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# Room temperature ferromagnetism in SnO<sub>2</sub> nanoparticles: An experimental and Density Functional study

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## Abstract

Room temperature ferromagnetism is observed in pristine SnO<sub>2</sub> nanoparticles prepared via sol-gel method. Microstructural studies with XRD and Raman reveal the coexisting of SnO and SnO<sub>2</sub> phase in vacuum annealed SnO<sub>2</sub>. Electron paramagnetic resonance (EPR) results confirm the presence of paramagnetic trapped hole in lattice oxygen atom and superoxide radical species in air and vacuum annealed SnO<sub>2</sub> respectively. It is seen that growth condition has significant influence on the magnetism of the nanoparticle. Air and vacuum annealed SnO<sub>2</sub> display clear ferromagnetic behavior at room temperature. However, the saturation magnetization reduces in vacuum annealed SnO<sub>2</sub>. We have compared the experimental results with ab initio density functional calculations and observed that the Sn vacancies are the main reason of magnetism in SnO<sub>2</sub>. The pattern of theoretical formation energy values shows that it is easier to form Sn vacancies in SnO<sub>2</sub> under O rich limit (similar to air annealing). This is in conformity with the experimentally determined higher saturation magnetization values in air annealed SnO<sub>2</sub>. Spin polarization of lattice oxygen atom by the presence of Sn vacancy gives rise to magnetism.

**Keywords:** Room temperature ferromagnetism, SnO<sub>2</sub>, DFT calculations, Oxygen vacancies.

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## 22 1. Introduction

23 In the field of spintronics it is important to design semiconductors with ferromagnetically  
24 polarized carriers at room temperature (RT).<sup>1</sup> The understanding of spin-based devices for quantum  
25 information depends upon the improvement of various nanoscale semiconductor architectures that allows  
26 perfect isolation and coherent control of the quantum confined spin states.<sup>2</sup> Room temperature  
27 ferromagnetism in transition metal doped semiconductor oxides is widely reported.<sup>2-5</sup> However, the  
28 formation of extrinsic magnetic impurity clusters stand as an obstacle in the practical applications of these  
29 materials. Hence, the quest is for RT ferromagnetism in a semiconductor material with appreciable  
30 saturation magnetization. Undoped oxide materials seem to be a better candidate for spintronics as the  
31 possibility of extrinsic magnetic phase separation due to magnetic impurities could be avoided. There are  
32 reports of ferromagnetism in undoped SnO<sub>2</sub> nanosystems.<sup>6-8</sup> Few authors<sup>7-10</sup> have claimed that the  
33 observed ferromagnetism in pristine SnO<sub>2</sub> is due to the presence of oxygen vacancies. On the other hand  
34 authors<sup>11,12</sup> have claimed on theoretical grounds that the main reason behind the observed ferromagnetism  
35 in the SnO<sub>2</sub> system is the presence of Sn<sup>4+</sup> vacancy. Espinosa et.al<sup>12</sup> confirmed that Sn<sup>4+</sup> vacancies are the  
36 main source of magnetism in SnO<sub>2</sub> thin film. Similarly, Rusydi *et al.*<sup>13</sup> and Qi *et al.*<sup>14</sup> justified with  
37 experimental and theoretical approach cationic (Ti) vacancy mediated ferromagnetism in Ta doped TiO<sub>2</sub>.  
38 Ghosh *et al.*, however, reported diamagnetism in pure SnO<sub>2</sub> nanoparticles.<sup>15</sup> Therefore, the exact  
39 origin of ferromagnetism in these nanoparticles is not clearly explained.

40 In this article we have performed both experimental and theoretical studies to establish the  
41 dependence of growth condition (oxygen rich or oxygen deficient) on the magnetism of SnO<sub>2</sub>  
42 nanoparticles. Initial characterization results of XRD and Raman show the coexisting of SnO and SnO<sub>2</sub>  
43 phases in vacuum annealed SnO<sub>2</sub>. We have used density functional (DFT) calculation using MedeA  
44 VASP (Vienna ab-initio simulation package) 5.2 for the investigation of intrinsic cause of magnetism in  
45 pristine SnO<sub>2</sub> prepared under both air and vacuum condition. Finally, we have verified on experimental

46 and theoretical ground the influence of cationic  $\text{Sn}^{4+}$  vacancy and anionic oxygen vacancy on  
47 ferromagnetism in  $\text{SnO}_2$ .

48 **2. Synthesis of  $\text{SnO}_2$  nanoparticles:** The synthesis of  $\text{SnO}_2$  followed sol-gel procedure.<sup>16</sup> All the  
49 reagents were of AR grade. The reagents used were  $\text{SnCl}_2 \cdot 2\text{H}_2\text{O}$  (98 % Merck), double distilled water,  
50 ethanol (Merck), HCl (Merck), and ammonia solution (Merck). 1g of  $\text{SnCl}_2 \cdot 2\text{H}_2\text{O}$  was mixed with 11 ml  
51 double distilled water and 5 ml ethanol. The above solution was stirred for 15 minutes and then few drops  
52 of HCl was added. Ammonia solution was added drop wise till the pH = 10 were attained. The prepared  
53 gel was divided into two parts. One part was dried at  $60^\circ\text{C}$  and annealed at  $200^\circ\text{C}$ . The other part was  
54 dried and annealed in vacuum at  $60^\circ\text{C}$  and  $200^\circ\text{C}$  respectively. The sample annealed in air was named  
55 TO200, while the sample annealed in vacuum was quoted in this manuscript as  $\text{TO}_{\text{Vac}200}$ .

56 **3. Experimental and computational details:** X-ray diffraction pattern of the nanoparticles was  
57 monitored in a Bruker AXS D8 focus X-ray diffractometer (XRD) equipped with intense  $\text{Cu K}\alpha$  radiation  
58 ( $\lambda=0.154$  nm). The scanning speed was 0.5 sec/step with an increment of 0.0057 degree. High resolution  
59 transmission electron microscope (HRTEM) images of the prepared nanoparticles were obtained with  
60 JEOL JEM 2010 transmission electron microscope operating at a voltage of 200 kV. Elemental analysis  
61 of the as prepared  $\text{SnO}_2$  nanopowder was studied with inductively coupled plasma optical emission  
62 spectroscopy (ICP-OES) in Perkin Elmer Optima 2100 DV by digesting the sample completely in small  
63 volume of conc.  $\text{HNO}_3$ . Raman spectra of the samples were obtained in Renishaw in Via Raman  
64 spectrometer with an  $\text{Ar}^+$  ion laser of 514.5 nm wavelength with spectral resolution of  $0.3 \text{ cm}^{-1}$ . X-ray  
65 photoelectron spectra (XPS) were recorded on a Kratos Axis 165 XPS spectrometer equipped with dual  
66 aluminum–magnesium anodes using  $\text{Mg K}\alpha$  radiation ( $h\nu = 1253.6$  eV). The dc magnetization  
67 measurement for the as-prepared sample was performed with a Quantum Design MPMS superconducting  
68 quantum interference device (SQUID) VSM Ever Cool system at room temperature (RT). The ZFC (zero  
69 field cooling) and FC (field cooling) temperature dependent magnetization (M-T) measurements were

