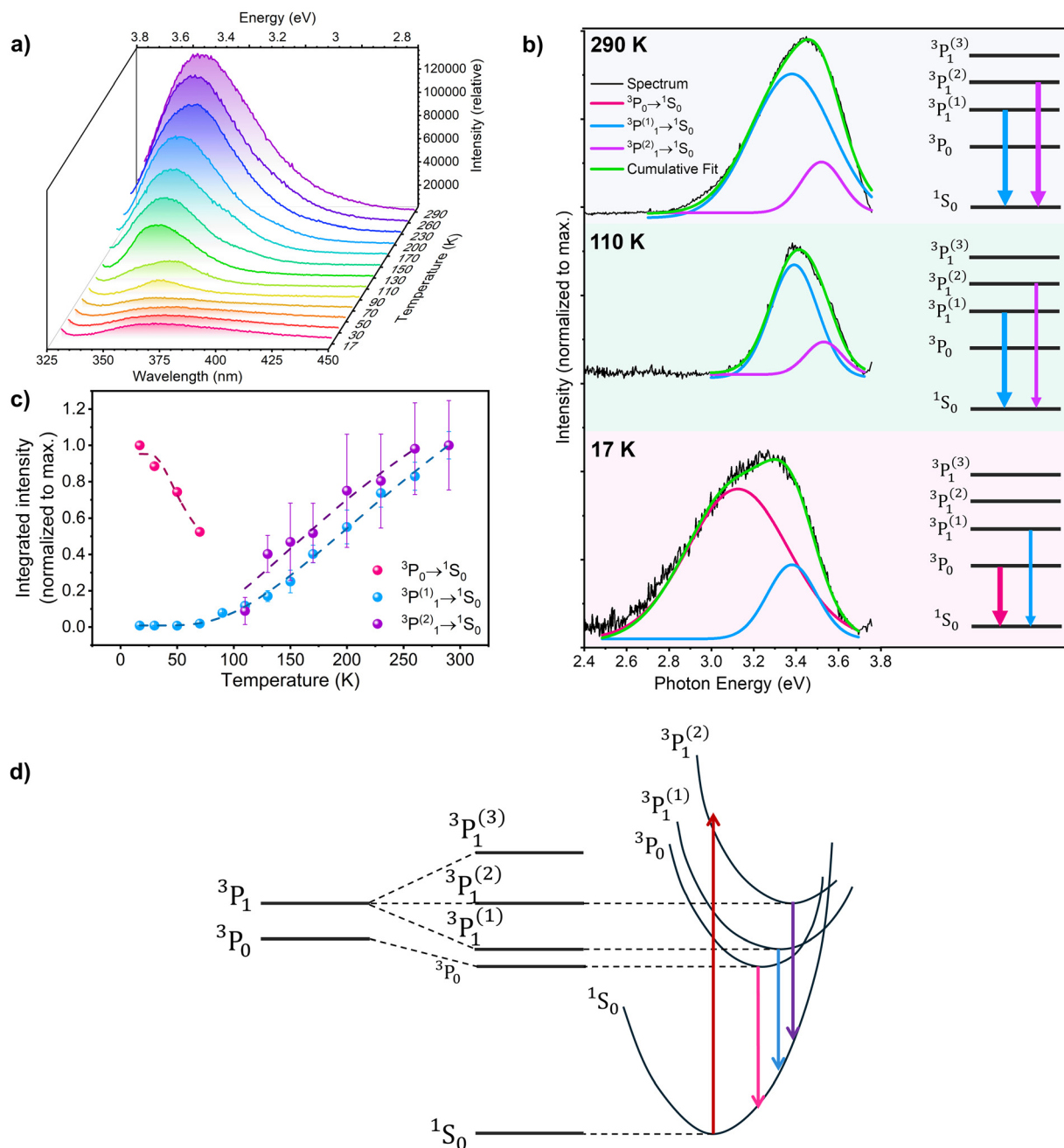
In the case of the PL acquired at both 250 nm and 300 nm photon excitation, the emission spectra display a maximum at $\sim$358 nm in both cases, with only a slight shift in the peak position, which will be further discussed in the next section. This value is in agreement with the previous reports in the literature regarding Bi-doped $LiYGeO_4$ samples, also presenting a similar spectral shape.[16,18] As already mentioned in the Introduction section, such a band has been attributed to the $^3P_1 \rightarrow {}^1S_0$ transition, with no further discussion. Nevertheless, the possibility of different Bi centres contributing to the observed transitions cannot be ruled out, as it is known that the sample contains other phases. Indeed, these additional phases may also present luminescence bands associated with the incorporation of Bi^{3+} ions. However, as noted by previous authors, the observed UV emission is unlikely to arise from the inclusion of Bi^{3+} in the secondary phases $Y_2Ge_2O_4$ or Bi_2GeO_5, in which this ion is known to emit in the green[63] or infrared[64] spectral regions. Conversely, for the Y_2GeO_5 phase, Bi^{3+}-related emission is expected to occur in the UV spectral region, peaking at approximately 370 nm.[63,65] Therefore, a minor contribution from this phase cannot be completely ruled out. The same applies to the Li_2GeO_3 phase. To the best of the authors' knowledge, there are currently no reports on Bi incorporation as a dopant in this matrix, thus, no comparison could be made that allowed us to discard such influence fully. Notwithstanding, as the Y_2GeO_5 and Li_2GeO_3 phases are not present in the LYGBi sample, and the obtained spectra for the two samples are very similar, if this contribution is indeed present, it should be negligible.

