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Long-term ambient air-stable cubic CsPbBr₃ perovskite quantum dots using molecular bromine⁺

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We report unprecedented phase stability of cubic CsPbBr₃ quantum dots in ambient air obtained by using Br₂ as halide precursor. Mechanistic investigation reveals the decisive role of temperature-controlled *in situ* generated, oleylammonium halide species from molecular halogen and amine for the long term stability and emission tunability of CsPbX₃ (X = Br, I) nanocrystals.

High photoluminescence quantum yield (PL QY), narrow emission linewidth, tunable band gap, large diffusion lengths and low exciton binding energies are some of the key attributes of all-inorganic caesium lead halide perovskite nanocrystals (LHP NCs) *i.e.*, CsPbX₃, X = I, Br, Cl.¹⁻⁶ This novel class of NCs has been shown to be highly "defect tolerant", i.e. defect states are either shallow or localized in the valence or the conduction band.^{1,2} Unlike conventional semiconductor NCs, the rigorous passivation of their surface via formation of core/shell structures or other methods is not required to achieve high QY. These LHP NCs are promising building blocks for light emitting diode,^{3,4} solar cell,^{5,6} laser,⁷ photocatalysis⁸ and detector.⁹ Despite the recent surge of studies on CsPbX₃ perovskite NCs, a persisting drawback is their poor phase stability in ambient air. For example, cubic (α) "black" phase CsPbI₃ ($E_g = 1.73$ eV) perovskite NCs undergo rapid phase transformation to nonluminescent orthorhombic (δ) "yellow" phase in ambient condition (Fig. S1[†]) leading to undesired changes of the band gap, optical and electrical properties.^{6,10,11} Similarly, cubic (α) CsPbBr₃ ($E_g = 2.25$ eV) is unstable at ambient condition.¹² For successful integration of these materials into devices, the issue

^bDepartment of Biomedical Sciences, University of Hull, Hull, HU6 7RX, UK ^cDepartment of Physics, University of Hull, Hull, HU6 7RX, UK of long-term phase stability must thus be addressed.6,13 Most of the reported strategies involve the use of additives such as halide salt,12 phosphinic acid,14 ammonium halide,11 2,2'-iminodibenzoic acid15 sulphides and metal ions16 and polymers5 or via special post-synthetic purification step.6,10 Herein, we report the first synthesis of highly stable, cubic α-CsPbBr₃ perovskite NCs using Br₂ as an independent halide precursor. In a typical synthesis, lead acetate is dissolved in 1-octadecene in the presence of oleyl amine (OAm) and oleic acid (OA). To this solution, Br₂ (warning: handle the liquid Br₂ in fume hood, Br₂ vapors are toxic) and caesium oleate solutions (both dissolved in ODE) are sequentially added. Phase stability and emission colour tunability are achieved by controlling the reaction temperature (75–200 °C) and amount of Br₂ (0.6–1.2 mmol) under air-free synthetic condition (cf. ESI; Experimental section[†]). The "three-precursor" nature¹⁶⁻¹⁸ of our synthetic scheme allows for independent tuning of the amount of the individual elements viz., Cs⁺, Pb²⁺ and X⁻ ions and in turn, allows for the precise control over the surface chemistry.17 Highly crystalline, monodisperse 7.62 \pm 1.0 nm sized cubic α -CsPbBr₃ NCs (Fig. 1a) were synthesised under optimized conditions using Cs : Pb : Br₂ ratio of 1 : 1 : 6 at 200 °C. Highresolution transmission electron microscopy (HRTEM) and powder X-ray diffraction (PXRD) analyses of the purified sample confirmed the pure cubic perovskite phase. HRTEM reveals (100) lattice fringes with a d-spacing of 0.58 nm (Fig. S2[†]). PXRD spectra of α-CsPbBr3 NCs sample exposed to ambient air (relative humidity of \sim 50–60%) for a period of 60 days showed no alteration (Fig. 1b). The observed stability is much higher than the previously reported (max. 8 days) achieved via passivation of cubic CsPbBr3 NCs with ZnBr2.12 Untreated pristine cubic CsPbBr₃ NCs, synthesised via conventional method typically transforms into orthorhombic phase within 1-2 days.¹² We reproduced halide salt passivated CsPbBr3 NCs reported by Woo et al. For comparison, we purified and stored the NCs (control) under conditions similar to our cubic CsPbBr₃ NCs.

