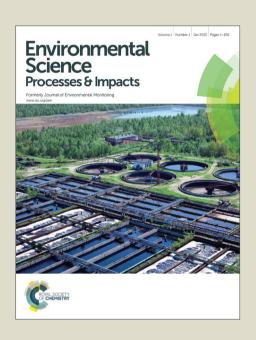
Environmental Science Processes & Impacts

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ARTICLE TYPE

Generating daily high spatial land surface temperatures by combining Aster and MODIS land surface temperature products for environmental process monitoring

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Abstract: There is a shortage of daily high spatial land surface temperature (LST) data for use in high spatial and temporal resolution environmental process monitoring. To address this shortage, this work used the Spatial and Temporal Adaptive Reflectance Fusion Model (STARFM), Enhanced Spatial and Temporal Adaptive Reflectance Fusion Model (ESTARFM), and the Spatial and Temporal Data Fusion Approach (STDFA) to estimate high spatial and temporal resolution LST by combining Advanced Spaceborne Thermal Emission and Reflection Radiometer (ASTER) LST and Moderate Resolution Imaging Spectroradiometer (MODIS) LST products. The actual ASTER LST products were used to evaluate the precision of the combined LST imageries using the correlation analysis method. This method was tested and validated in study areas located in Gansu Province, China. The results show that all the models can generate daily synthetic LST imagery with a high correlation coefficient (r) of 0.92 between the synthetic imagery and the actual ASTER LST observations. The ESTARFM has the best performance, followed by the STDFA and the STARFM. Those models had better performance in desert areas than in cropland. The STDFA had better noise immunity than the other two models.

20 Keywords: MODIS; ASTER; Land surface temperature; Multi-sensor fusion;

Environmental impact

The land surface temperature (LST) product is very important for understanding the impact of global climate change and human activities on climate change. However, there is a shortage of LST data for high spatial and temporal environmental process monitoring. To solve this problem, the Spatial and Temporal Adaptive Reflectance Fusion Model (ESTARFM), and the Spatial and Temporal Data Fusion Approach (STDFA) were tested and validated to estimate high spatial and temporal resolution land surface temperature data by combining Advanced Spaceborne Thermal Emission and Reflection Radiometer (ASTER) LST and Moderate Resolution Imaging Spectroradiometer (MODIS) LST products in study areas located in Gansu Province, China. The results show that all the models can accurately generate daily synthetic LST imagery, with a correlation coefficient higher than 0.92. The generated daily synthetic LST images enable high spatial and temporal resolution environmental process monitoring of environmental pollution, the ecological environment, and agriculture.

1. Introduction

Land surface temperature (LST) data on the earth's surface is used for the analysis and simulation of important surface energy balance parameters, and is also an important input parameter in soil moisture, drought, and crop yield estimation models. LST is also widely used in regional hydrology ecological and environmental research. Because the response of different

vegetation types or ecosystems to temperature is an important part of global change research, LST plays a large role in monitoring global climates⁶⁻⁸ and studying urban heat island effects.⁹⁻¹¹ Thus, the LST product is very important to understanding the impact of global climate change and human activities on climate change.¹²

The main method of estimating LST is an inversion approach using thermal infrared remote sensing images. ¹³ Currently, global LST products can be estimated from Advanced Spaceborne

Thermal Emission and Reflection Radiometer (ASTER) and Moderate Resolution Imaging Spectroradiometer (MODIS) data.14-15 The thermal infrared data from Landsat and the Huan Jing Constellation satellite (HJ) can also be used to estimate 5 LST. 16-17 However, due to limitations in satellite data acquisition and the influence of clouds, there is a lack of high spatial and temporal resolution LST data. MODIS LST products (MOD11A1) have a high temporal resolution and have been widely used in global or large-scale area monitoring. 18 However, these products 10 are not suitable for high precision monitoring in small areas due to their lower spatial resolution. The LST products of ASTER, Landsat, and HJ have a high spatial resolution and have been widely used in small study areas, for example, to examine urban heat islands. 19 However, these products are not suitable for time 15 series monitoring due to the long review cycle of the satellites. Therefore, there is a pressing need to combine those two types of data to achieve both high spatial and high temporal resolutions.