3.3.2 Temperature-dependent PL. To provide deeper insights into the origin of the observed luminescence, the sample LYG was selected for the temperature-dependent PL

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To further analyse the observed transitions, the spectra were deconvoluted into Gaussian functions, as displayed in Fig. 7(b). At 17 K, the emission maxima are shifted to longer wavelengths, peaked at ∼380 nm (∼3.26 eV), when compared to the RT spectrum. It can also be seen that the band is strongly asymmetric, pointing to the overlap of more than one electronic transition. Moreover, at 17 K, the emission is much broader than in the case of RT. The spectral deconvolution allowed us to conclude that the emission is dominated by a component peaked at ∼3.13 eV (pink line in Fig. 7(b)), with a small contribution from another band at ∼3.38 eV (blue line in Fig. 7(b)). For temperatures higher than 70 K, the low-energy component is no longer observed. Note that up to 70 K, the overall PL intensity decreases, while from 90 K to RT, an increase in the intensity was progressively observed (Fig. 7(a)), accompanied by a high-energy shift up to RT. At 110 K, the main emission shifts to high energy, compared to emission at 17 K, with the emission band now peaked at 363 nm (∼3.42 eV). By deconvolution of this band, the appearance of a higher-energy component (at ∼3.5 eV, violet line) becomes clear, which is seen to contribute to a shift in the overall emission towards higher energies, in line with what was displayed in Fig. 6(a) for the RT data. When raising the temperature to RT the main maximum is peaked at 358 nm (∼3.46 eV) and composed of two components, which were attributed to the splitting of the $^3P_1$ level due to the crystal field, since the ion is placed in a C_s/C_1 symmetry, thus explaining the slight shift in the peak of the emission band depending on the excitation wavelength observed in Fig. 6(a).

