Journal of Materials Chemistry A



View Article Online

PERSPECTIVE

Check for updates

Cite this: J. Mater. Chem. A, 2024, 12, 19

Received 3rd September 2023 Accepted 1st November 2023

DOI: 10.1039/d3ta05315f

rsc.li/materials-a

Temperature and pressure induced structural transitions of lead iodide perovskites

Pratap Vishnoi ()* and C. N. R. Rao ()*

During the past fourteen years, conventional lead iodide perovskites, $APbI_3$ [A = Cs⁺, methylammonium $(CH_3NH_3^+)$ and formamidinium $(NH_2CHNH_2^+)]$, have emerged as the forerunner materials for optoelectronic applications, including photovoltaics. However, their photoactive phases suffer from structural instabilities including phase transitions under extrinsic factors such as heat, light, pressure etc. They even show chemical instability in water, oxygen and humid air. Consequently, their optical and electronic properties show marked changes under device working conditions, which is one of the major obstacles in long-term use. At the same time, phase transition offers plenty of opportunities for modulating structures to physical properties. Hence, there is a need for a comprehensive understanding of the structural instabilities and phase transitions of APbl₃ perovskites. This perspective focuses on three conventional lead iodide perovskites, CsPbl₃, MAPbl₃ and FAPbl₃, which are among the most widely studied materials in this family. We critically review the experimentally known structures as a function of temperature or pressure. We describe crystal structures of these three perovskites as well as highlight the reason for their instabilities, mechanism of structural transitions, and fundamental insights into their optoelectronic properties. In addition to temperature- and pressure-driven structural changes, we highlight structural changes due to chemical instability towards water and oxygen. We focus mainly on the structures reported based on X-ray, synchrotron, and neutron diffraction data. We conclude the perspective by discussing current challenges in these exciting materials, possible future opportunities to improve them and our thoughts on new directions.

New Chemistry Unit, International Centre for Materials Science, School of Advanced Materials, Jawaharlal Nehru Centre for Advanced Scientific Research (JNCASR), Jakkur, Bangalore, 560064, India. E-mail: pvishnoi@jncasr.ac.in; cnrrao@jncasr.ac.in



Pratap Vishnoi

Pratap Vishnoi is an Assistant Professor jointly in the New Chemistry Unit (NCU) and the International Centre for Materials Science (ICMS) at the Jawaharlal Nehru Centre for Advanced Scientific Research (JNCASR), Bangalore. He received his PhD in chemistry from the Indian Institute of Technology Bombay (IIT Bombay) under the supervision of Prof. R. Murugavel. He then worked as a Department of Science & Technology post-

doctoral fellow with Prof. C. N. R. Rao at JNCASR. Subsequently, he carried out post-doctoral studies at the University of California Santa Barbara (UCSB) with Prof. Ram Seshadri and Prof. Sir Anthony K. Cheetham. Recently, he was named an Emerging Investigator-2022 by the Journal of Materials Chemistry A. His current research interests are in multifunctional hybrid organic-inorganic materials, halide perovskites and layered nanomaterials.



C. N. R. Rao is the Linus Pauling Research Professor at the JNCASR, Bangalore and an honorary professor at the Indian Institute of Science, Bangalore. He is a member of several scientific academies including the Royal Society and the US National Academy of Sciences. He is the recipient of numerous prestigious awards including the Royal Medal of the Royal Society, the Dan David Prize, the Illy Trieste Science Prize, the

C.N. R. Rao

August von Wilhelm Hoffmann medal of the German Chemical Society, the von Hippel Award of the Materials Research Society, the Eni International Award for Research in Energy Frontiers and the Sheikh Saud International Prize for Materials Science. His research interests are in the chemistry of materials, especially 2D materials and their chemical functionalization, reduction of water and CO_2 and aliovalent substitutions.