Consistent with their report, the halide passivation did improve the stability from 2 days to 6 days, thereafter the phase

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Fig. 1 (a) TEM image of cubic CsPbBr₃ perovskite NCs (size: 7.62 \pm 1.0 nm) (b) XRD patterns showing film stability of CsPbBr₃ NCs over a period of 60 days in the air under ambient condition (relative humidity ~50–60%). The pattern of α -CsPbBr₃ (JCPDS 00-054-0752) is indicated as black bars for comparison. (c) UV-vis absorption and normalised PL spectra of as-prepared (day 1) α -CsPbBr₃ NCs and the same sample stored in ambient air for 60 days as the colloidal solution (hexane). (d) Integrated photoluminescence vs. absorbance plot²⁹ for day 1 (black) and day 60 samples (red). Absorbance is measured at the excitation wavelength (400 nm).

changed to orthorhombic phase (Fig. S3[†]). In our case, the synthesized α -CsPbBr₃ the diffraction peak at 30.5 degrees (2 θ) did not split even after 60 days (Fig. S4[†]) ruling out the presence of the orthorhombic phase. In the case of conventionally prepared CsPbBr₃ NCs, the peaks become visibly sharper and split within a week of exposure to ambient air, indicative of the phase transformation into orthorhombic (Pnma) phase.12,19,20 Furthermore, comparison of the integrated photoluminescence vs. absorbance plot for the freshly prepared NCs (day 1) and for the sample stored as the colloidal solution at ambient temperature for 60 days shows the insignificant difference between the two slopes signifying no change in quantum efficiency within the experimental error (Fig. 1c and d). Previous studies have shown that the phase stability of α-CsPbX3 NCs is extremely sensitive to surface chemistry.6,11,17 To understand the cause of the exceptional stability in our case, a series of additional experiments were conducted. FTIR spectra of the purified NCs show the signature of protonated amine groups $(-NH_3^+)^{21}$ at 1641 cm⁻¹ in addition to the peak associated to oleate anions $(-COO^{-})$ at 1576 cm⁻¹ (Fig. 2a).

Furthermore, the ¹H NMR shows a multiplet at 7.2 ppm (Fig. 2b) which is uncorrelated with any other proton peak in ${}^{1}H{-}^{1}H$ COrrelated SpectroscopY (COSY) NMR spectra (Fig. S5†). This feature is attributed to diastereotopically coupled protons of ammonium ions bound to the surface of the NCs.^{11,22} The presence of oleylammonium ions tightly bound to the NC surface is further established with X-ray photoelectron spectroscopy (XPS) analyses. The XPS spectra corrected using the maximum of the C 1s signal at 284.8 eV (Fig. S6†) shows



Fig. 2 (a) FTIR spectra of the purified CsPbBr₃ NCs (black) compared to pure oleyl amine (red). (b) ¹H NMR of purified α -CsPbBr₃ NCs in CDCl₃ (black) (c) XPS N 1s and (d) Pb 4f core level XPS spectra CsPbBr₃ NCs.

a prominent N 1s peak at 401.8 eV characteristic of bound -NH₃⁺ group (Fig. 2c).²¹ Furthermore, the Pb 4f core level spectra of ambient stable show two peaks located at 138.4 and 143.2 eV assigned to the Pb $4f_{7/2}$ and Pb $4f_{5/2}$ levels respectively (Fig. 2d). The observed binding energy value is on the higher side, consistent with observation reported for CsPbBr₃ NCs prepared via passivation with ZnBr₂ (ref. 12) and tetrafluoroacetate.⁴ Analysis of the N 1s, Cs 3d, Pb 4f/5d, and Br 3d/3s peaks revealed that the obtained a-CsPbBr3 NCs have bromide ion rich surface (Br/Pb > 3) (Fig. S7[†]). The estimated N : Cs ratio is 1.56. These results confirmed the presence of ammonium ligand on the lead halide-rich surfaces. The synthetic approach is equally applicable to cubic α-CsPbI3 quantum dots, which are known to be even more sensitive to phase transformation in ambient air (cf. ESI[†]). In general, under X₂-rich condition and with optimized Cs^+ ion concentration ($[Cs^+] \leq [Pb^{2+}]$), the surface of the CsPbX₃ NCs can be concomitantly passivated with halide and ammonium ions to achieve long-term stability. Highly crystalline, phase-stable 11.11 nm sized α -CsPbI₃ perovskite NCs (Fig. S8[†]) were obtained using a Cs : Pb : I₂ precursor ratio of 1:1:2 at 200 °C. They were stable in ambient air (relative humidity of \sim 50-60%, room temperature) for a period of 20 days (Fig. S9[†]) and for over 4 months in inert condition (Fig. S10[†]). Similar to the CsPbBr₃ NCs surface, the presence of ammonium species (Cs:N \sim 1.12) and slightly excess halide ions (Cs : Pb : I = 0.8 : 1 : 3.4) were confirmed by XPS studies (Fig. S11 and S12[†]). The presence of surface-bound ammonium ion was further confirmed by FTIR (Fig. S13[†]) and 1D/2D NMR (Fig. S14 and S15[†]). The presence of firmly surfacebound oleylammonium ligands is considered important for the in situ stabilisation of the cubic phase of CsPbX₃ NCs.¹¹ In fact, the Goldschmidt tolerance factor in cubic $CsPbX_3$ (X = Br, I) perovskite is lower than the ideal value of 1, mainly due to the