Although the transition from the $^3P_0$ level to the ground state is not expected to contribute to absorption due to its forbidden nature, previous works on other matrices[24,33,34,66] have reported that at lower temperatures (<80 K), and as a result of mixing the $^3P_0$ and $^3P_1$ states induced by the second derivatives of strains caused by nearby point defects, the $^3P_0 \rightarrow {}^1S_0$ transition can be observed. Indeed, several authors have reported this transition as the dominant one at low temperatures.[24,66] Therefore, in the present case, the low-energy component is suggested to correspond to such a transition. According to Boulon et al.,[33] despite being considered metastable, the $^3P_0$ level does make an important contribution to the total emission at very low temperatures, in which the $^3P_1 \rightarrow {}^1S_0$ transition has a low intensity due to the lower population of the $^3P_1$ level at such temperatures. As the temperature increases, the thermal population from the $^3P_0$ to the $^3P_1$ levels takes place, resulting in the increased intensity of the $^3P_1 \rightarrow {}^1S_0$ transition, which is accompanied by a blue-shift of the overall PL emission. The temperature dependence of the PL intensity of each transition is depicted in Fig. 7(c), highlighting the mentioned behaviour. Moreover, it also shows that both high-energy transitions increase their intensity up to RT, presenting nearly the same intensity dependence, which is a strong indication that both are related to the $^3P_1$ excited level, being assigned to the lower energy $^3P_1^{(1)}$ sub-level and the intermediate $^3P_1^{(2)}$ at higher energies, respectively. This is further corroborated by their energy separation (∼140 meV).

A schematic representation of these levels is shown in Fig. 7(d). Hence, the emission observed at RT is composed of an overlap of the $^3P_1^{(1)} \rightarrow {}^1S_0$ and $^3P_1^{(2)} \rightarrow {}^1S_0$ transitions, explaining the broad and asymmetric band that is typically observed in this material.[16,18,27]

However, it is important to note that the above interpretation represents the explanation that seems most likely to the authors, taking into account the data presented here. Nevertheless, the possibility that the emission observed at RT arises from an overlap of the $^3P_2$ and $^3P_1$ states cannot be entirely excluded. Even so, the fact that the temperature behaviour is the same for the two higher energy components strongly supports the present assumption.

In general, and as shown in Fig. 7(c), by exciting the sample with 325 nm photons, the intensity of the $^3P_0 \rightarrow {}^1S_0$ decreases over the entire temperature range in which this transition is observed (17 K to 70 K), while, as already mentioned, the $^3P_1^{(1)} \rightarrow {}^1S_0$ and $^3P_1^{(2)} \rightarrow {}^1S_0$ increase up to RT. As mentioned above, this behaviour indicates the presence of a thermal population mechanism among the energy levels of the Bi^{3+} ion. As so, the effect of thermal depopulation of the $^3P_0$ state on its luminescence intensity, i.e., the non-radiative competitive processes involved in its temperature dependence, can be expressed by the classical model,[1,67,68]

$$\frac{I(T)}{I_0} = \frac{1}{1 + C\exp\left[-E_{\text{depop}}/k_B T\right]}, \quad (1)$$

in which T is the absolute temperature, I_0 is the intensity at 17 K, C is the ratio of the effective degeneracies of the electronic levels, k_B the Boltzmann constant and E_{depop} is the activation energy for the non-radiative processes. However, for the $^3P_1^{(1)}$ and $^3P_1^{(2)}$ states, in which only a thermal population with increasing temperature was observed, the intensity behaviour with increasing temperature can be described by the model,

$$\frac{I(T)}{I_0} = 1 + C_{\text{pop}}\exp\left[-\frac{E_{\text{pop}}}{k_B T}\right] \quad (2)$$

where E_{pop} corresponds to the activation energy for the population of the states. The values derived from the best fits to the proposed models are presented in Table 3. The values obtained for the activation energies corroborate the assumption of thermalization from the $^3P_0$ to the $^3P_1$ sublevels with increasing temperature.

3.3.3 Thermoluminescence analysis. In order to obtain more information about the trap states present in the LYG and LYGBi samples, TL measurements were performed by monitoring the Bi^{3+} emission. The T_M–T_{STOP} method was employed to identify the number of different traps present in two of the prepared samples, as mentioned in the Experimental details section. Fig. 8(a) and (b) show the evolution of the glow curves as a function of the pre-heating temperature T_{STOP}, which allows the gradual depopulation of the trapping states. The temperature position T_M of the first maximum of the glow curves was monitored for increasing values of the pre-heating temperature T_{STOP}, revealing the presence of at least four traps. Moreover, this plot also provides insight into the kinetics