70 done in the temperature range of 5-350 K. Electron paramagnetic resonance (EPR) was performed using  
71 JES - FA200 ESR Spectrometer in X band, sensitivity  $-7 \times 10^9$  spins/0.1mT and resolution - 2.35(micro) T  
72 or better .We performed density functional calculations using generalized gradient approximation (GGA)  
73 functional with Perdew-Burke-Ernzerhof (PBE) to describe the electron-electron exchange and  
74 correlation effects. The density functional theory (DFT) equations were solved via projector augmented  
75 wave (PAW) method using plane wave basis set as implemented in Vienna ab initio Simulation Package  
76 (VASP) and interfaced with MedeA technology platform. All the systems were calculated on a real space  
77 grid defined with a plane-wave energy cut off of 480 eV. A spacing of  $0.3 \text{ (\AA)}^{-1}$  was used for K point  
78 sampling in both supercell and surface calculation. All atomic positions were fully relaxed until all  
79 atomic forces were smaller than  $0.01 \text{ eV/\AA}$ . A  $2 \times 2 \times 2$  supercell was built with the unit cell of SnO<sub>2</sub> in  
80 rutile form and with a  $10 \text{ \AA}$  vacuum space between two supercells. The space group of pure SnO<sub>2</sub> was  
81 taken as  $P_{42/mmm}$  and the Supercell contained 48 atoms (Sn<sub>16</sub>O<sub>32</sub>). The effect of Sn vacancy was  
82 investigated by removing Sn atom from centre position of the Supercell. Similarly process was done for O  
83 vacancy but for Sn interstitial an extra Sn was added in the above mentioned supercell (Sn<sub>16</sub>O<sub>32</sub>). In all  
84 the cases, spin polarized structural optimization was performed such that the systems obtained its  
85 minimum energy configuration by relaxing all the atoms in all direction and calculated the magnetic  
86 moments. For the surface (110) we also considered 48 atoms and relaxed the structure. A  $10 \text{ \AA}$  vacuum  
87 slab was considered between two surfaces.

## 88 4. Results and discussion

### 89 4.1 Structural characterizations

90 X-ray diffraction patterns of the prepared SnO<sub>2</sub> nanoparticles were shown in Fig. 1. All the  
91 diffraction peaks were indexed to rutile (JCPDS-41-1445) phase. XRD graph indicated presence of SnO  
92 (101), (200) and (112) characteristic peaks along with SnO<sub>2</sub> phase in TO<sub>vac</sub>200, whereas these peaks are

93 missing in TO200. Surface oxygen vacancies are easily formed in the as prepared SnO<sub>2</sub> under vacuum  
 94 annealing. SnO<sub>2</sub> tries to compensate these oxygen vacancies by transforming itself to SnO.<sup>17</sup> Crystallite  
 95 sizes of all the prepared systems are calculated using Scherer's formula in the EVA software equipped  
 96 with the XRD instrument. The determined values are 5.44 nm for TO200 and 12.5 nm for TO<sub>Vac</sub>200. We  
 97 believe that larger crystallite in TO<sub>Vac</sub>200 is most likely due to the coexistence of SnO<sub>2</sub> and SnO  
 98 crystalline phases. We further studied the morphology, size and shape of SnO<sub>2</sub> with high resolution  
 99 transmission electron microscope (HRTEM). Fig. 2 (a-b) displays the HRTEM images of TO<sub>Vac</sub>200 and  
 100 TO200 respectively. These images show that particles are of varied sizes and shapes. Particles in  
 101 TO<sub>Vac</sub>200 are of larger size and lesser agglomerated than in TO200. The size distribution plot shows that  
 102 the particles of sizes 13 nm and 6 nm are highly populated in TO<sub>Vac</sub>200 and TO200 respectively. This  
 103 compliments the crystallite sizes obtained from XRD.

104 Raman spectroscopy is an efficient tool to explore the vibrational modes within the system. The  
 105 Raman spectra of both the prepared system is shown in Fig. 3. A normal mode can be described by the 3n  
 106 (n is the number of atoms in the primitive cell) coordinates giving the displacements of the atoms from  
 107 their equilibrium positions. Rutile SnO<sub>2</sub> having 6 atoms unit cell give out 18 possible vibrations. The  
 108 normal lattice vibration at the  $\Gamma$  point of the Brillouin zone is as follows.<sup>18,19</sup>

$$109 \quad \Gamma = \Gamma_1^+(1A_{1g}) + \Gamma_2^+(1A_{2g}) + \Gamma_3^+(1B_{1g}) + \Gamma_4^+(1B_{2g}) + \Gamma_5^-(1E_g) + \Gamma_1^-(1A_{2u}) + 2\Gamma_4^-(B_{1u}) + 4\Gamma_5^+(E_u) \quad (1)$$

110 Two modes ( $A_{2u}$  and triply degenerate  $E_u$ ) are infrared active.  $E_u$  has two acoustic and one optic mode  
 111 while  $A_{2u}$  has one optic and one acoustic mode. Four modes, viz.,  $A_{1g}$ ,  $B_{1g}$ ,  $B_{2g}$  and doubly degenerate  $E_g$   
 112 are found to be Raman active and two ( $A_{2g}$ ,  $B_{1u}$ ) modes are silent. Raman active modes for rutile SnO<sub>2</sub>  
 113 (bulk) are expected at 123, 476, 634 and 778 cm<sup>-1</sup> for  $B_{1g}$ ,  $E_g$ ,  $A_{1g}$  and  $B_{2g}$  respectively.<sup>20</sup> In our case only  
 114  $A_{1g}$  peak is prominent in both the samples. Both samples exhibit a hump at 574 cm<sup>-1</sup> which is referred as  
 115 the  $A_1$  mode and is due to surface phonon in SnO<sub>2</sub> nanocrystallites.<sup>21</sup> One small peak is observed at 210  
 116 cm<sup>-1</sup> in TO<sub>Vac</sub>200 which is otherwise absent in TO200. This peak appears owing to the presence of small

117 fraction of SnO phase alongside SnO<sub>2</sub> phase.<sup>22,23</sup> The Raman active modes of SnO<sub>2</sub> (or SnO) are due to  
118 the polarization occurring due to the displacements of Sn and oxygen in Sn and oxygen sub-lattice in  
119 SnO<sub>2</sub>. Any changes in the stoichiometry (by the presence of oxygen vacancies) in the basic SnO<sub>6</sub> unit (in  
120 SnO<sub>2</sub>) and in SnO<sub>4</sub> (in SnO) may significantly reduces the polarization of Sn-O bond and thus, might  
121 influence the position, intensity of Raman modes. We have observed blue shifting in the position of the  
122 A<sub>1g</sub> as well as surface A<sub>1</sub> phonon Raman mode in TO<sub>Vac</sub>200 as compared to TO200. The blue shifting can  
123 be associated with presence of oxygen vacancies in the TO<sub>Vac</sub>200. Presence of oxygen vacancies reduces  
124 the numbers of polarized Sn-O bond in TO<sub>Vac</sub>200, resulting in the shifting in the position of Raman peak.  
125 Similar type of oxygen defect assisted blue shifting of Raman peak is reported in vacuum annealed SnO<sub>2</sub>  
126 as compared to its air annealed counterpart by Shek *et al.*<sup>21</sup>

127 Room temperature EPR spectra TO<sub>Vac</sub>200 and TO200 are shown in Fig. 4. The *g* value is  
128 calculated using the relation

$$129 \quad g = \frac{h\nu}{\mu_B B} \quad (2)$$

130 Where *h* is the Planck's constant, *ν* is the applied microwave frequency (9.878 GHz), *μ<sub>B</sub>* is the Bohr  
131 magneton and *B* is the resonance magnetic field. TO200 contains an intense EPR signal at *g* = 2.04. This  
132 signal can be attributed to hole trapped at lattice oxygen site (O<sup>•</sup> center).<sup>24,25</sup> TO<sub>Vac</sub>200 contains a broad  
133 EPR signal at *g* ~ 2.17. This EPR signals resemble the *g*-value corresponding to superoxide radical.<sup>26,27</sup>  
134 The superoxide radicals (O<sub>2</sub><sup>•</sup>) are formed by the transfer of trapped electron present in oxygen vacancies  
135 to the surface adsorbed O<sub>2</sub> molecule in SnO<sub>2</sub>. These superoxide radicals are strongly attached to Sn<sup>4+</sup> or  
136 Sn<sup>2+</sup> species on the surface. Surfaces in TO<sub>Vac</sub>200 containing high density of oxygen vacancies with  
137 trapped electrons are active sites for easy adsorption of O<sub>2</sub> molecule. Therefore, TO200 contain EPR  
138 signal due to hole trapped at oxygen lattice site, but does not contain any signal corresponding to oxygen  
139 vacancy or superoxide species. On the other hand TO<sub>Vac</sub>200 contains EPR signal corresponding to oxygen