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small size of Cs⁺ ions compared to the void.²³ This effect is accentuated in the case of X = I. Consequently, $[PbX_6]^{4-}$ octahedra undergo distortion to decrease the extra space resulting in a transformation from the cubic to the orthorhombic phase. Partial substitution of Cs⁺ ions on NC surface layers with larger alkyl ammonium ions reduces this distortion, resulting in the observed enhanced stability of the cubic phase. In addition to the presence of ammonium ions, the detection of stoichiometricallyexcess halide ions by XPS in our case suggests a possible concomitant role of the halide ions in stabilising the NCs. Kim *et al.* have demonstrated that halide-amine passivation of fractional dangling electrons satisfies the 8-electron rule on the highly faceted tetrahedral shaped InP NCs surfaces.²⁴ Based on this we propose the co-passivation of α -CsPbX₃ NCs with ammonium ions and halide for enhanced stability.

To further gain mechanistic insight into the origin of surface-bound ammonium ions, we examined the fate of OAm in the presence of OA and Br_2 at reaction temperature (~200 $^{\circ}$ C). In the presence of amine, X₂ is reduced to a strong acid HX in situ.25 Therefore, the molecular bromine plays the double role of supplying halide ions and modulating the acid/base (OAm/ OA) ligand chemistry via protonation of OAm in the solution. In conventional ligand system, OAm is protonated by OA in solution, resulting in the formation of oleylammonium oleate. As expected Fig. 3a shows the downfield shift of α -CH₂ resonance peak of pure OAm from 2.66 to 2.85 ppm (25 °C) on the addition of OA. However, upon heating of the OA/OAm mixture to 200 °C an upfield shift of the α-CH₂ resonance peak was observed, which is due to the dissociation of oleylammonium oleate. Its formation from OA and OAm being an exothermic reaction, Le Chatelier's principle predicts indeed that the equilibrium shifts toward the reactants when the temperature is increased.²⁶ The depletion of oleylammonium species at high temperature in typical OA/OAm synthetic methods leads to the instability of the cubic phase. Therefore, CsPbX₃ perovskite NCs prepared by conventional methods involving the OA/OAm ligand system is thermally unstable and additional oleylammonium halide ions are required to impede rapid phase change at elevated temperature.11 Typically, synthesis is carried out at 160 °C or lower temperature followed by rapid quenching in an ice bath.11 To the contrary, we found that in the presence of Br₂, the OAm/OA equilibria behave differently with a change in temperature. Fig. 3b presents the NMR study of OAm/Br₂ mixture at 25 °C and 200 °C. The α-CH₂ proton resonance of OAm changes from 2.66 to 2.70 ppm due to protonation of OAm



Fig. 3 (a) ¹H NMR spectra of free OAm (black), OAm/OA mixture at 25 °C (red) and OAm/OA mixture at 200 °C (blue) showing change in α -CH₂ (amine) resonance peak. (b) ¹H NMR spectra of free OAm (black); OAm/Br₂ mixture at 25 °C (red) and at 200 °C (blue).

with HBr formed in situ. Importantly, when the OAm/Br₂ solution is heated to the reaction temperature (~ 200 °C), the α -CH₂ (amine) resonance peak shifts further downfield to 2.87 ppm showing the presence of thermally stable ammonium halide ion species. We propose that oleylammonium bromide has been stabilised via self-association or complexation with excess amine present in the mixture.25 The peak broadening at 2.87 ppm also supports that different ammonium halide species (self-associated) are in equilibrium, collectively contributing to the time-averaged NMR spectra. A similar effect of temperature on the NMR chemical shift was reported for pyridine (Py) and HCl mixtures, attributed to the formation of $[Py-H-Py]^+$ and $(PyHCl)_2$ species.²⁷ Furthermore, OAm/I₂ mixture exhibited similar characteristics (Fig. S16†), underlining the general nature of the mechanism of phase stabilization in $CsPbX_3$ (X = Br, I) NCs.