1. Introduction

For several decades, conventional oxide perovskites have been intensely explored as phase transition materials by condensed matter physics and solid-state chemistry communities. Some common examples of oxides include ferroelectric titanate materials and high-temperature cuprate superconductors.^{1,2} On the other hand, lead halide perovskites, with the general chemical formula of APbI₃, have come to the fore very recently, opening an entirely new avenue of phase transition materials. They are composed of a three-dimensional [PbI₃]⁻ framework of corner-sharing [PbI₆]⁴⁻ octahedra and mono-cation, known as A-cation, in the interstices (Fig. 1a). The A-cation is typically a large inorganic cation, such as Cs⁺ or an organic cation, such as methylammonium $(CH_3NH_3^+)$ and formamidinium (NH_2^-) CHNH₂⁺). APbI₃ perovskites with these three A-cations and their mixed cation variants are among the most widely studied compounds of the lead halide perovskite family. Their semiconducting behaviour combined with narrow, tuneable and direct bandgaps (Fig. 1b) gives rise to exceptional optoelectronic properties.^{3,4} They exhibit high absorption coefficients, broadspectrum light absorption, low exciton binding energies, high charge-carrier mobilities, long carrier diffusion lengths and defect tolerance.^{5,6} The widespread success of APbI₃ perovskites is also due to their ability to be solution-processed into thin films. These are all beneficial for high efficiency of solar cells and other optoelectronic devices.

Another fascinating feature of lead iodide perovskites is that they offer substantial opportunities for modulating their structures and physical properties through phase transitions.⁷ The ideal perovskite adopts a cubic $Pm\bar{3}m$ structure typically at high temperature with no structural distortion. Low symmetry phases are obtained from this structure by changing temperature or pressure. The A-cation tumbling dynamics and the $[PbI_6]^{4-}$ octahedral tilting are the primary reasons for phase transitions. The phase transition is intriguing, yet a flipside exists: their phase transition temperatures range across room temperature, causing marked changes in the physical properties under device working conditions, which is one of the major obstacles to their use.⁸⁻¹⁰ Additionally, some of them form photoinactive non-perovskites under ambient conditions.¹¹ Thus, understanding crystal structures of the different phases is



Fig. 1 (a) Crystal structure of an archetypical $APbI_3$ perovskite in the cubic $Pm\overline{3}m$ symmetry. (b) Schematic presentation of optical bandgaps.

crucial because all the physical properties are dependent on the structure in addition to chemical composition. Table 1 lists the structural transition pathways of CsPbI₃, MAPbI₃ and FAPbI₃ found under heat and pressure. Tables 2 and 3 list the structural parameters of the temperature and pressure induced phases, respectively. Clearly, there are diverse views about the transition pathways as well as crystallographic symmetries and there is a need for having a deeper understanding of their structures.

In this perspective, we critically examine the structures of various phases of CsPbI₃, MAPbI₃, and FAPbI₃ perovskites, reported in the crystallographic database. Our aim is to understand structures and underlying optoelectronic properties. Our perspective is expected to raise the current understanding level of structures at the atomic level and provide some new strategies to further design and synthesize stable halide perovskites.

2. Thermodynamics of solid-to-solid phase transitions

A solid-to-solid phase transition of a material is a transition between its different crystalline forms (also known as phases or polymorphs), each of which possesses uniform structure and composition throughout and separates from the other form by the phase boundary.³⁵ Thermodynamically, phase transition is the transformation of a thermodynamic system from one phase into another in response to a change in one or more thermodynamic variables (such as temperature or pressure) which is distinguishable by a change in one or more thermodynamic quantities or order parameters (such as entropy, enthalpy, heat capacity *etc.*).³⁶ Temperature and pressure can be precisely controlled to study changes in the free energy and the order parameters of a phase. Entropy (*S*) and volume (*V*) are first order parameters, *i.e.* first derivatives of the Gibbs free energy as defined using equations:

$$-\left(\frac{\partial G}{\partial T}\right)_{\rm P} = S \tag{1}$$

$$\left(\frac{\partial G}{\partial P}\right)_{\rm T} = V \tag{2}$$

Heat capacity, volume thermal expansivity, compressibility and magnetic susceptibility are some of the second order parameters, *i.e.* second derivatives of the Gibbs free energy as defined using equations:

$$\left(\frac{\partial^2 G}{\partial T^2}\right)_{\rm p} = -\left(\frac{\partial S}{\partial T}\right)_{\rm p} = -\frac{C_{\rm p}}{T}$$
(3)

$$\left(\frac{\partial^2 G}{\partial P \partial T}\right) = \left(\frac{\partial V}{\partial T}\right)_{\rm P} = V\alpha \tag{4}$$

$$\left(\frac{\partial^2 G}{\partial P^2}\right)_{\rm T} = \left(\frac{\partial V}{\partial P}\right)_{\rm T} = -V\beta \tag{5}$$

$$\left(\frac{\partial^2 G}{\partial H^2}\right)_{\rm T} = \left(\frac{\partial M}{\partial H}\right)_{\rm T} = \chi_{\rm T} \tag{6}$$

Perspective

 Table 1
 Temperature and pressure induced phases of APbI₃ perovskites, and their transition pathways reported in different studies. The orange and blue highlighted rows indicate temperature and pressure induced transitions, respectively

Phase transition pathways	Ref.			
CsPbI ₃				
Cubic $Pm\overline{3}m \xrightarrow{533 \text{ K}}$ tetragonal P4/mbm $\xrightarrow{448 \text{ K}}$ orthorhombic Pnma	12			
Cubic $Pm\overline{3}m \xrightarrow{602 \text{ K}}$ orthorhombic <i>Pnma</i> (non-perovskite)	13			
Orthorhombic <i>Pnma</i> (non-perovskite) $\xrightarrow{3.9 \text{ GPa}}$ monoclinic <i>P</i> 2 ₁ / <i>m</i> (non-perovskite)	14			
Orthorhombic <i>Pnma</i> (non-perovskite) $\xrightarrow{6.9 \text{ GPa}}$ monoclinic <i>C</i> 2/ <i>m</i> (non-perovskite)	15			
MAPbl ₃				
Cubic $Pm\overline{3}m \xrightarrow{327.4 \text{ K}}$ tetragonal /4/mcm $\xrightarrow{162.2 \text{ K}}$ orthorhombic $Pna2_1$	16			
Cubic $Pm\overline{3}m \xrightarrow{327-330 \text{ K}}$ tetragonal <i>I</i> 4/ <i>mcm</i> $\xrightarrow{161 \text{ K}}$ orthorhombic <i>Pnma</i>	17–20			
Tetragonal P4mm $\xrightarrow{333 \text{ K}}$ tetragonal I4cm $\xrightarrow{\sim 150 \text{ K}}$ γ -phase (space group not provided)	21			
Tetragonal I4/mcm $\xrightarrow{0.35 \text{ GPa}}$ cubic Im $\overline{3} \xrightarrow{2.5 \text{ GPa}}$ cubic Im $\overline{3} \xrightarrow{>2.5 \text{ GPa}}$ amorphous	22			
Tetragonal I4/mcm $\xrightarrow{0.4 \text{ GPa}}$ cubic Im $\overline{3} \xrightarrow{2.5 \text{ GPa}}$ orthorhombic Immm $\xrightarrow{2.7 \text{ GPa}}$ amorphous	23			
Tetragonal I4/mcm $\xrightarrow{0.4 \text{ GPa}}$ orthorhombic Imm_2	24			
Tetragonal I4/mcm $\xrightarrow{0.26 \text{ GPa}}$ orthorhombic $Imm_2 \xrightarrow{3.0 \text{ GPa}}$ amorphous	25			
Orthorhombic <i>Fmmm</i> $\xrightarrow{0.6 \text{ GPa}}$ cubic <i>Im</i> $\overline{3} \xrightarrow{2.9 \text{ GPa}}$ amorphous	26			
FAPbl ₃				
Cubic $Pm\overline{3}m \xrightarrow{300 \text{ K}}$ hexagonal $P6_3/mmc$ (non-perovskite)	27			
Cubic $Pm\overline{3}m \xrightarrow{300 \text{ K}} \text{hexagonal } P6_3/mmc$ (non-perovskite) $\xrightarrow{175 \text{ K}} \text{hexagonal } P6_3/m$	28 and 29			
Cubic $Pm\overline{3}m \xrightarrow{285 \text{ K}} \text{tetragonal } P4/mbm \xrightarrow{140 \text{ K}} \text{tetragonal } P4/mbm$	29–31			
Cubic $Pm\overline{3}m \xrightarrow{0.30 \text{ GPa}}$ orthorhombic $Imm_2 \xrightarrow{1.7 \text{ GPa}}$ orthorhombic $Immm \xrightarrow{>4 \text{ GPa}}$ amorphous	32			
Cubic $Pm\overline{3}m \xrightarrow{0.46 \text{ GPa}} \text{tetragonal } P4/mbm \xrightarrow{3.13 \text{ GPa}} \text{cubic } Im\overline{3} \xrightarrow{6.59 \text{ GPa}} \text{amorphous}$	33			
Cubic $Pm\overline{3}m \xrightarrow{0.49 \text{ GPa}}$ tetragonal <i>P4/mbm</i>	34			
Hexagonal $P6_3/mmc$ (non-perovskite) $\xrightarrow{1.26-1.73 \text{ GPa}}$ orthorhombic $Cmc2_1$ (non-perovskite)	33			