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Table 3 Activation energies obtained for the population/depopulation processes of $^3P_0 \rightarrow {}^1S_0$, $^3P_1^{(1)} \rightarrow {}^1S_0$ and $^3P_1^{(2)} \rightarrow {}^1S_0$ transitions excited with 325 nm ($\sim$3.81 eV)

Transition	E_{pop} (meV)	E_{depop} (meV)
$^3P_0 \rightarrow {}^1S_0$	—	9.3 ± 3.5
$^3P_1^{(1)} \rightarrow {}^1S_0$	36.0 ± 1.6	—
$^3P_1^{(2)} \rightarrow {}^1S_0$	25.1 ± 15.6	—

governing the TL phenomenon, namely regarding the likelihood of an electron getting detrapped and/or retrapped. In particular, if the retrapping processes are negligible, the trap is said to conform to 1st order kinetics, whereas if the detrapping and retrapping processes have similar probabilities, the trap is said to follow 2nd order kinetics. The T_M–T_{STOP} results suggest that trap 1 follows 2nd order kinetics due to the smooth increase in T_M, whereas traps 2–4 are likely closer to 1st order kinetics due to the abruptness of the corresponding steps in the plots. On the other hand, the results are very similar for both samples, which suggests that the trapping system is likely similar for each of them.

Bearing this in mind, the curve was fitted to the sum of four general-order[69] TL peaks using the TL-Fitter software, which is freely available on GitHub,[70] according to the following model:

$$I(T) = A + \sum_{i=1}^{4} \frac{n_{0i} S_i''}{\beta} \exp\left(-\frac{E_i}{kT}\right)$$

$$\times \left[1 - (1 - b_i) \frac{S_i''}{\beta} \int_{T_0}^{T} \exp\left(-\frac{E_i}{kT'}\right) dT' \right]^{\frac{b_i}{(1-b_i)}}, \quad (3)$$

where n_{0i}, S'', E_i and b_i are, respectively, the concentration, frequency parameter, activation energy and kinetics order parameter for the ith trap, and k, β, and T_0 are Boltzmann's constant, the heating ramp and the initial temperature (300 K), respectively. The A parameter accounts for a possible constant background due to the dark current of the photomultiplier tube or the eventual contribution of the blackbody radiation. The fits are shown in Fig. 8(c) and (d) for each of the two measured samples.