Both CsPbI₃ and CsPbBr₃ NCs exhibit emission tunability in a wide spectral range by varying the reaction temperature from 75 °C to 200 °C (Fig. 4). The obtained NCs exhibit narrow PL emission (FWHM: 29–48 nm) and high PL QYs (\sim 40–79%). Lower emission wavelengths could easily be obtained by simply decreasing the reaction temperature (Fig. 4) or increasing the X₂ concentration (Fig. S17†).

The crystallite sizes calculated from XRD reveals that the change in emission wavelength is due to change in size of the NCs (Fig. S18[†]). Furthermore, NCs synthesized at higher temperature show lower absorbance (nearly 3-fold) at same photon energy (350 nm), compared to smaller sized NCs



Fig. 4 (a) UV illuminated photographs of colloidal solutions of different sized CsPbX₃ NCs (X = Br, I). Absorption (b) and emission spectra (c) of CsPbBr₃ NCs at different reaction temperatures (75–200 °C). Absorption (d) and emission spectra (e) of CsPbI₃ NCs at different reaction temperatures (75–200 °C).

synthesized at lower temperature (Fig. S19[†]). This suggests that the higher temperature produces less nuclei, though of an overall larger final size. On the other hand, at the same temperature increasing the halide concentration leads to increase in supersaturation and hence the increased nucleation rate. By combining the emission ranges of the two materials, the entire visible spectrum from blue to red (478-714 nm) can be accessed. The blue-emitting CsPbBr₃ QDs (PL peak 478 nm) obtained at low temperature (~75 °C) under Br₂-rich condition deserve particular attention. Attaining this wavelength is considered difficult due to the small size of CsPbBr₃ QDs and their high surface sensitivity.28 It is noteworthy that phase stability of smaller sized NCs synthesized at lower temperature under X2-rich condition were relatively poor compared to ones synthesized at higher temperature (~ 200 °C) at same X₂ concentration (Table S1[†]). It has been suggested that due to bulky nature of OAm⁺ ions, its binding on the surface of the NCs is less feasible at lower temperature.11

In conclusion, we demonstrate a versatile synthetic scheme for achieving long-term ambient stability and emissiontunability of cubic CsPbX₃ (X = Br, I) perovskite NCs. The use of molecular X_2 allows for the precise control of surface composition, crystal structure and optical properties of the NCs. The key mechanistic insights obtained herein are relevant for the halide perovskite NC synthesis in general.

Conflicts of interest

The authors declare no conflicts of interest.

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References

- 1 Q. A. Akkerman, G. Rainò, M. V. Kovalenko and L. Manna, *Nat. Mater.*, 2018, **17**, 394–405.
- 2 D. Aldakov and P. Reiss, J. Phys. Chem. C, 2019, 123, 12527-12541.
- 3 X. Zhang, H. Lin, H. Huang, C. Reckmeier, Y. Zhang, W. C. H. Choy and A. L. Rogach, *Nano Lett.*, 2016, **16**, 1415–1420.
- 4 H. Wang, X. Zhang, Q. Wu, F. Cao, D. Yang, Y. Shang, Z. Ning, W. Zhang, W. Zheng, Y. Yan, S. V. Kershaw, L. Zhang, A. L. Rogach and X. Yang, *Nat. Commun.*, 2019, 10, 665.