Published on 03 2023. Downloaded by Fail Open on 23/07/25 10:34:24.

Here, C_p is the heat capacity at constant pressure, α is the thermal expansivity, β is the compressibility and χ is the magnetic susceptibility.

According to Ehrenfest,³⁷ a phase transition is defined as an n^{th} order transition, if at-least one of the n^{th} order derivatives of the Gibbs free energy, such as entropy (*S*), volume (*V*), enthalpy (*H*), heat capacity (C_p) *etc.*, exhibits discontinuity at the transition point. Thus, in the first order phase transition, at-least one of the first derivatives exhibits a discontinuity. Additionally, the specific heat capacity (C_p) shows extremely sharp and symmetrical shape of the peak at the transition

point in a first order transition.^{38,39} According to modern thermodynamics, a first order phase transition includes latent heat (or change in enthalpy) ($\Delta H = T\Delta S$), making the first order transition materials promising heat storage systems.⁴⁰ When first order derivatives change continuously but at-least one of the second-order derivatives (such as C_p) shows discontinuity, the phase transition is defined to be a second order. Higher order transitions are also possible but are perhaps not very useful as the difference between the thermodynamic quantities becomes smaller.

Table 2 Structural parameters of the structures reported for $APbI_3$ at ambient pressure and different temperatures (7). The exact conditions are given in parentheses