 Recently, several data fusion approaches have been proposed to blend high spatial resolution data and high temporal resolution data. These approaches generate synthetic high spatial resolution imagery with a high temporal resolution. Gao *et al.*²⁰ introduced the Spatial and Temporal Adaptive Reflectance Fusion Model (STARFM) for blending MODIS and Landsat imagery. Several studies applied this model and demonstrated the use of STARFM within a mainly coniferous area, as well as for urban environmental variables extraction, vegetated dry land ecosystem monitoring, public health studies, and the generation of daily land surface temperatures. ²¹⁻²⁵ Zhu *et al.*²⁶ enhanced the STARFM for complex heterogeneous regions. Other researchers proposed

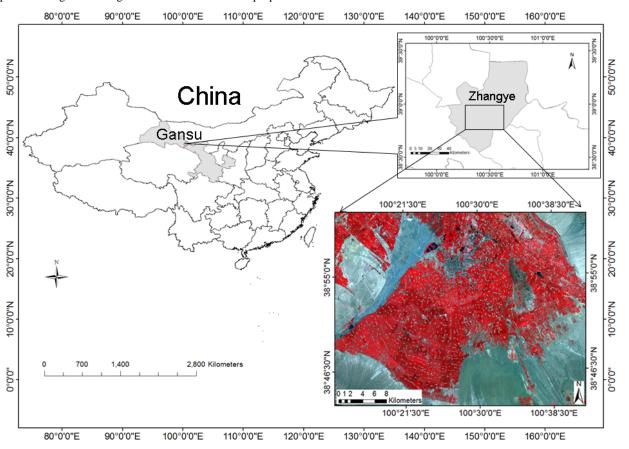
methods that fuse Landsat and MODIS data to generate high temporal resolution synthetic Landsat data based on a linear mixed model. ²⁷⁻²⁹ Wu *et al.* ³⁰ introduced a Spatial and Temporal Data Fusion Approach (STDFA) based on linear mixing theory and applied it to the estimation of a high-resolution Leaf Area ³⁵ Index³¹ and crop mapping. ³² In these cases, the methods were originally proposed fusing Landsat and MODIS reflectance images. Thus, there is a need to test the ability of these methods with other sensors and products.

The present work seeks to apply these methods in the fusion of ASTER and MODIS LST products. The objectives of this study are to (1) compare the suitability of applying STARFM, ESTARFM, and STDFA in the fusion ASTER and MODIS LST products; and (2) evaluate the applicability of these three models for different land use types.

45 2. Study area and data preparations

2.1. Study area

Zhangye country, Gansu province, China was selected as the study area for this research. The latitude and longitude range for the region are from 38° 02′ 32.23″ to 38° 09′ 24.17″ N, and from 100° 00′ 29.63″ to 100° 02′ 45.89″ E, respectively. Zhangye has a continental dry climate with an annual average temperature of 6 °C. The coldest month is January, while the hottest month is July. The main land forms of Zhangye are plains, cities, and deserts. The plains in the region are mainly devoted to ⁵⁵ agriculture.



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2.2. Data preparation

datasets (MOD11A1) were used in this study (Table 1).