140 vacancies only but no signals of hole trapped oxygen. The relative number of spins participating in the  
141 broadening and intensity of EPR signal are calculated by the relation<sup>28</sup>

$$142 \quad N_s \propto I (\Delta H)^2 \quad (3)$$

143 I and ( $\Delta H$ ) are the intensity and FWHM of the intense peak calculated from the integrated spectrum of  
144 each sample. It is found that relative number of spins ( $N_s$ ) due to electron trapped in oxygen vacancy is  
145 higher than that of hole trapped at oxygen lattice site, viz.,  $N_s$  in  $\text{TO}_{\text{vac}200} > \text{TO}200$ .

146 We have carried out XPS analysis of  $\text{TO}_{\text{vac}200}$  (Fig. 5) and  $\text{TO}200$  (Fig. 6). The standard  $3d_{5/2}$   
147 spectrum for  $\text{Sn}^{4+}$  and  $\text{Sn}^{2+}$  are found to be at 486.72 and 487.4 eV respectively.<sup>29,30</sup> The separation  
148 between  $\text{Sn } 3d_{5/2}$  and  $\text{Sn } 3d_{3/2}$  should be around 8.5 eV.<sup>30</sup> In  $\text{TO}200$  the peaks found for  $\text{Sn}^{4+}$  are at 486.9  
149 and 495.4 eV while in  $\text{TO}_{\text{vac}200}$  the peaks for  $\text{Sn}^{4+}$  are obtained at 486.7 and 495.2 eV. The peaks of  $\text{Sn}^{2+}$   
150 in  $\text{TO}_{\text{vac}200}$  are found to be at 488.1 and 496.6 eV. We have not observed any  $\text{Sn}^{2+}$  peak in  $\text{TO}200$  which  
151 confirms the absence of  $\text{SnO}$  in  $\text{TO}200$ . The separation between  $3d_{5/2}$  and  $3d_{3/2}$  spectra of  $\text{Sn}^{4+}$  and  $\text{Sn}^{2+}$  is  
152 8.5 eV, although peaks are found to be shifted to higher energy in comparison to standard spectrum. The  
153 peaks of  $\text{Sn}^{2+}$  are found to be shifted to higher energy by 0.7 eV with respect to the standard sample value  
154 (given above). O 1s spectra of both the sample is deconvoluted into two peaks. The peak in lower energy  
155 can be ascribed to lattice oxygen ( $\text{O}_L$ ) while the higher energy one is due to adsorbed oxygen at the  
156 surface ( $\text{O}_H$ ).<sup>29n31</sup> The percentage concentration of  $\text{O}_L$  and  $\text{O}_H$  is found to be 59.5 and 40.5 respectively  
157 for  $\text{TO}200$  while it is 53.6 and 46.4 respectively for  $\text{TO}_{\text{vac}200}$ . The amount of adsorbed oxygen in  
158  $\text{TO}_{\text{vac}200}$  increases than in  $\text{TO}200$ . Actually, in  $\text{TO}_{\text{vac}200}$  the number of oxygen vacancies increases such  
159 that the surface becomes more reactive and hence adsorb large numbers of free oxygen molecules.

## 160 4.2 Magnetic properties study

161 Prior to the magnetic measurements ICP-OES elemental analysis for the magnetic impurity Fe,  
162 Co and Ni are carried out. The concentration of Fe, Co and Ni is found to be 0.088, 0.003 and 0.085 ppm



163 respectively. The form of these impurities can be in pure metal phase or in metal oxide phase. FeO and  $\alpha$ -  
164 Fe<sub>2</sub>O<sub>3</sub> are anti ferromagnetic whereas  $\gamma$ -Fe<sub>2</sub>O<sub>3</sub> is ferromagnetic. Fe<sub>3</sub>O<sub>4</sub> can easily oxidize to  $\gamma$ -Fe<sub>2</sub>O<sub>3</sub>.  
165 Therefore, in our case we have to consider the magnetization of Fe as metal or as  $\gamma$ -Fe<sub>2</sub>O<sub>3</sub>. The  
166 magnetization of Fe and  $\gamma$ -Fe<sub>2</sub>O<sub>3</sub> for 0.088 ppm is of 10<sup>-3</sup> order. Similarly, metallic Co and Ni are  
167 ferromagnetic, whereas Co<sub>3</sub>O<sub>4</sub> and NiO are antiferromagnetic. The maximum contribution from both the  
168 phases of Co is of the order of 10<sup>-4</sup> emu/g and from Ni is 10<sup>-3</sup> emu/g. In our case we have got  
169 magnetization of SnO<sub>2</sub> of the order of 10<sup>-1</sup> emu/g. Therefore, we conclude that the magnetization values  
170 reported in this work are not due to impurities and it is intrinsic in SnO<sub>2</sub>. Fig. 7(a) shows room  
171 temperature (300 K) hysteresis loops of TO<sub>vac</sub>200 and TO200. Even though both samples show RT  
172 ferromagnetism, the saturation magnetization (M<sub>s</sub>) appears to be higher in TO200 than in TO<sub>vac</sub>200. The  
173 values of M<sub>s</sub> and H<sub>c</sub> (coercivity) and M<sub>r</sub> (remanence) of both the samples are inserted in Table 1. M<sub>s</sub> & H<sub>c</sub>  
174 for TO200 are found to be 0.16 emu/g and 63 Oe respectively. As table I shows, H<sub>c</sub> decreases with  
175 increase in M<sub>s</sub>, which is also consistent with the equation<sup>32</sup> H<sub>c</sub> = 2K<sub>1</sub>/M<sub>s</sub>, where K<sub>1</sub> is the anisotropy  
176 constant. Arrot-Belov-Kouvel (ABK)<sup>28</sup> plot of TO<sub>vac</sub>200 and TO200 is demonstrated in Fig. 7(b). In the  
177 ABK plot, M<sup>2</sup> is plotted against H/M and the spontaneous magnetization is measured from the intercept  
178 of the linear fitting of the curve on Y-axis. ABK plot signifies the presence of ferromagnetic interaction  
179 and measures the spontaneous magnetization in the sample. The spontaneous magnetizations of TO200  
180 and TO<sub>vac</sub>200 are 0.03 and 0.02 emu/g respectively. The increase in spontaneous magnetization of TO200  
181 indicates the presence of long range ferromagnetic ordering.