			Atomic coordinates				
Compound	Space group	Unit cell constants	Atom	x	у	z	Occ.
CsPbI ₃	Cubic $Pm\bar{3}m$ (no. 221) ¹² (α -phase)	$a = b = c = 6.29658(2)$ Å, $\alpha = \beta = \gamma =$ 90° ($T = 646.6$ K)	Cs	0.5	0.5	0.5	1
			Pb	0.0	0.0	0.0	1
			Ι	0.5	0.0	0.0	1
	Tetragonal <i>P</i> 4/ <i>mbm</i> (no. 127) ¹² (β-phase)	a = b = 8.82690(3) Å and $c =$	Cs	-0.5	0.0	0.5	1
		6.29902(3) Å, $\alpha = \beta = \gamma = 90^{\circ} (T = 511.7 \text{ K})$	Pb	0.0	0.0	0.0	1
			I1	0.0	0.0	0.5	1
			I2	-0.212	0.288	0.0	1
	Orthorhombic <i>Pbnm</i> (no. 62) ¹² (γ-phase)	a = 8.61979(6) A, b = 8.85176(6) A and $c = 12.50128(7) \text{ Å}; \alpha = \beta = \gamma = 90^{\circ} (T = 324.6 \text{ K})$	Cs	-0.0165	0.4451	0.25	1
			Pb	0.0	0.0	0.5	1
			I1	0.1973	0.2993	0.5306	1
			I2	-0.0631	-0.0017	0.75	1
	Orthorhombic <i>Pbnm</i> (no. 62) ⁵⁸ (δ-phase, non-	a = 10.4581(3) A, $b = 4.80171(13)$ A and $c = 17.7761(5)$ Å; $\alpha = \beta = \gamma = 90^{\circ}$ (T = 298 K)	Cs	0.4156	0.25	0.82924	1
			Pb	0.16049	0.25	0.06185	1
	perovskite)		I1	0.2997	0.25	0.2127	1
			I2	0.3355	0.75	0.9979	1
			13	0.0331	0.25	0.88541	1
MAPbI ₃	Cubic $Pm3m$ (no. 221) ¹³	$a = b = c = 6.31728(27) \text{ A}; \alpha = \beta = \gamma =$	Pb	0.0	0.0	0.0	1
	$(\alpha$ -phase)	$90^{\circ} (T = 352 \text{ K})$	I	0.5	0.0	0.0	1
	Tetragonal $P4mm$ (no. 99) ²¹	a = b = 6.3115(2) Å and $c = 6.3161(2)Å; \alpha = \beta = \gamma = 90^{\circ} (T = 400(2) \text{ K})$	Pb	0.5	0.5	0.50813	1
	(α-phase)		I1	0.5	0.0	0.5122	1
			12	0.5	0.5	-0.0006	1
	Tetragonal <i>I4/mcm</i> (no. 140) ¹⁹ (β-phase)	a = b = 8.80625(28) A and $c = 8.80625Å; \alpha = \beta = \gamma = 90^{\circ} (T = 180 \text{ K})$	Pb	0.5	0.5	0.5	1
			I1	0.5	0.5	0.75	1
		· · · · · · · · · · · · · · · · · · ·	12	0.20122	0.29878	0.5	1
	Tetragonal <i>I4cm</i> (no. 108) ³⁹	a = b = 8.8438(3) A and $c = 12.6215(5)$	Pb	0.0	0.0	0.00012	1
	(β-phase) A;	A; $\alpha = \beta = \gamma = 90^{\circ}$ (T = 293 K)	I1	0.0	0.0	0.25	1
	Tetragonal <i>I</i> 422 (no. 97) ⁶⁰ $a = b = 8.88375(18)$ Å and $c = (\beta-phase)$ 12.7010(3) Å; $\alpha = \beta = \gamma = 90^{\circ}$ ($T = 293$ K)		12	0.2135	0.7135	0.0007	1
		a = b = 8.88375(18) A and $c =$	Pb	1.0	1.0	0.7498	1
		12.7010(3) A; $\alpha = \beta = \gamma = 90^{\circ} (T = 293)$	I1_1	1.0	1.0	0.5	1