Eleven ASTER LST data sets and twenty-nine MODIS LST

5 Table 1 Main characteristics of ASTER and MODIS images used in the research

ASTER	Usage		MOD11A1	ASTER	Usage		MOD11A1	ASTER	Usage		MOD11A1
06/15/2012	Reference		06/15/2012		-		07/31/2012				08/28/2012
			06/18/2012	08/02/2012	Reference Validation	and	08/02/2012				08/29/2012
			06/19/2012				08/04/2012	09/03/2012	Reference Validation	and	09/03/2012
			06/20/2012	08/11/2012	Reference Validation	and	08/11/2012	09/12/2012	Reference Validation	and	09/12/2012
			06/21/2012				08/14/2012				09/15/2012
06/24/2012	Reference Validation	and	06/24/2012				08/15/2012				09/17/2012
			07/04/2012	08/18/2012	Reference Validation	and	08/18/2012	09/19/2012	Reference Validation	and	09/19/2012
07/10/2012	Reference Validation	and	07/10/2012				08/21/2012				09/21/2012
			07/14/2012				08/22/2012				09/26/2012
			07/19/2012	08/27/2012	Reference Validation	and		09/28/2012	Reference		09/28/2012

2.2.1. ASTER data

ASTER is one of five remote sensory devices on board the Terra satellite. ASTER provides earth images every 16days in 14 different bands in a spectral range from visible to thermal infrared light. Its spatial resolution ranges between 15 to 90 meters. ASTER data are widely used to create detailed maps of land surface temperatures, emissivity, reflectance, and elevation.

Eleven ASTER LST datasets, provided by the Cold and Arid 15 Regions Science Data Center at Lanzhou, were used in this study (Table 1). The data were acquired under clear sky conditions and were estimated from ASTER L1B thermal infrared data using the temperature emissivity separation (TES) inversion algorithm. The thermal infrared data were atmospherically corrected using the 20 water vapour scaling (WVS) method in MODTRAN with MODIS atmosphere profile data products (MOD07) acquired on the same day as the ASTER data. The ASTER L1B thermal infrared data were then georeferenced using a second-order polynomial warping approach based on the selection of an 25 appropriate number of ground control points (GCPs). This procedure used ASTER L3 data with the nearest neighbour resampling method and a position error within 0.7 ASTER pixels. The accuracy of this data was evaluated using the ground measured LST data. The results showed that the average 30 deviation of this product was less than 0.5K with an RMSE less than 2K.33-34

2.2.2. MODIS LST data

MODIS is a key instrument aboard the Terra and Aqua satellites. MODIS provides earth images every 1 to 2 days in 36 spectral bands, ranging from visible to thermal infrared light. Its spatial resolution ranges from 250 to 1000 meters.

Twenty-nine MODIS LST data sets (MOD11A1, 1 km,

collection5) obtained under clear sky conditions were used in this study (Table 1). These MODIS images were re-projected from the native sinusoidal projection to a UTM-WGS84 reference system, and were resized to the selected study areas using the MODIS reprojection tool (MRT). The MODIS data were then georeferenced by a second-order polynomial warping approach based on the selection of an appropriate number of GCPs on 1000m ASTER images, using a nearest neighbour resampling method. The 1000m ASTER images were generated from georeferenced ASTER images with the pixel aggregate resampling method, and the position error was within 0.78 ASTER pixels. The MODIS LST data included some default data. Median filtering methods were used to remove these default values.

3. Approach

3.1 Model introduction

3.1.1 STARFM

55 In STARFM model, the relationship between the fine resolution image and the coarse resolution image is described as:

$$r(x, y, t_k) = R(x, y, t_k) + \xi(t_k)$$
(1)

where $r(x, y, t_k)$ is the fine resolution reflectance of target pixels (x,y) at time t_k ; $R(x, y, t_k)$ is the coarse resolution reflectance; $\xi(t_k)$ is the sensor difference. Suppose the ground coverage type and system errors does not change over prediction date, the STARFM model predicts synthetic high spatial imageries from low spatial imageries as follows: 20

$$r(x, y, t_k) = R(x, y, t_k) + r(x, y, t_0) - R(x, y, t_0)$$
 (2)