Considering two of the common acceptance criteria for TL fitting, namely the TL figure of merit below 3% and the acceptability parameter below 0.40, the fitting quality is quite good.[71,72] The obtained activation energies E_i are shown in Fig. 8(c) and (d) next to the corresponding peaks. Comparing both samples, the estimated energy values differ at most by 0.12 eV,
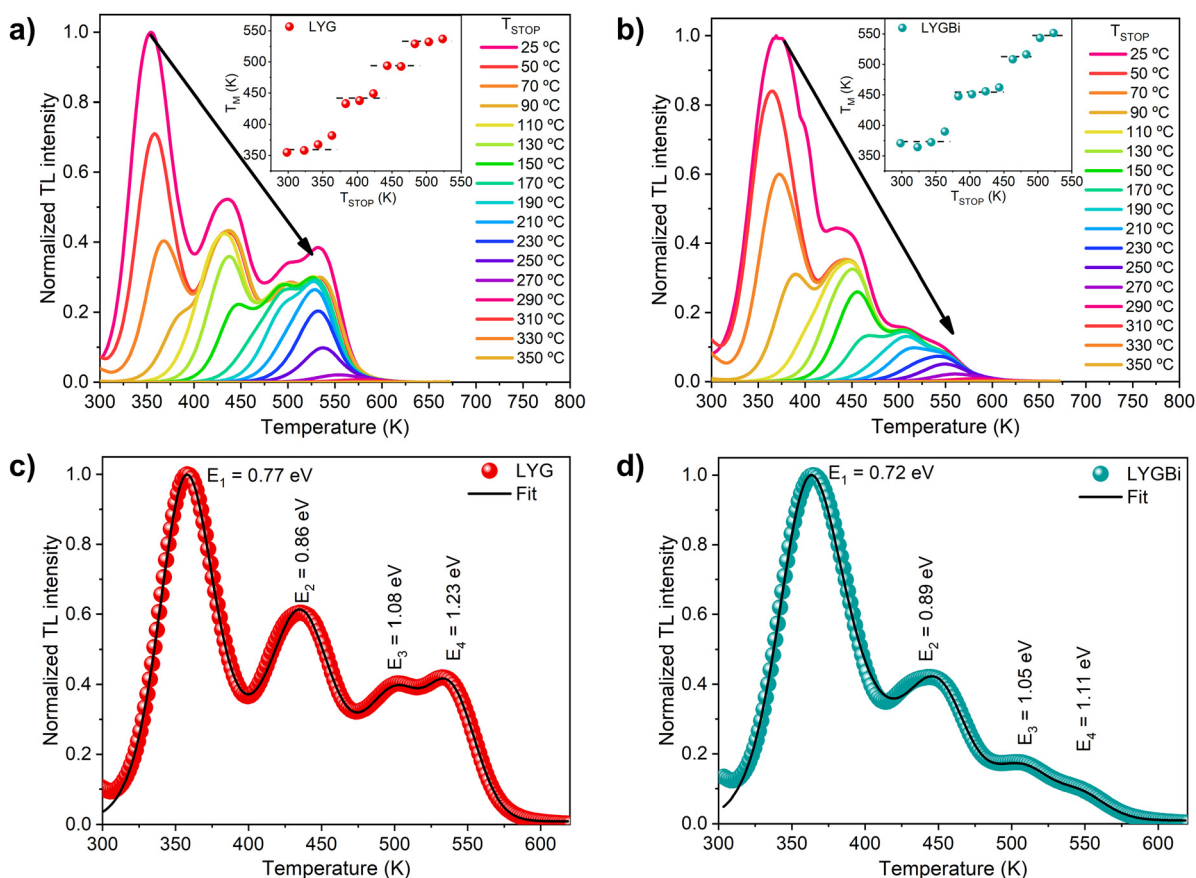
Fig. 8 Thermoluminescence study of the two produced samples (LYG and LYGBi) by the T_M–T_{STOP} method (panels a and b) and by glow curve computational fitting (panels c and d). The insets in the upper panels show the evolution of the temperature position of the first, T_M, as a function of the pre-heating temperature, T_{STOP}, used to gradually depopulate the traps.

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which further suggests that the traps are indeed similar in these samples. The differences observed in the glow curves are likely related to differences in the trap concentrations and in the frequency parameters, which depend on the structural properties of the samples.

It should be noted that the extraction of unique parameters from glow curve fitting is notoriously difficult due to the ambiguity in the role of different parameters, considering the number of free variables. However, the energies obtained from the fits are in close agreement with those reported in the literature for this material. For instance, Lyu and Dorenbos[73] report a trap in $LiYGeO_4$ with an activation energy of 0.76 eV, which is attributed to Bi^{2+} and compares very well with trap 1 in the present work. On the other hand, the results are also quite close to several traps estimated by Qiao et al.[10] using DFT. In particular, trap 1 can tentatively be assigned to the +/0 level of Ge_{Li}, (predicted at an energy of 0.71 eV), trap 2 to the +/0 level of Li_i (predicted at an energy of 0.84 eV) and traps 3 and 4 to the +/− and 2 +/+ levels of Y_{Li} (predicted at energies of 1.08 and 1.25 eV, respectively). Notably, the latter is expected in LYGBi, as the incorporation of Bi into the lattice may displace Y atoms, which in turn may occupy Li sites.[16,18] Moreover, the obtained activation energy for trap 1 is also close to the absorption band measured in these samples (absorption band identified at $\sim$212 nm/$\sim$5.85 eV, which corresponds to $\sim$0.68 eV below the CBM), as discussed above, which further stresses its role in the PersL phenomenon.