		K)	I1_2	1.0	1.0	1.0	1
		· · · · · · · · · · · · · · · · · · ·	12	0.28604	0.78604	0.75	1
	Orthorhombic <i>Fmmm</i> (no. 69) ²⁶	a = 12.4984(7) A, b = 12.5181(7) A and $c = 12.6012(8) \text{ Å}; \alpha = \beta = \gamma = 90^{\circ} (T = 296(2) \text{ K})$	Pb	0.25	0.25	0.5	1
			11	0.2241	0.5	0.5	1
			I2A	0.5	0.2981	0.5	0.75
FAPbI ₃			12B	0.5	0.267	0.462	0.125
			I3A Iap	0.25	0.25	0.25	0.75
			I3B	0.294	0.25	0.25	0.125
	Orthorhombic Phma (no. 62) ¹⁷ (γ -phase) Trigonal P3m1 (no. 156) ²¹ (α -phase)	a = 8.8362(11) A, b = 12.5804(15) A	Pb	0.5	0.0	0.0	1
		and $c = 8.5551(10)$ A; $\alpha = \beta = \gamma = 90^{\circ}$	11	0.485720	0.25	-0.05291	1
		(I = 100 K)	IZ Dh 1	0.1902	0.01/19	0.18615	1
		u = 8.9817(13) A, v = 8.9817(13) A and	PD1 Db2	0.000007	0.3333333	0.837860	1
		$l = 11.000(2)$ A; $\alpha = \beta = 90^{\circ}$ and $\gamma = 1209(T = 202(2))$ K)	PD2 Db2	0.0	0.0	0.50454	1
		$120^{-}(1-293(2) \text{ K})$	PD5 11	0.000007	0.3333333	0.17092	1
			11	0.1002	0.8558	0.53757	1
			12	0.85505	0.10097	0.07107	1
	Cubic $Pm\bar{3}m$ (no. 221) ⁶¹ $a = b$ (α -phase) 90° (? Tetragonal $P4/mbm$ $a = b$ (no. 127) ³⁰ (β -phase) 6.326 200 H 200 H Hexagonal $P6_3mc$ $a = b$ (no. 186) ²¹ (δ -phase, non- $\dot{A}; \alpha =$ perovskite) 293(2) Hexagonal $P6_2/mmc$ $a = b$	$a = b = c = 6.3620(8) \text{ Å}; \alpha = \beta = \gamma = 90^{\circ} (T = 298 \text{ K})$ $a = b = 8.922756(12) \text{ Å and } c = 6.326256(14) \text{ Å}; \alpha = \beta = \gamma = 90^{\circ} (T = 200 \text{ K})$ a = b = 8.6603(14) Å and c = 7.9022(6)	IJ Dh	0.49950	0.0044	0.00447	1
			I	0.5	0.0	0.0	1
			Ph	0.0	0.0	0.0	1
			1 D 11	0.0	0.0	0.5	1
			12	0.230	0.730	0.0	1 1
			Ph	0.230	0.0	0.0	- 1
		$\dot{A} = \beta = 0.0005(14) \text{ A and } t = 7.5022(0)$ $\dot{A} = \beta = 90^{\circ} \text{ and } \gamma = 120^{\circ} (T - 10^{\circ})$	I	0.83117	0.16883	0.251	1 1
		A; $\alpha = \beta = 90^{\circ}$ and $\gamma = 120^{\circ}$ ($T = 293(2)$ K) a = b = 8.5891(5) Å and $c = 7.853(7)$ Å:		0.0011/	0.10005	0.201	*
			Pb	0.0	0.0	0.0	1
	(no. 194) ³¹ (δ -phase. non-	$\alpha = \beta = 90^{\circ}$ and $\gamma = 120^{\circ} (T = 298 \text{ K})$	I	0.82957	0.17043	0.25	1
	perovskite) $\alpha = \beta = 30^{\circ} \text{ and } \beta = 120^{\circ} (1 - 230^{\circ} \text{ K})^{\circ}$	-				-	