182 For a better understanding of the magnetic property in the system we have also carried out zero  
183 field cooling (ZFC) and field cooling (FC) measurements of TO200 in the temperature range of 4 to 350  
184 K at an applied field of 100 Oe, as shown in Fig. 8. The bifurcation between FC and ZFC up to 350 K  
185 indicates the presence of ferromagnetic interaction in the system. Several factors might affect this  
186 bifurcation including magnetic anisotropy in the system, coercive field and the way by which FC and  
187 ZFC measurements are being carried out. Absence of characteristic sharp peak or hump in the ZFC curve

188 signifies that the observed ferromagnetism is intrinsic in nature and there is no blocking of spins. Both M-  
189 H and M-T measurements indicate presence of room temperature ferromagnetism in the system. The high  
190 value of  $M_s$  of TO200 as compared to vacuum annealed sample indicate that vacuum annealing do affect  
191 the magnetic property. It is known that annealing of SnO<sub>2</sub> nanoparticles under vacuum removes the  
192 loosely bound oxygen atoms on the surface and generates oxygen vacancies. The paramagnetic oxygen  
193 species that may contribute towards magnetism are (i) oxygen vacancies with single trapped electron<sup>8</sup> (F<sup>+</sup>)  
194 and (ii) single electron trapped oxygen molecules (O<sub>2</sub><sup>-</sup>). Although we have not found direct confirmation  
195 of the presence of F<sup>+</sup> center from EPR, presence of O<sub>2</sub><sup>-</sup> signal might indicate that F<sup>+</sup> centers are present on  
196 the surface. The single trapped electron might be easily transformed to adsorbed O<sub>2</sub> molecule and forms  
197 O<sub>2</sub><sup>-</sup> species. We could believe that vacuum annealing has increased the density of oxygen vacancies and  
198 thus, the density of paramagnetic superoxide species. Because of the presence of higher numbers of  
199 electron spins in TO<sub>vac</sub>200 one could expect high magnetization in TO<sub>vac</sub>200 as compared to TO200.  
200 Instead, the lowering of both saturation and spontaneous magnetization in TO<sub>vac</sub>200 might indicate that  
201 the oxygen vacancies (or to mention superoxide species) have negative impact on the total magnetization.

202 Therefore, it is likely that cationic Sn<sup>4+</sup> vacancy might be affecting the ferromagnetism. Earlier  
203 reports summarized that cationic vacancies could produce magnetism in different oxide systems.<sup>12-14,33,34</sup>  
204 Initially prepared SnO<sub>2</sub> nanocrystallites may contain considerable numbers of Sn<sup>4+</sup> and oxygen vacancies  
205 on the surface. On air annealing of as prepared SnO<sub>2</sub> powder at 200 °C, most of the oxygen vacancies  
206 disappeared and thus, only Sn<sup>4+</sup> vacancies are left. These Sn<sup>4+</sup> vacancies should contribute to magnetism  
207 but there is ambiguity in the literature reports on the observed magnetism in SnO<sub>2</sub>.<sup>7,9,35</sup> Our experimental  
208 results indicate that oxygen and Sn<sup>4+</sup> vacancies have respectively negative and positive contribution to the  
209 observed magnetism. In order to establish the roles of oxygen and Sn<sup>4+</sup> vacancies in ferromagnetism, we  
210 also performed theoretical calculations based on DFT. We have thus calculated electronic properties of  
211 several point defects starting from vacancies to interstitial in two different model structures of SnO<sub>2</sub>  
212 {Surface (110) and Supercell}.

### 213 4.3. Electronic structure and magnetic calculation

214 We have computed both surface and supercell structures and did a detail DFT based investigation  
215 on different kind of vacancies to arrive at a model which explains the observed magnetism in pristine  
216 SnO<sub>2</sub> nanosystem annealed in vacuum and air. We have considered SnO<sub>2</sub> (110) surface having 48 atoms  
217 and relaxed the structure to settle for minimum energy configuration. As we know the surface to volume  
218 ratio in nano regime is very high. The contribution from surface part is very prominent and its effect on  
219 magnetism has to be considered. We have chosen the surface calculation because most of the Sn atoms  
220 and oxygen atoms are not adequately compensated with bonds. The magnetism arises in the system due to  
221 2p state of O (a). Both these oxygen are single bonded and are not compensated. Fig. 9 shows the total  
222 density of states (TDOS) plot for SnO<sub>2</sub> where the partial density of states (PDOS) of O (a) split up near  
223 the Fermi level (0 eV). The TDOS of SnO<sub>2</sub> (110) shows the hybridization of Sn 5s and Sn 5p states with  
224 O 2p states at the top of the valence band where Sn 5s state is dominant over Sn 5p, which indicates the  
225 formation of free holes.<sup>36</sup> The asymmetric distribution of spin density around the Fermi level indicates the  
226 presence of magnetism in the system. If we put any further Sn<sup>4+</sup> vacancy or oxygen vacancy or a Sn  
227 interstitial, the magnetism appears to be decreased. This shows that the resultant magnetic moment of 1.2  
228  $\mu_B$  is due to the reduction of SnO<sub>2</sub> from bulk to surface and any further impurity does not increase the  
229 magnetic moment. The PDOS of oxygen proves the fact that these are the spin polarized oxygen which  
230 induces magnetism in pure layer of SnO<sub>2</sub>.

231 We have also investigated the supercell of SnO<sub>2</sub> containing a total of 48 atoms. We  
232 studied the effect of intrinsic defects on the structure and the magnetism. Pure SnO<sub>2</sub> supercell Sn<sub>16</sub>O<sub>32</sub> has  
233 no magnetism at all and this is quite expected as there are no dangling bonds of Sn or oxygen. We  
234 initially introduced oxygen vacancies, next Sn vacancies and then Sn interstitial. The defect formation  
235 energy  $E_f(D)$  is calculated using

$$236 E_f(D) = E_d - E_p + \frac{1}{2} n_O E(O_2) + n_{Sn} E(Sn) + n_O \mu_O + n_{Sn} \mu_{Sn} \dots\dots\dots(iv)$$

237 Where  $E_d$ ,  $E_p$ ,  $E(O_2)$  and  $E(Sn)$  are the total energy of the supercell with defect, pure supercell, oxygen  
238 molecule and single Sn atom.  $n_O$  and  $n_{Sn}$  are the number of O and Sn vacancy, while  $\mu_O$  and  $\mu_{Sn}$  are the  
239 chemical potential of O and Sn respectively. The values of  $\mu_O$  and  $\mu_{Sn}$  are external environment dependent  
240 i.e., it depends on the oxygen partial pressure. The two extreme conditions are known as O poor and O  
241 rich condition.<sup>17</sup> The value of formation energy of  $Sn_{15}O_{32}$  (single Sn vacancy),  $Sn_{16}O_{31}$  (single O  
242 vacancy) and  $Sn_{17}O_{32}$  (single Sn interstitial) are given in table II. We have seen that forming an oxygen  
243 vacancy under O poor condition and forming a Sn vacancy under O rich condition is energetically  
244 favorable. The formation of Sn interstitial neither in O rich nor O poor condition seems to be energetically  
245 favorable. This implies that formation of oxygen vacancy in  $TO_{Vac}200$  and the formation of Sn vacancy in  
246  $TO200$  are easier to achieve. The formation energy values shown in table II cannot be directly compared  
247 to the required experimental formation energy values in  $TO_{Vac}200$  and  $TO200$ , but still the pattern of  
248 formation of defects can be understood. The magnetic moment values determined based DFT based  
249 results are listed in table II. We checked it for different positions for oxygen vacancies but it did not result  
250 any magnetism. From table II it is clear that in an ordered  $SnO_2$ , we get maximum value of magnetization  
251 when  $Sn^{4+}$  vacancies are introduced. From our calculation we have found that the major contribution in  
252  $Sn_{15}O_{32}$  comes from the surrounding oxygen named as O (a-f), as shown in magnetization density  
253 mapping of Fig. 10. Further O (a) has same contribution as that of O (c), similarly O (b) is same as O (d)  
254 and O (e) is same as O (f). We have checked the splitting of oxygen near the Fermi level by plotting  
255 PDOS of oxygen as shown in Fig. 11. The asymmetric distribution of spin up and spin down states for  
256 oxygen (a-f) near and above the Fermi level confirms the magnetic nature of the system. We have also  
257 calculated the formation energy of oxygen and Sn vacancy on the surface (110) plane of  $SnO_2$  and  
258 inserted the values in table II. We have found that the formation energy of oxygen vacancy in O deficient  
259 condition is very low, while its formation in O rich condition is relatively harder (with high formation  
260 energy). The formation energy of Sn vacancy in O deficient condition is very high, while the  
261 corresponding value is very low in O rich condition. In the bulk calculation the formation energy of