3.3.4 Persistent luminescence. Persistent luminescence was recorded after excitation of the samples for 10 minutes with photon wavelengths of 250 nm (4.96 eV) and 300 nm (4.13 eV). It was observed that the latter, i.e. direct excitation of the excited states of the Bi^{3+} ion, does not promote PersL. However, with the 250 nm photon excitation, it was possible to activate this phenomenon, and the intensity of the transitions remained detectable for over 7 hours in both LYG and LYGBi samples. Fig. 9 illustrates, for both samples, the RT intensity

decay of luminescence after ceasing the excitation (afterglow). It is also important to note that the PersL profile resembles the one identified for steady-state PL at RT, indicating that the observed emission is due to a combination of the $^3P_1^{(1)} \rightarrow {}^1S_0$ and $^3P_1^{(2)} \rightarrow {}^1S_0$ transitions.

Analysing the decay profiles reveals a rapid intensity drop within the initial minutes, with a $\sim$97% reduction after 1 hour, after which the decrease of the PersL intensity becomes more gradual over time. To investigate the mechanism behind this behaviour, namely the defect trap level distribution involved in capturing and de-trapping of the photoexcited carriers, the measured data were analysed. The best fit of the decay curves was obtained using a combination of stretched exponential functions (see Fig. 9), following the equation,

$$I(t) = \sum_{i=1}^{n} I_i \exp\left[-\delta_i \left(\frac{t}{\tau_i}\right)\right] \tag{4}$$

where $0 < \delta < 1$ stands for the dispersion factor, in which $\delta = 1$ is analogous to a simple exponential decay, and considering $n = 4$. The fitting parameters are presented in Table 4. The estimated dispersion factors varied between 0.32 and 1. These results demonstrate that we face a complex and disordered system, in which a wide distribution of trap levels is available for the storage of the charge carriers, in line with what was stated above.[1,74,75] The measured decay times are likely to be related to the duration that the charge carriers spend in the traps, with the faster decays corresponding to shallower traps. It is important to note that no intrinsic lifetime was recorded in this case since, as reported by Shao et al.,[16] the lifetime of the transition is of the order of hundreds of nanoseconds. Such a fast intrinsic lifetime was expected given that the $^3P_1$ level is already split, and so, the transitions between the sublevels of this state and the ground state have a high probability of occurring, leading to very short lifetimes. Hence, this lifetime could not be assessed in the present measurements since the equipment used and the afterglow measurement method do not provide time resolution on the ns scale. The fact that several decay times were found points to the presence of several trap defects that participate in the PersL process. Indeed, four distinct decay times were found, in line with the four discrete levels that were estimated by the above TL measurements. We can thus infer that the shorter decay times correspond to
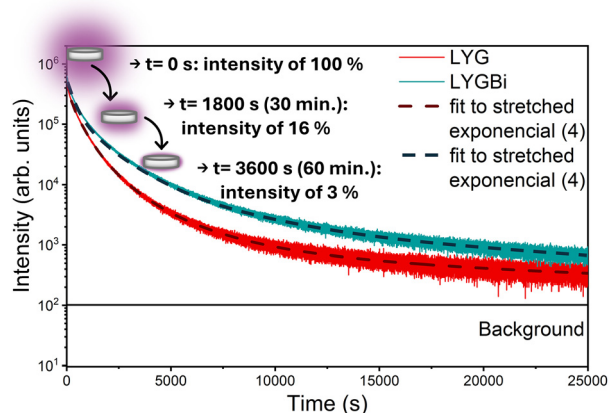

Fig. 9 Decay profiles acquired for 7 h after exciting the samples with 250 nm photons for 10 min. The illustration of the pellets depicts a representation of the reduction in emission over time. The percentage intensity values shown in the figure are relative to the initial intensity measured at $t = 0$ s (immediately after the excitation was turned off).