- 5 B. Li, Y. Zhang, L. Fu, T. Yu, S. Zhou, L. Zhang and L. Yin, *Nat. Commun.*, 2018, **9**, 1–8.
- 6 A. Swarnkar, A. R. Marshall, E. M. Sanehira,
 B. D. Chernomordik, D. T. Moore, J. A. Christians,
 T. Chakrabarti and J. M. Luther, *Science*, 2016, 354, 92–95.
- 7 Y. Wang, X. Li, J. Song, L. Xiao, H. Zeng and H. Sun, *Adv. Mater.*, 2015, 27, 7101–7108.
- 8 X. Zhu, Y. Lin, Y. Sun, M. C. Beard and Y. Yan, *J. Am. Chem. Soc.*, 2019, **141**, 733–738.
- 9 X. Chen, H. Hu, Z. Xia, W. Gao, W. Gou, Y. Qu and Y. Ma, *J. Mater. Chem. C*, 2017, **5**, 309–313.
- 10 A. Dutta and N. Pradhan, ACS Energy Lett., 2019, 4, 709-719.
- 11 A. Dutta, S. K. Dutta, S. Das Adhikari and N. Pradhan, *Angew. Chem., Int. Ed.*, 2018, **57**, 9083–9087.
- 12 J. Y. Woo, Y. Kim, J. Bae, T. G. Kim, J. W. Kim, D. C. Lee and S. Jeong, *Chem. Mater.*, 2017, **29**, 7088–7092.
- 13 N. Park, M. Grätzel, T. Miyasaka, K. Zhu and K. Emery, *Nat. Energy*, 2016, **1**, 16152.
- 14 C. Wang, S. R. Chesman and J. J. Jasieniak, *Chem. Commun.*, 2017, **53**, 232–235.
- 15 J. Pan, Y. Shang, J. Yin, M. De Bastiani, W. Peng, I. Dursun, L. Sinatra, A. M. El-Zohry, M. N. Hedhili, A.-H. Emwas, O. F. Mohammed, Z. Ning and O. M. Bakr, *J. Am. Chem. Soc.*, 2018, **140**, 562–565.
- H. Sun, Z. Li, L. Kong, B. Wang, C. Zhang, Q. Yuan, S. Huang,
 Y. Liu and L. Li, *Chem. Commun.*, 2018, 54, 9345–9348.
- 17 M. Imran, V. Caligiuri, M. Wang, L. Goldoni, M. Prato, R. Krahne, L. De Trizio and L. Manna, *J. Am. Chem. Soc.*, 2018, **140**, 2656–2664.
- 18 Q. A. Akkerman, L. Mart, L. Goldoni, M. Imran, D. Baranov,
 H. J. Bolink, F. Palazon and L. Manna, *Chem. Mater.*, 2018,
 30, 6915–6921.
- 19 Y. Cai, L. Wang, T. Zhou, P. Zheng, Y. Li and R.-J. Xie, *Nanoscale*, 2018, **10**, 21441–21450.
- 20 C. Tenailleau, S. Aharon, B.-E. Cohen and L. Etgar, *Nanoscale Adv.*, 2019, **1**, 147–153.
- 21 J. Liu, K. Song, Y. Shin, X. Liu, J. Chen, K. X. Yao, J. Pan, C. Yang, J. Yin, L.-J. Xu, H. Yang, A. M. El-Zohry, B. Xin, S. Mitra, M. N. Hedhili, I. S. Roqan, O. F. Mohammed, Y. Han and O. M. Bakr, *Chem. Mater.*, 2019, DOI: 10.1021/ acs.chemmater.9b00680.
- 22 V. K. Ravi, P. K. Santra, N. Joshi, J. Chugh, S. K. Singh, P. Ghosh and A. Nag, *J. Phys. Chem. Lett.*, 2017, **8**, 4988–4994.
- 23 A. Swarnkar, W. J. Mir and A. Nag, ACS Energy Lett., 2018, 3, 286–289.
- 24 K. Kim, D. Yoo, H. Choi, S. Tamang, J. H. Ko, S. Kim, Y. H. Kim and S. Jeong, *Angew. Chem., Int. Ed.*, 2016, 55, 3714–3718.
- 25 M. J. Kogan and J. C. Schug, J. Magn. Reson., 1973, 11, 406-415.
- 26 G. Almeida, L. Goldoni, Q. Akkerman, Z. Dang, A. H. Khan, S. Marras, I. Moreels and L. Manna, ACS Nano, 2018, 12, 1704–1711.
- 27 C. A. Angell and J. W. Shuppert, *J. Chem. Phys.*, 1977, **67**, 3050–3056.
- 28 V. Malgras, J. Henzie, T. Takei and Y. Yamauchi, *Angew. Chem., Int. Ed.*, 2018, **57**, 1–6.
- 29 S. Tamang, G. Beaune, I. Texier and P. Reiss, *ACS Nano*, 2011, 5, 9392–9402.