262 oxygen vacancy and Sn vacancy in two different conditions are found to be of same pattern though values  
263 are higher in this case compared to surface (110). Our claim that Sn vacancies are the source of  
264 magnetism can be verified by the result of formation of Sn vacancies in SnO<sub>2</sub>. Both in bulk and surface  
265 Sn vacancies formation are easier to achieve in O rich condition than in O deficient condition. On the  
266 other hand oxygen vacancies instantaneously form under O deficient condition in SnO<sub>2</sub> (110) surface and  
267 at slightly high energy cost in bulk SnO<sub>2</sub>.

#### 268 4.4. Correlation between DFT and experimental results

269 From the experimental and theoretical investigations it seems that the origin of magnetism in  
270 SnO<sub>2</sub> is not due to oxygen vacancies but because of Sn vacancy or the magnetism is hole mediated. From  
271 DFT calculation it is observed that the valence band splitting and spin polarization is observed near the  
272 Fermi level in Sn<sub>15</sub>O<sub>32</sub>. This splitting is due to the formation of four holes (Sn<sup>4+</sup> vacancy). The holes are  
273 distributed on the lattice oxygen nearest to Sn vacancy and the surrounding oxygen is spin polarized.  
274 However, in TO<sub>Vac</sub>200 the number of lattice oxygen surrounding Sn vacancy will be diminished.  
275 Therefore, loss of polarized oxygen around Sn vacancy will lead to reduction in average magnetization in  
276 TO<sub>Vac</sub>200.<sup>37</sup> Although our EPR results indicated that the relative numbers of unpaired spins on the oxygen  
277 vacancies are higher in TO<sub>Vac</sub>200, these unpaired spins have no positive influence on the total  
278 magnetization. Zhang *et al.* reported that point defects with local moments may not necessarily  
279 contribute towards ferromagnetism.<sup>38</sup> In the present study anionic point defects (oxygen vacancies) have  
280 no role in ferromagnetism, whereas cationic defects (Sn vacancy) has certainly the major contribution to  
281 ferromagnetism. From our results we have found that formation of Sn vacancy is favorable in TO200 and  
282 formation of oxygen vacancy is favorable in TO<sub>Vac</sub>200. Sn vacancy polarizes the nearest lattice oxygen  
283 and induces magnetic moment in them, whereas presence of oxygen vacancies reduces the number of  
284 polarized oxygen and reduces the magnetic moment. Apart from Sn<sup>4+</sup> vacancies, the smaller size of  
285 TO200 than TO<sub>Vac</sub>200 might be another possible reason of superior magnetization of TO200 over

286 TO<sub>Vac</sub>200. The first effect (magnetization due to Sn<sup>4+</sup> vacancy) is described by the supercell approach  
287 while the second effect (reduction in size) is understood by considering surface (110) calculation. In the  
288 present study the adsorbed O<sub>2</sub> molecule has negative impact on the magnetization. The numbers of these  
289 adsorbed O<sub>2</sub> molecule are expected to be very high on the surface of TO<sub>Vac</sub>200 than in TO200. We believe  
290 that because of the presence of high density of oxygen vacancy sites on the surface of TO<sub>Vac</sub>200, the O<sub>2</sub>  
291 molecule can be easily adsorbed on the surface oxygen vacancies. However, in TO200 the numbers of  
292 surface oxygen vacancies are comparatively less and thus, the adsorption sites for O<sub>2</sub> are expected to be  
293 very few. Unlike TO200 where hole trapped in lattice oxygen atom, surrounding Sn vacancy, becomes  
294 advantageous for high magnetization, the additional surface adsorbed O<sub>2</sub> molecule with trapped electron  
295 (or O<sub>2</sub><sup>-</sup>) in TO<sub>Vac</sub>200 has no contribution to the total magnetization.

## 296 5. Conclusion

297 A detailed study on structural and magnetic property is done on the SnO<sub>2</sub> nanosystems annealed in  
298 vacuum and air at 200 °C. XRD and Raman study reveals that TO<sub>Vac</sub>200 have both SnO and SnO<sub>2</sub> phase  
299 in the system. Room temperature ferromagnetism is because of the presence of Sn vacancy. Formation of  
300 Sn vacancy releases hole which are trapped on the nearest lattice oxygen, and on application of magnetic  
301 field the spins in 2*p* orbital of oxygen are polarized with induced magnetic moment. Vacuum annealing  
302 will result in the removal of polarized lattice oxygen surrounding Sn vacancy and thus, a reduced  
303 magnetic moment is resulted. DFT spin polarized structural optimization on SnO<sub>2</sub> (110) and SnO<sub>2</sub>  
304 supercell supports the experimental results and predicts that magnetism in this system might arise due to  
305 the presence of Sn<sup>4+</sup> vacancies. Oxygen vacancies have no contribution to ferromagnetism in SnO<sub>2</sub>.  
306 Defect formation energy values shows that formation of Sn vacancies in O rich system is energetically  
307 favorable compared to other point defects. The DFT prediction seems to justify the observed higher  
308 saturation magnetization of TO200 as compared to oxygen deficient TO<sub>Vac</sub>200.

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385 **Figure captions:**386 Fig.1 XRD pattern of TO200 and TO<sub>vac</sub>200.387 Fig.2 High resolution TEM images of TO<sub>vac</sub>200 (a & b) and TO200 (c & d) respectively. Histogram in  
388 the inset of Fig. 2a and Fig. 2c shows size distribution of the nanoparticles.389 Fig. 3 Room temperature Raman Spectra of TO<sub>vac</sub>200 and TO200. The position of A<sub>1g</sub> peak is shifted  
390 towards higher energy side in TO<sub>vac</sub>200.391 Fig.4 Room temperature electron paramagnetic resonance (EPR) spectra of TO<sub>vac</sub>200 and TO200.392 Fig.5 X-ray photoelectron spectra (XPS) of TO<sub>vac</sub>200.

393 Fig.6 X-ray photoelectron spectra (XPS) of TO200.

394 Fig.7 (a) Room temperature M-H curves of TO200 and TO<sub>vac</sub>200. (b) Arrot-Belov-Kouvel (ABK) plot for  
395 the determination of spontaneous magnetization in TO200 and TO<sub>vac</sub>200 respectively.396 Fig.8 Zero field cooling (ZFC) and field cooling (FC) temperature dependent magnetization (M-T) curves  
397 of TO200. Inset of the figure shows the extended part of the curve in the temperature range of 200-350 K  
398 to display the complete separation of ZFC and FC upto 350 K.399 Fig.9 Total density of states (TDOS) plot of SnO<sub>2</sub> (110) and PDOS plot of O (a), Sn 5s and Sn 5p. Red  
400 lines indicate *p* state while black line indicated *s* state. The Fermi level is shown by blue line (at 0 eV).401 Fig. 10 (a) Magnetization density develops in SnO<sub>2</sub> surface (110). (b) 2x2x2 supercell of SnO<sub>2</sub> (right) and  
402 magnetization density develops in the system due to Sn vacancy. Yellow atom denotes Sn, red denotes  
403 Oxygen and white denotes Sn vacancy.404 Fig.11 TDOS plot of Sn<sub>15</sub>O<sub>32</sub> supercell (with one Sn vacancy) and PDOS of O (a, b, e). Red lines indicate  
405 *p* state while black line indicated *s* state whereas Fermi level is shown by blue line (at 0 eV).