 By introducing additional information from the neighbouring pixels to reduce the influences of land cover change, surface heterogeneity, and solar geometry bidirectional reflectance 5 distribution function (BRDF) changes, a weighted STARFM model can be determined as follows:

$$r(\mathbf{x}, \mathbf{y}, t_k) = \sum_{i=1}^{w} \sum_{j=1}^{w} \sum_{k=1}^{n} W_{ijk} \left(R(\mathbf{x}_i, \mathbf{y}_j, t_k) - R(\mathbf{x}_i, \mathbf{y}_j, t_0) + r(\mathbf{x}_i, \mathbf{y}_j, t_0) \right)$$
(3)

Where k is the number of pixels (x_i,y_j) in the window w; w is the search window size; W_{ijk} is the weight determined by the spectral difference S_{ijk} and temporal difference T_{ijk} between the fine and low resolution data, and D_{ijk} is the distance between the target pixel and the candidate pixel. Those parameters are calculated as follows:

$$W_{ijk} = S_{ijk} \times T_{ijk} \times d_{ijk} \tag{4}$$

$$S_{ijk} = \left| R(\mathbf{x}_{i}, \mathbf{y}_{j}, t_{k}) - r(\mathbf{x}_{i}, \mathbf{y}_{j}, \mathbf{t}_{k}) \right| \tag{5}$$

$$T_{ijk} = |R(x_{i}, y_{j}, t_{k}) - R(x_{i}, y_{j}, t_{0})|$$
 (6)

$$D_{iik} = 1 + d_{iik} / A \tag{7}$$

$$d_{ijk} = \sqrt{(x_i - x)^2 + (y_i - y)^2}$$
 (8)

²⁰ where A is the constant used to determine whether the spectral similarity weight or the distance weight is more important.

3.1.2 ESTARFM

The ESTARFM is based on the assumption that the change of reflectance of each endmember is linear over time. Thus the ratio v_k of the change of reflectance for kth endmember to the change of reflectance for a coarse pixel can be described as:

$$\frac{(r(t_k) - r(t_0))}{(R(t_k) - R(t_0))} = v_k \tag{9}$$

Similar to the STARFM, a weighted ESTARFM model considers information from neighbouring pixels as follows:²⁶

$$r(x, y, t_k) = r(x_i, y_j, t_0) + \sum_{i=1}^{w} \sum_{j=1}^{w} \sum_{k=1}^{n} W_{ijk} \times v_{ijk} \left(R(x_i, y_j, t_k) - R(x_i, y_j, t_0) \right)$$
(10)

The weight W_{ijk} of ESTARFM is calculated as follows:

$$W_{i} = (1/D_{i}) / \sum_{i=1}^{n} (1/D_{i})$$
(11)

$$D_i = (1 + S_i) \times d_i \tag{12}$$

$$d_i = 1 + \sqrt{(x - x_i)^2 + (y - y_i)^2} / (w/2)$$
 (13)

$$S_{i} = \frac{E[(r_{i} - E(r_{i})) \times (R_{i} - E(R_{i}))]}{\sqrt{D(r_{i})} \times \sqrt{D(R_{i})}}$$
(14)

3.1.3 STDFA

The STDFA is based on the linear mixing theory. According to linear mixing theory, the reflectance of a coarse-resolution spatial ⁴⁰ pixel measured by a sensor is its composite value. The response of each coarse spatial resolution pixel is assumed to be a linear combination of the responses of each land cover class contributing to the mixture which was expressed as: ³⁵

$$R(x, y, t) = \sum_{c=0}^{K} f_c(x, y, c) \times \bar{r}(c, t) + \xi(x, y, t)$$
 (15)

45 Constrained:
$$\sum_{c=0}^{k} f_c(x, y, c) < 1$$
; and $f_c(x, y, c) \ge 0$ for all

where $f_c(x, y, c)$ is the fractional cover of class c in coarse pixel (x,y) at time t, which is usually assumed not to change over time, $\bar{r}(c,t)$ is the mean reflectance of fine resolution homogeneous pixels belonging to land cover class c and $\xi(x,y,t)$ is the residual $\xi(x,y,t)$ error term. Using the ordinary least squares technique, time series mean reflectance values can be obtained by solving the linear system of Equation (15). Then based on the assumption that the temporal variation properties of each fine resolution pixels in the same class are constant, the STDFA model predicts synthetic $\xi(x,y,t)$ high spatial imagery as follows: $\xi(x,y,t)$

$$r(x, y, t_k) = r(x, y, t_0) + r(c, t_k) - r(c, t_0)$$
 (16)

where $\overset{-}{r}(c,t_k)$ and $\overset{-}{r}(c,t_0)$ are the mean reflectance of land cover c