Table 3 Structural parameters of the crystal-structures reported for $APbI_3$ at ambient temperature and different pressures (*P*). The exact pressure conditions are given in parentheses

Compound	Space group & unit cell constants
CsPbI ₃	Monoclinic $P2_1/m$ (no. 11) $(P = 3.9 \text{ GPa})^{14} a = 9.832(2) \text{ Å}, b = 4.543(1) \text{ Å and } c = 16.919(2) \text{ Å}; \alpha = \gamma = 90^{\circ} \text{ and } \beta = 90.34(3)^{\circ}$
MAPbI ₃	Monoclinic C2/m (no. 12) ($P = 40$ GPa) ¹³ $a = 15.3402(3)$ A, $b = 3.52911(2)$ A and $c = 18.1236(5)$ A; $\alpha = \gamma = 90^{\circ}$ and $\beta = 132.892(2)^{\circ}$
	Cubic $Im\bar{3}$ (no. 204) $(P = 0.45(2) \text{ GPa})^{22} a = b = c = 12.4067(5) \text{ Å}; a = \beta = \gamma = 90^{\circ}$
	Orthorhombic Immm (no. 71) $(P = 3.21 \text{ GPa})^{23} a = 12.1360(4) \text{ Å}, b = 12.1022(2) \text{ Å and } c = 11.8177(5) \text{ Å}; \alpha = \beta = \gamma = 90^{\circ}$
FAPbI ₃	Orthorhombic Imm2 (no. 44) $(P = 0.40 \text{ GPa})^{23} a = 12.408(3) \text{ Å}, b = 12.419(3) \text{ Å} and c = 12.25(5) \text{ Å}; \alpha = \beta = \gamma = 90^{\circ}$
	Orthorhombic <i>Imm</i> 2 (no. 44) $(P = 0.34 \text{ GPa})^{32} a = 12.4950 \text{ Å}, b = 12.4900 \text{ Å} and c = 12.5428 \text{ Å}; \alpha = \beta = \gamma = 90^{\circ}$
	Orthorhombic <i>Immm</i> (no. 71) ($P = 0.67$ GPa) ³² $a = 12.2264$ Å, $b = 12.4900$ Å and $c = 12.3583$ Å; $\alpha = \beta = \gamma = 90^{\circ}$
	Tetragonal <i>P4/mbm</i> (no. 127) $(P = 0.49 \text{ GPa})^{34} a = b = 8.9148(6)$ Å and $c = 6.3138(7)$ Å; $\alpha = \beta = \gamma = 90^{\circ}$
	Cubic $Im\bar{3}$ (no. 204) $(P = 3.13 \text{ GPa})^{33} a = b = c = 12.0654(6) \text{ Å}; a = \beta = \gamma = 90^{\circ}$
	Orthorhombic $Cmc2_1$ (no. 36) $(P = 1.73 \text{ GPa})^{33}$ (non-perovskite) $a = 14.0957(8)$ Å, $b = 10.2072(4)$ Å and $c = 13.9906(2)$ Å;
	$lpha=eta=\gamma=90^{\circ}$

3. A brief history of the lead iodide perovskite renaissance

The discovery of halide perovskite dates to 1893 when Wells synthesized CsPbX₃ (X = Cl, Br, or I).⁴¹ The X-ray structures were first reported in 1957 by Møller.42 At room temperature, the chloride and the bromide were shown to exhibit 3-dimensional structures, while the iodide had a 1-dimensional chain structure.43 Furthermore, at higher temperatures, all these perovskites were reported to show phase transitions to intensively coloured phases, which was also the first report on the structural transition of lead halide perovskites. In 1978, Weber reported hybrid lead halide perovskites MAPbX₃ (MA = methylammonium; X = Cl, Br or I), which were suggested to adopt cubic symmetry.44 Since then research on lead halide perovskites progressed almost steadily, focusing largely on understanding the nature of structural phase transitions45,46 and electronic properties.^{47,48} The breakthrough happened in 2009 when Kojima et al.,⁴ fabricated dye-sensitized solar cells (DSSCs) by using MAPbBr₃ and MAPbI₃ as photosensitizers. This development quickly captured the wide attention of the PV community. As a consequence, APbI₃ perovskites have become one of the fastest growing PV materials with the solar to power conversion efficiency (PCE) for single-junction solar cells reaching 25% (certified 25.5%),49-52 surpassing the organic photovoltaics by a large margin and rivalling silicon solar cells. The field is still thriving, and many more developments are anticipated in the near future.

4. Origin of structural transitions in lead iodide perovskites

The highest symmetry cubic structure of APbI₃ perovskites consists of a $[PbI_3]^-$ framework of corner-sharing $[PbI_6]^{4-}$ octahedra with all Pb–I–Pb bond angles of 180°, similar to the ReO₃ structure.⁵³ Between the octahedra, large A-cation occupies the 12-coordinated interstices and serves as a structural template for stabilization of the perovskite structure. Due to flexibility of the $[PbI_3]^-$ sub-lattice and dynamic nature of the

A-cation, the cubic perovskite tends to form octahedrally tilted lower-symmetry phases to achieve structural and thermodynamic stability. Without exception, more than one structure is reported for each of the APbI₃ perovskites, depending on temperature, pressure and even the preparation method. To be consistent with the nomenclature of different phases mentioned in the literature, in this article we designate the highest symmetry phase as the α -phase, the phases after first and second transitions as the β -phase and γ -phase, respectively, and the non-perovskite phase as the δ -phase.