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## FIGURE FILES

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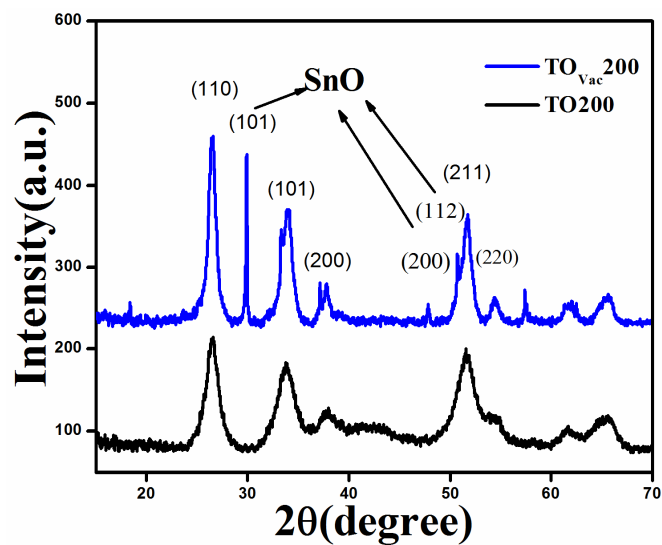
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**Figure 1**

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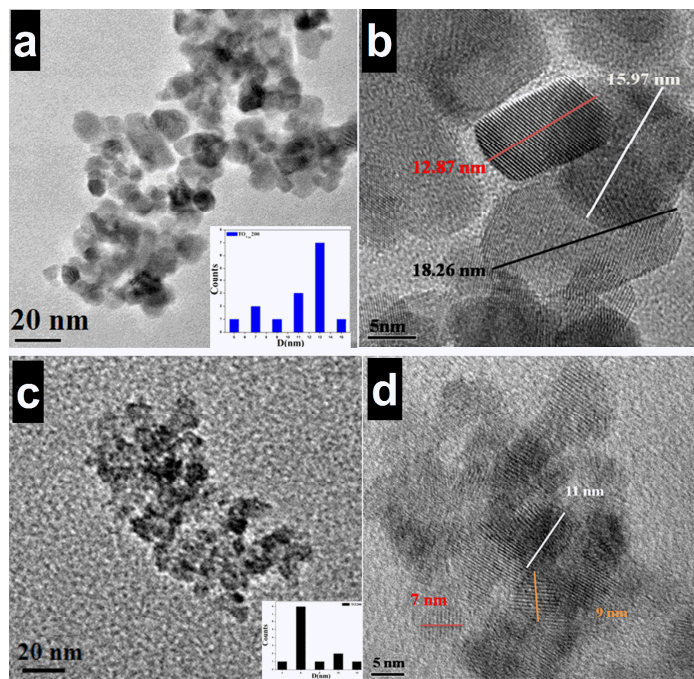
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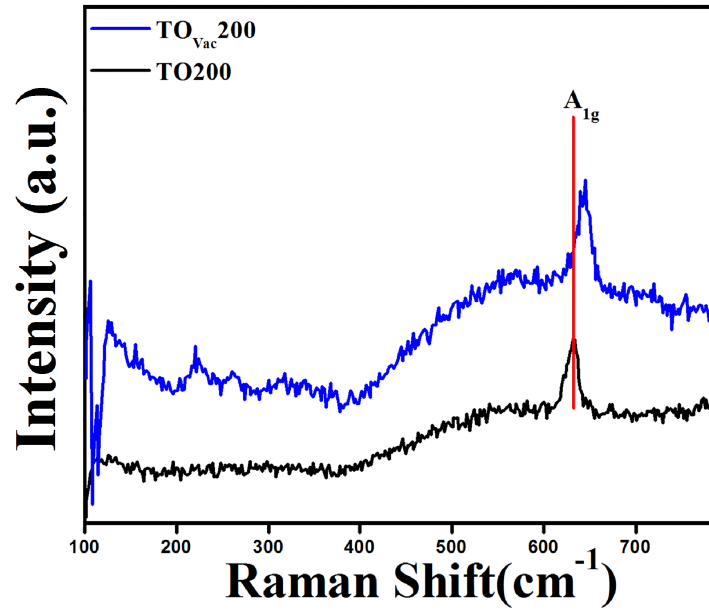
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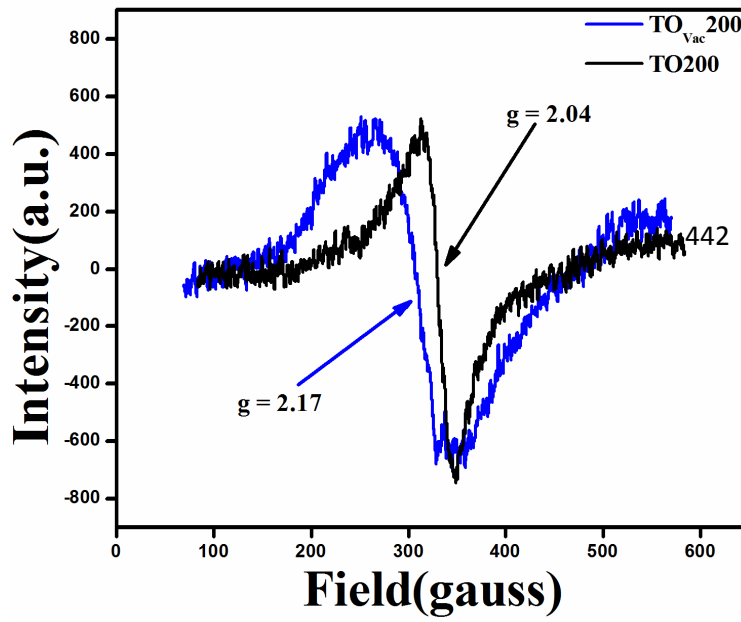
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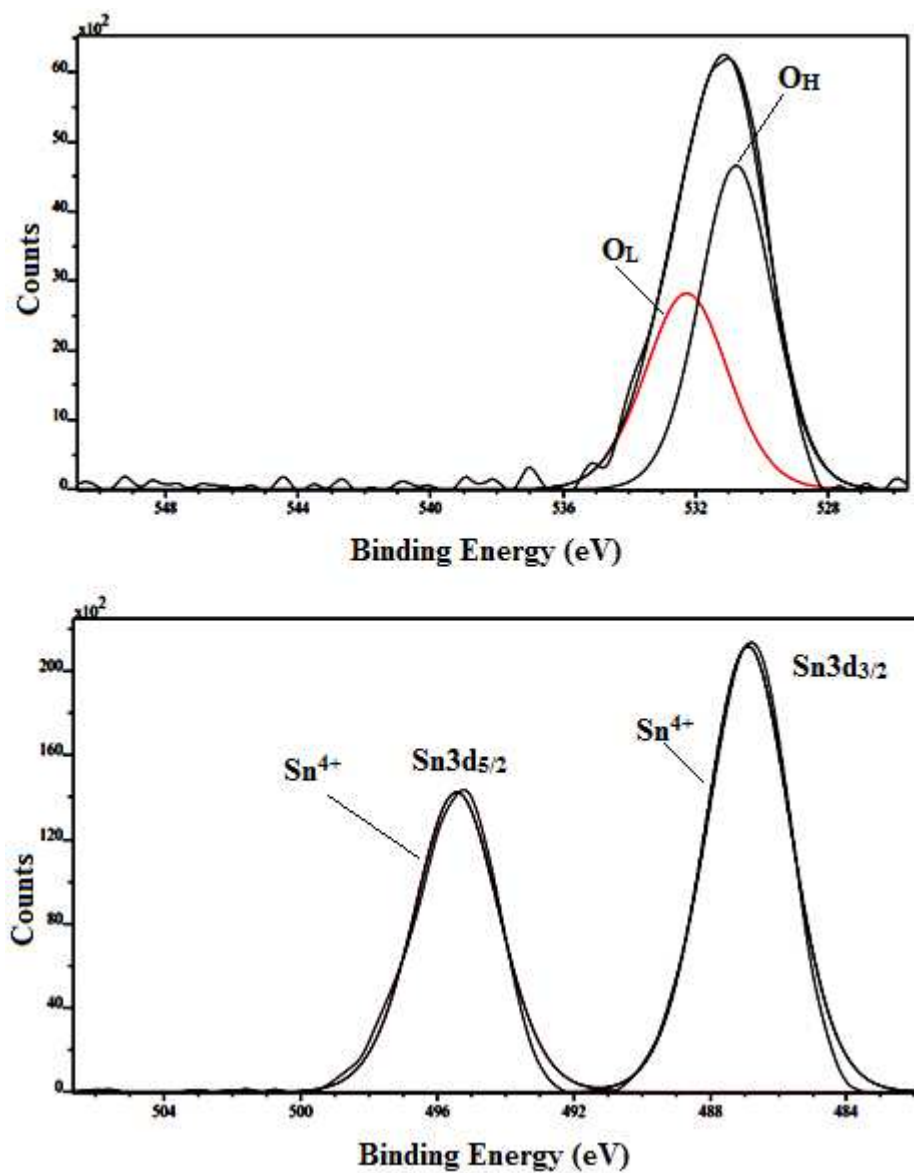
**Figure 2**