3.2 Application of spatial and temporal fusion models

60 For the STARFM, ESTARFM, and STDFA methods, two types of data must be input: the reference images and the time series MODIS LST products. These latter must contain MODIS LST data acquired on the same day as the reference image, and at least

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one MODIS LST data set acquired on the same day as the ASTER LST image that we want to predict. The selection of reference images has an important influence on the fusion accuracy.³⁶ Usually, the ASTER LST data that are from a date 5 near the predicted date are used as the reference images. However, since the temporal change in LST over time may not be constant, similar dates do not necessarily indicate similarity in LST. Therefore, in this work, the ASTER LST data used for the reference images were selected based on the day with the highest 10 correlation of MODIS data between that day and the day being predicted. Two-day ASTER and MODIS LST images are also needed in the ESTARFM. The early day ASTER and MODIS LST images were acquired on the same day as the reference ASTERLST image. The later day ASTER and MODIS LST 15 images were acquired on a date later than the day being predicted. Two days of ASTER LST images (one day for the reference ASTERLST image, the second after the prediction date) were also needed for classification in the STDFA. For the STARFM and ESTARFM, the MODIS LST images must be resampled to 20 the same resolution of the ASTER LST data by the nearest neighbour resampling method.

3.3 Evolution of spatial and temporal fusion models

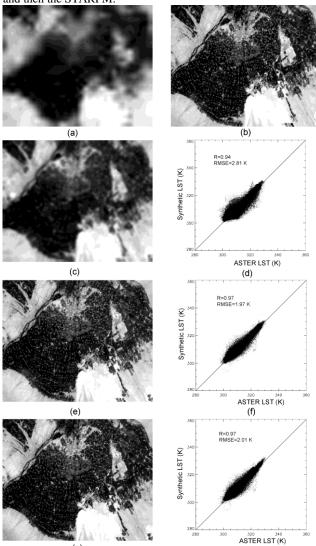
Since the objective of the STARFM, ESTARFM, and STDFA methods was to generate synthetic ASTER LST data, nine real 25 ASTER LST data sets were used to validate the algorithm. Obviously, the closer the synthetic ASTER LST imagery is to the actual imagery, the higher the precision. The algorithm was validated using two methods. First, visual interpretation was used to qualitatively evaluate the difference between the synthetic and 30 real ASTER LST imagery. If the synthetic and real ASTER LST imagery can be distinguished visually, the accuracy of the method is not very high. Second, a correlation analysis was used to quantitatively evaluate the similarity between the actual observations and the synthetic imagery. Several indicators, such 35 as the coefficient (r), variance, mean absolute difference(MAD), bias, and root mean square error(RMSE), were used to represent the precision of this model. Higher r and lower variance, MAD, bias, and RMSE indicate a higher accuracy.

4. Results

40 4.1. Overall accuracy comparison

Using the STARFM, ESTARFM, and STDFA, 29 synthetic ASTER LST images were generated using each method. Nine actual ASTER LST images were used to evaluate the accuracy of those three methods. Figure 2 shows the actual MODIS LST data, 45 the actual ASTER LST data, and three synthetic ASTER LST images generated by those three methods on 10 July 2012. Through visual interpretation, we found that the synthetic ASTER LST data generated by the ESTARFM and STDFA were very similar to the actual ASTER LST data, while the resolution of the synthetic ASTER LST data generated by the STARFM was some what lower the actual ASTER LST data. However, all

synthetic ASTER LST images had a higher resolution than the MODIS LST product. Table 2 shows the result of the correlation analysis. From Table 2, it is seen that all methods generate ASTER LST images very similar to the actual ones, with coefficient (r) higher than 0.92 and RMSE lower than 3.4k. The ESTARFM had the best performance, followed by the STDFA and then the STARFM.