Table 4 Fit parameters of the data presented in Fig. 9. using a sum of stretched exponentials (eqn (4))

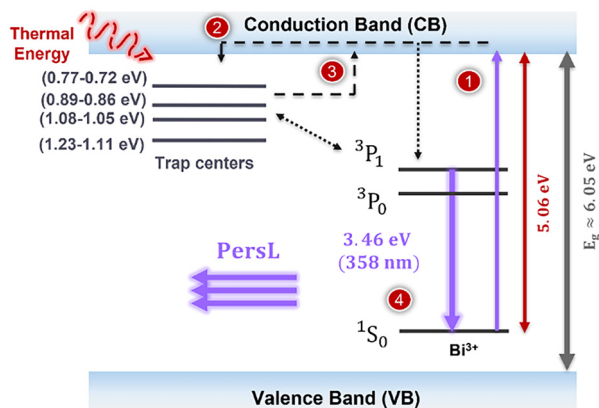| Sample	τ (s)	δ
LYG	$\tau_1 = 3.0 \pm 0.5$	$\delta_1 = 0.733 \pm 0.159$
	$\tau_2 = 136.8 \pm 4.1$	$\delta_2 = 0.198 \pm 0.005$
	$\tau_3 = 342.5 \pm 7.7$	$\delta_3 = 1.000 \pm 0.047$
	$\tau_4 = 892.2 \pm 62.2$	$\delta_4 = 0.750 \pm 0.015$
LYGBi	$\tau_1 = 137.9 \pm 4.0$	$\delta_1 = 0.74 \pm 0.03$
	$\tau_2 = 310.1 \pm 4.3$	$\delta_2 = 0.32 \pm 0.45$
	$\tau_3 = 312.9 \pm 90.7$	$\delta_3 = 0.77 \pm 0.04$
	$\tau_4 = 343.1 \pm 1.6$	$\delta_4 = 0.78 \pm 0.07$

Fig. 10 Schematic diagram illustrating the PersL phenomenon in LiYGeO$_4$: Bi upon excitation in the MMCT band. The depths of the defect traps were labelled according to the absorption and TL results. The PersL phenomenon involves exciting electrons from the ground state of Bi^{3+} ($^1S_0$) to the CBM (process 1), followed by their capture by defect traps below the CBM (process 2). Once the excitation ceases, thermal energy releases the trapped electrons to the CB again or directly to excited states of Bi^{3+}, leading to the emission of UVA PersL as the electrons return to the ground state (processes 3 and 4).

the shallower traps, while the longer decays should be linked with the deeper traps.

A simplified scheme of the process described can be seen in Fig. 10. Since the PersL is only observed when using 250 nm excitation light, which, as indicated by the PLE results, corresponds to the excitation of the MMCT transition, it is reasonable to consider that the electrons must be excited from the ground level of the Bi^{3+} ($^1S_0$) state to the CBM (process labelled as 1 in Fig. 10). After which the electron should be captured by the intrinsic point defects located below the CBM (process 2). Once the excitation is removed, thermal energy from the surrounding environment stimulates the release of the trapped electron, starting the de-trapping process. The shallower the depths of the trap, the faster the detrapping process, corresponding to shorter decay times. This process will then populate the excited states of Bi^{3+}, either through the conduction band or directly from the defect levels (process 3). The electrons then return to the ground state of the emitting levels, emitting photons, and resulting in the observed UVA PersL (process 4).

4. Conclusions

Lithium yttrium germanate samples, nominally undoped and intentionally doped with 0.5 mol% of Bi, were prepared by high-temperature solid-state reaction, at 1000 °C for 40 hours. The band structure of the material was calculated using the HSE06 hybrid functional, which predicts a room temperature direct bandgap at the Γ point with a value of 6.01 eV. The calculations also confirm that the most favourable position for the Bi atoms in the LiYGeO$_4$ structure is at the Y site, in line with previous reports of other authors. Additionally, Bi located at the Y site appears to be thermodynamically easier to form when there are Li vacancies, both by replacing Y atoms and by

pushing Y atoms to Li vacancies, as it tends to be attracted to regions where these vacancies are present.