438 **Figure 3**



447 **Figure 4**



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**Figure 5**

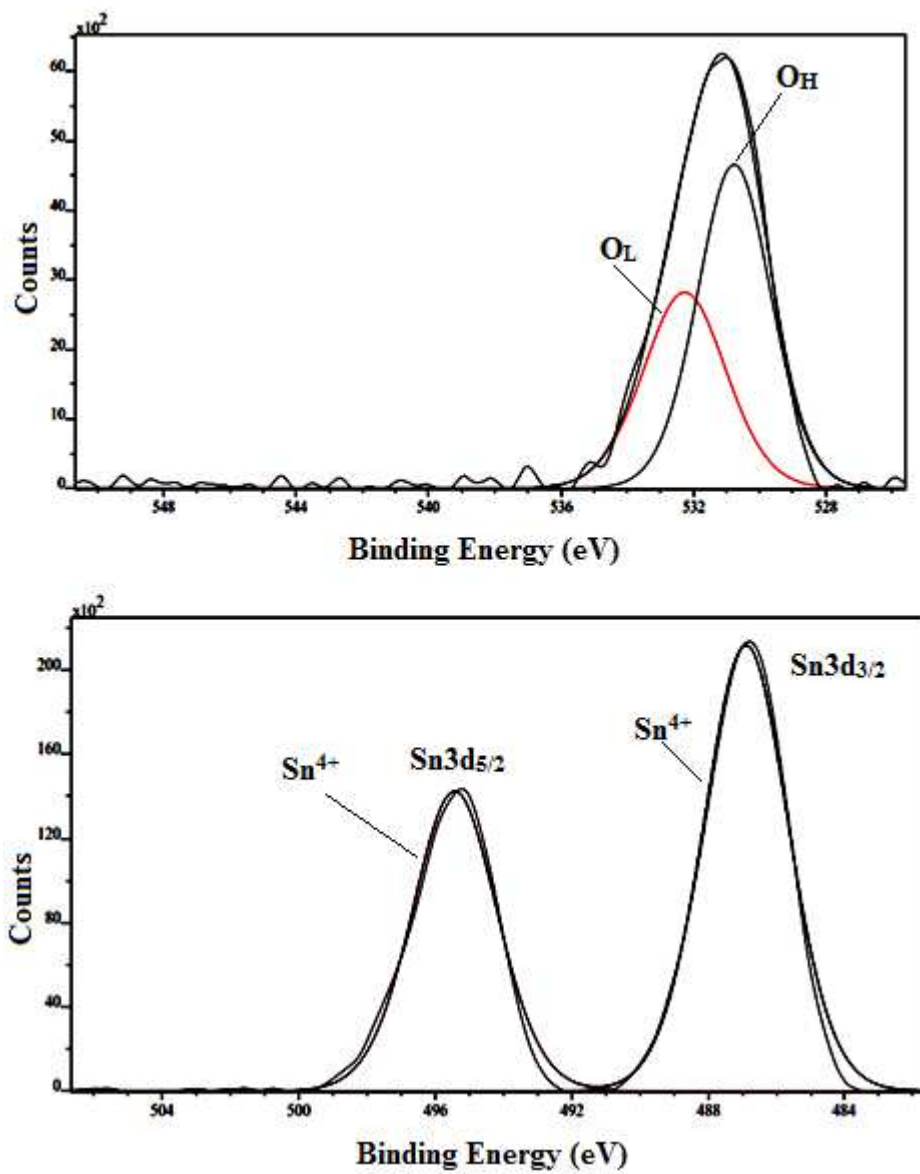
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**Figure 6**

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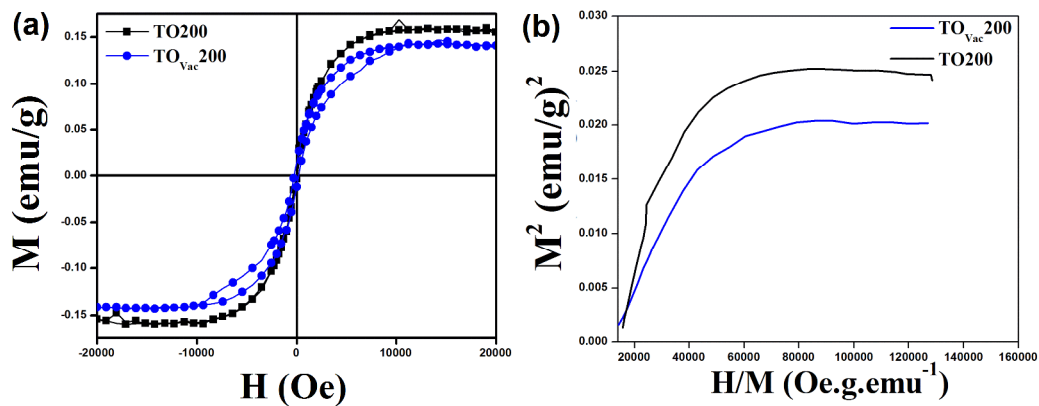
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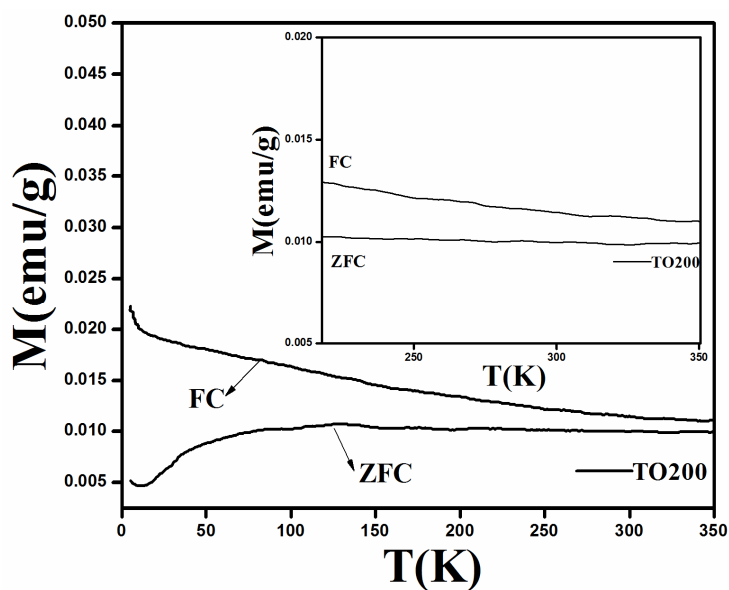
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**Figure 7**