LST data generated by the STARFM, ESTARFM, and STDFA acquired on 10 July 2012: (a) MOD11A1 LST data; (b) ASTER LST data; (c) synthetic LST data generated by the STARFM; (d) scatter plot between the ASTER LST and synthetic LST generated by the STARFM; (e) synthetic LST data generated by the ESTARFM; (f) scatter plot between the ASTER LST and synthetic LST generated by the ESTARFM; (g) synthetic LST data generated by the STDFA; (h) scatter plot between the

ASTER LST and synthetic LST generated by the STDFA.

Fig.2A comparison between MODIS LST, ASTER LST, and synthetic

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Table 2Comparison between actual LST and synthetic LST images

D			STARF	M				ESTARI	FM				STDF	A	
Day	R	Var	MAD	RMSE	bias	R	Var	MAD	RMSE	bias	R	Var	MAD	RMSE	bias
06/24/2012	0.94	8.92	2.26	2.99	0.09	0.96	5.87	1.79	2.43	-0.09	0.96	6.42	1.90	2.54	-0.09
07/10/2012	0.94	7.79	2.06	2.81	-0.29	0.97	3.73	1.38	1.97	-0.38	0.97	3.82	1.40	2.01	-0.48
08/02/2012	0.95	6.57	1.85	2.60	0.45	0.97	3.73	1.38	1.97	0.38	0.97	3.62	1.37	1.96	0.48
08/11/2012	0.94	6.32	1.89	2.59	-0.64	0.97	3.48	1.36	2.01	-0.75	0.95	5.06	1.69	2.35	-0.66
08/18/2012	0.92	6.44	1.98	2.71	-0.95	0.95	7.54	2.24	2.93	-1.03	0.94	6.62	2.14	2.81	-1.13
08/28/2012	0.92	11.15	2.56	3.38	-0.49	0.93	9.36	2.29	3.12	-0.58	0.94	7.85	2.19	2.85	-0.53
09/03/2012	0.93	6.74	1.94	2.67	0.63	0.92	9.36	2.29	3.12	0.58	0.93	7.50	2.15	2.79	0.54
09/12/2012	0.93	6.99	1.94	2.67	-0.37	0.96	4.18	1.41	2.09	-0.46	0.95	5.36	1.58	2.38	-0.55
09/19/2012	0.94	5.26	1.68	2.35	0.52	0.96	4.18	1.41	2.10	0.46	0.94	5.25	1.58	2.35	0.54

4.2 Accuracy comparison of each land cover types

⁵ To demonstrate the performance of the STARFM, ESTARFM, and STDFA for each land cover type, an accuracy comparison for each land cover was conducted on those nine days. Table 3 showed the correlation analysis results for those three methods with each land cover type. Figure 3 shows scatter plots with the

actual and synthetic ASTER LST images in each class, which fit well to the 1:1 lines. From Table 3 and Fig. 3, we find that the ESTARFM and STDFA have the same precision evaluated using all nine days of images in the desert class. The ESTARFM had slightly better performance than the STDFA in the crops and is cities classes. Further more, the STARFM had the worst performance in each land cover type.