The produced samples are predominantly composed of the LiYGeO$_4$ crystalline phase, although small amounts of other phases were also present, resulting from the instability of the LiYGeO$_4$ at high temperatures, which may result in the breakdown of the desired stoichiometry due to lithium sublimation.

Elemental analysis revealed a Bi contamination in the undoped sample, introduced due to trace contaminations in the percursor powders. Furthermore, the Bi content in the intentionally doped sample differs from the nominal one, likely due to its evaporation during the synthesis process, which was highlighted in our thermal analysis. The loss of Li was confirmed by the elemental analysis, leading to the formation of Li vacancies.

The optical analysis showed very similar results for both samples. Absorption measurements indicate a steep slope band at shorter wavelengths, suggesting a direct bandgap for the material with a value of ∼6.05 eV, consistent with our hybrid DFT results. Under 250 nm and 300 nm photon excitation, the PL spectra revealed a strong UVA emission band peaked at ∼358 nm. Further assessment of the temperature-dependent PL provided valuable insights into the origin of the observed luminescence. Under 325 nm photon excitation, which corresponds to direct excitation of the Bi ion's excited levels, when the sample is cooled to cryogenic temperatures, a redshift in the peak position (regarding RT) is observed, accompanied by a broadening of the band. By analysing the behaviour of the PL spectra with rising temperature, we propose that at 17 K the emission arises from an overlap of the $^3P_0 \rightarrow {}^1S_0$ and $^3P_1^{(1)} \rightarrow {}^1S_0$ transitions, with the $^3P_0 \rightarrow {}^1S_0$ transition being the dominant one. As the temperature increases, a thermal population from the $^3P_0$ to the $^3P_1$ levels occurs, leading to increased intensity of the $^3P_1^{(1)} \rightarrow {}^1S_0$ transition, accompanied by a blueshift of the overall PL emission. At temperatures above 70 K, the $^3P_0$ component is no longer observed, and the development of another component, from 110 K, was detected, which we associate with the $^3P_1^{(2)} \rightarrow {}^1S_0$ transition. Thus, at RT, the PL emission peaked at 358 nm is attributed to the splitting of the $^3P_1$ level and is composed of the overlap of the $^3P_1^{(1)} \rightarrow {}^1S_0$ and $^3P_1^{(2)} \rightarrow {}^1S_0$ transitions.

It was also inferred that PersL is only observed by direct excitation into the MMCT state, which corresponds to the excitation of an electron from the Bi^{3+} ground state to the minimum of the host conduction band, as this phenomenon could only be activated when photons with 250 nm were used to illuminate the samples, matching the estimated energy for this transition. Longer wavelengths were not effective for the observation of this phenomenon. As so, attaining PersL requires the excitation of an electron from the $^1S_0$ ground state of the Bi^{3+} to the CBM, which is then captured by the trap levels. Afterglow measurements revealed that the intensity of the PersL could be detected for about 7 hours after turning off the excitation. The fact that 4 different decay times were needed to describe the afterglow profile is well in line with the 4 trap levels identified by thermoluminescence, suggesting that those are indeed the

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traps involved in the trapping/detrapping process that mediates the observed PersL. Thermoluminescence measurement allowed us to conclude that at least four traps were present in the analysed samples, with activation energies of 0.71–0.77 eV, 0.86–0.89 eV, 1.05–1.08 eV and 1.11–1.23 eV. The fact that the estimated energy values exhibit only a slight difference (less than 0.12 eV), suggests that similar traps are involved in the two samples, however with distinct defect concentrations, with different trapping/detrapping kinetics. Those traps were tentatively assigned to the +/0 level of Ge_{Li}, the +/0 level of Li_i and the +/− and 2+/+ levels of Y_{Li}, as previously reported by DFT calculation.

Author contributions

Data availability

Data for this article are available at Mendeley Data at **https://data.mendeley.com/datasets/zgw9v3h4nw/1**.

Conflicts of interest

The authors declare no conflict of interest.

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