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**Figure 8**

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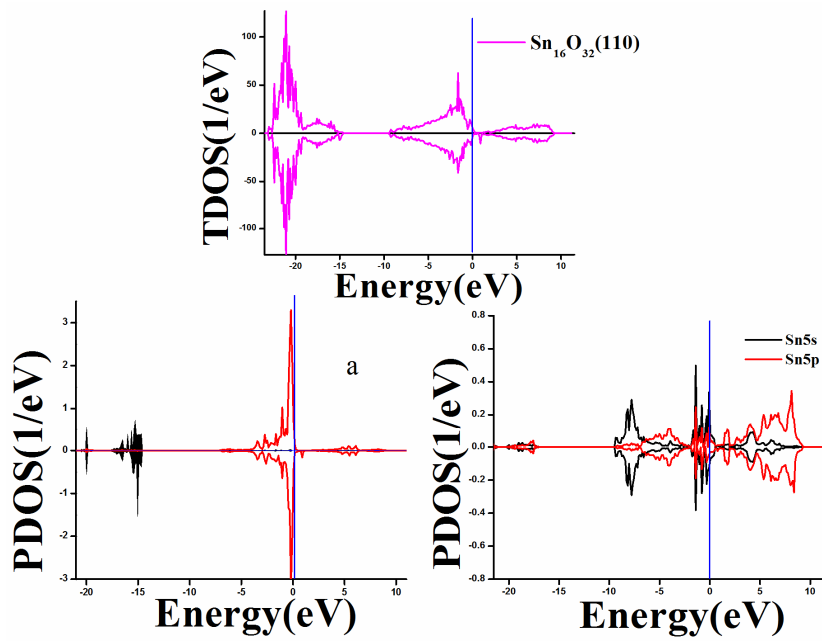
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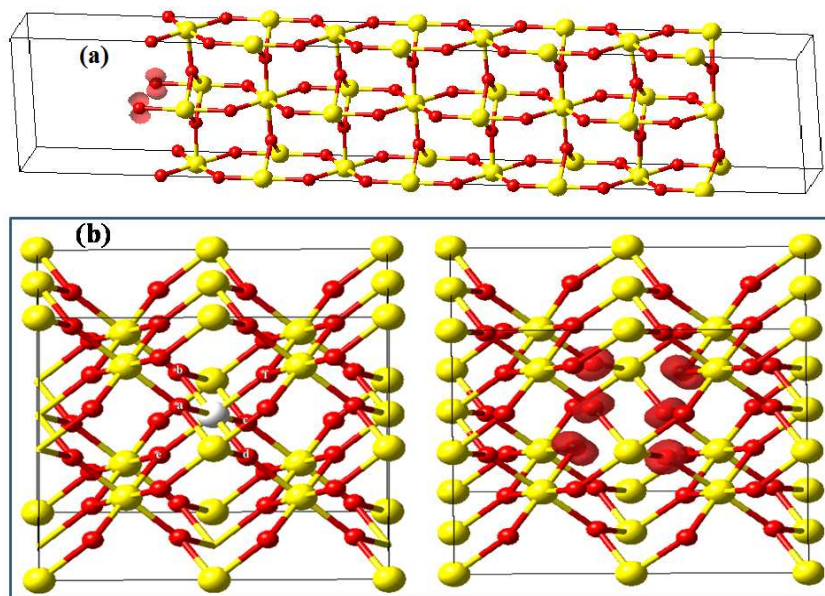
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**Figure 9**

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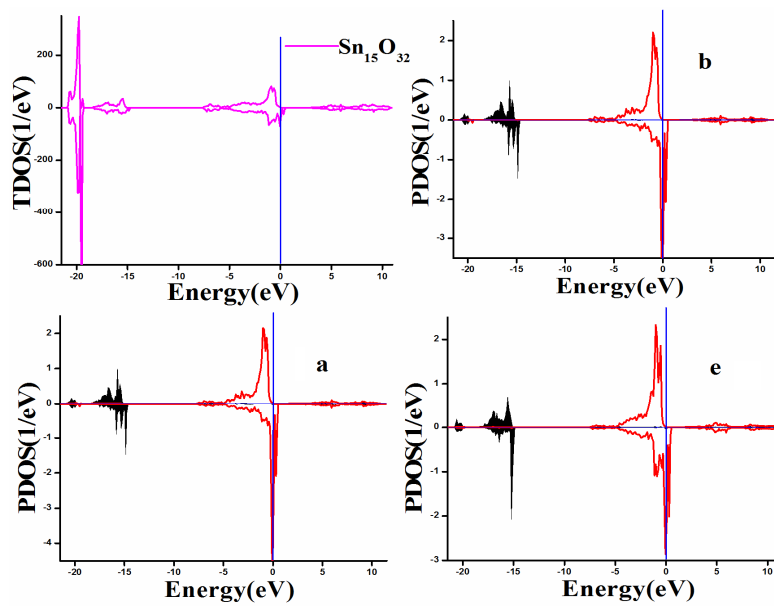
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**Figure 10**

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**Figure 11**

500 **Table I Crystallite size, saturation magnetization ( $M_s$ ), coercivity ( $H_c$ ) and remanence ( $M_r$ ) of**  
 501 **TO200 and TO<sub>Vac</sub>200**

| System                | Crystallite Size<br>(nm) from Scherer<br>formula | $M_s$<br>(emu/g) | $H_c$ (Oe) | $M_r$<br>(emu/g) |
|-----------------------|--|------------------|------------|------------------|
| TO200                 | 5.44   | 0.16             | 63         | 0.005            |
| TO <sub>Vac</sub> 200 | 12.50  | 0.14             | 200        | 0.011            |

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526 **Table II: Different systems, defect formation energy (eV) in O excess and O poor condition and**  
527 **total Magnetic moment**

| System                              | Defect formation energy $E_f(D)$ (eV) |                  | Total magnetic moment ( $\mu_B$ ) |
|-------------------------------------|---------------------------------------|------------------|-----------------------------------|
|                                     | O excess condition                    | O poor condition |                                   |
| $\text{Sn}_{16}\text{O}_{32}$       | --                                    | --               | 0.0                               |
| $\text{Sn}_{15}\text{O}_{32}$       | 1.3                                   | 7.5              | 3.9                               |
| $\text{Sn}_{16}\text{O}_{31}$       | 3.5                                   | 0.4              | 0.0                               |
| $\text{Sn}_{17}\text{O}_{32}$       | 14.3                                  | 8.2              | 0.0                               |
| $\text{SnO}_2(110)$                 | --                                    | --               | 1.2                               |
| $\text{SnO}_2(110)$ with Sn vacancy | 0.9                                   | 7.1              |                                   |
| $\text{SnO}_2(110)$ with O vacancy  | 1.6                                   | -1.4             |                                   |

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