Table 3Comparison between actual LST and synthetic LST imageries of different land cover types

	A	All nine days			Crops			Cities			Desert	
	STARF M	ESTARF M	STDF A	STARF M	ESTARF M	STDF A	STARF M	ESTARF M	STDF A	STARF M	ESTARF M	STDF A
R	0.73	0.96	0.96	0.15	0.62	0.55	0.35	0.80	0.79	0.75	0.96	0.96
Var	50.00	6.01	6.05	65.08	5.33	6.00	76.88	10.10	10.83	23.90	3.33	3.61
MAD	2.20	1.79	1.84	2.15	1.79	1.94	3.21	2.45	2.55	1.74	1.30	1.37
RMS E	7.07	2.46	2.47	8.13	2.31	2.45	8.80	3.22	3.34	4.94	1.83	1.90
bias	0.01	-0.21	-0.21	-1.00	0.00	0.11	0.77	-0.51	-0.58	0.70	0.12	0.06

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5. Discussion

- 5 The correlation analysis between the actual observations and the synthetic ASTER LST images showed high correlation with r higher than 0.92. This demonstrated that spatial and temporal data fusion methods, such as the STARFM, ESTARFM, and STDFA, can be used to combine ASTER LST and MODIS LST 10 data to generate daily synthetic ASTER LST data. However, issues remain that should be addressed in the application of these methods.
- (1) Influence of noise. ASTER and MODIS image noise will reduce the accuracy of the three models. Table 4 shows the 15 correlation analysis results between the actual observations and the synthetic images generated without removing the default data from the MODIS LST data. Comparing Table2 and Table4, we see that the STDFA had the smallest decrease in correlation coefficient r and had the highest tolerance for noise, while the 20 ESTARFM had the greatest decrease in correlation coefficient r and had the lowest tolerance for noise. This is because the STARFM and ESTARFM use a 3×3 window in their prediction,

while the STDFA uses a 40×40 window. Noise values appearing in the window would therefore have a greater influence on the FSTARFM

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Table 4 Comparison of ASTER LST and synthetic LST data on 10 July 5 2012 generated by the STARFM, ESTARFM, and STDFA. MODIS LST data is input with noise.

	STARFM	ESTARFM	STDFA
R	0.92	0.77	0.96
Var	10.35	36.90	6.68
MAD	2.17	2.57	2.13
RMSE	3.22	6.09	2.92
bias	-0.20	0.45	1.36

- (2) Influence of spatial variability. The spatial and temporal data fusion methods assume that the temporal effects are constant over time. This may be true in classes like deserts and cities, but may be violated in classes like crops. The temporal changes in crops are easily affected by farm management practices such as irrigation, so there would be reduced accuracy for crops. Different methods were used to handle the spatial variability of reflectance. The STARFM and ESTARFM use additional information from a 3 × 3 window to reduce the influence of spatial variability, while the STDFA uses a 40 × 40 window. Thus, spatial variability still exists in the 40×40 window of the STDFA, which results in its lower accuracy compared to the 20 ESTARFM. The ESTARFM is more suitable for complex surface areas.
- (3) The spatial and temporal data fusion methods use optical images that are easy affected by cloudy weather. For example, during the 106 days from 15 June 2012 to 28 September2012, 25 there have been only 29 MODIS images without clouds. The effective data rate is only 23%. Therefore, the development of an optical and radar data fusion algorithm is an important direction for multi-source remote sensing data. 37-42

6. Conclusions

- 30 The STARFM, ESTARFM, and STDFA methods were compared and validated in the generation of daily ASTER LST data in Zhangye Country, Gansu province, China. The results showed the following:
- (1) All three methods can generate synthetic ASTER LST images very similar to the actual ASTER LST imageries, with r values higher than 0.92 and RMSE values lower than 3.4k. The ESTARFM had the best performance, followed by the STDFA and then the STARFM.
- (2) All three methods had their best performance with desert ⁴⁰ land types. The next best performance was for cities, and the worst performance was for crops. The ESTARFM and STDFA have the same precision in the desert class, while the ESTARFM had slightly better performance than the STDFA in the crops and cities classes. The STARFM method had the worst performance ⁴⁵ for each land cover type.
- (3) ASTER and MODIS image noise will reduce the accuracy of these three models. The STDFA had the best tolerance to noise, while the ESTARFM had the worst tolerance to noise.

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65 Notes and references

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