The structural stability of perovskites can be empirically prescreened using the tolerance factor (t),⁵⁴ and the octahedral factor (μ) ,⁵⁵ calculated from the ionic radii of the constituent ions using the expressions: $t = (r_A + r_X)/\sqrt{2}(r_X + r_B)$ and $\mu = r_B/r_X$ (here, r_A , r_B , and r_X are the radii of the A-site cation, B-site metal ion and X-site halide, respectively). The octahedral factor assesses formability of the [PbI3]⁻ framework which essentially needs octahedral coordination around the B-site metal ion. Its geometric limit is between 0.414 and 0.732. For lead iodide perovskites, the octahedral factor is 0.541 ($r_{\rm Pb} = 1.19$ Å; $r_{\rm I} = 2.2$ Å) which is well within the perovskite range. The tolerance factor assesses the fitting of the A-cation within the A-site interstices. For a perovskite structure, the value of the tolerance factor should be in the 0.8-1.0 range. At a tolerance factor of 1.0, a perfect fit of the A-cation, thereby an un-tilted cubic structure is expected. A tolerance factor of 0.8 < t < 1.0 indicates that the A-cation is undersized and to strengthen the ionic interaction with the A-cation, the octahedra show tilting with Pb-I-Pb angles less than 180°. The tilting often leads to symmetry lowering from cubic to tetragonal to orthorhombic systems.⁵⁶ With the tolerance factor below 0.80, structures, such as ilmenite-type (FeTiO₃),⁵⁷ are expected to be stable, while structures with face-sharing octahedra are expected to form with a tolerance factor above 1.0.62 In the case of hybrid perovskites, the hydrogen bonding and the order-disorder dynamics of the A-cation also play an important role in addition to the tolerance factor. Thus, the phase transition of the hybrid systems is more complex compared to that of allinorganic analogues.



Fig. 2 Goldschmidt tolerance factors of different $APbI_3$ perovskites with commonly used A-cations, Cs^+ , MA^+ or FA^+ . The tolerance factors of bromides and chloride analogs are also provided for comparison.

The A-site cavity of APbI₃ perovskites is large and, to our knowledge, only three mono-cations, Cs⁺ (1.88 Å) MA⁺ (2.17 Å) and FA⁺ (2.53 Å) are of appropriate size. Some reports ascribe the phase instability of CsPbI₃ to its tolerance factor being less than 0.9.⁶³ The increase in A-cation size in the Cs⁺ < MA⁺ < FA⁺ order adjusts the tolerance factor close to 1.0 (Fig. 2), thereby structurally stabilizing the highest symmetry phase at lower temperature. Furthermore, a decrease in the halide radius adjusts the tolerance factor closer to 1.0 and halide mixing is, therefore, a powerful strategy for making stable perovskite structures.⁶⁴

5. Temperature induced phase transition

5.1. Cesium lead triiodide (CsPbI₃)

CsPbI₃ in the black perovskite phase exhibits a narrow bandgap of 1.7 eV and can effectively absorb visible light when applied as an absorber layer in perovskite-solar cells. However, the black phase is structurally stable only at high temperature and undergoes two types of phase transitions upon decreasing the temperature (Fig. 3a): (i) perovskite to perovskite transition the α -phase (cubic $Pm\bar{3}m$) to the β -phase (tetragonal P4/mbm) to the γ -phase (orthorhombic *Pnma*) and (ii) perovskite to nonperovskite transition – the α -phase to the δ -phase or 'yellow phase' (orthorhombic *Pnma*). Each of the perovskite phases (α -, β - or γ -) is composed of a 3-dimensional [PbI₃]⁻ framework, with Cs⁺ ions in the A-site voids. The perovskite phases differ from each other in the magnitude and the direction of octahedral tilts. The α -phase is the un-tilted form with a Glazer's tilt notation of $a^0 a^0 a^0$. The β -phase shows one in-phase tilt (Glazer's tilt notation $a^0 a^0 c^+$), while the γ -phase shows one in-phase and one out-of-phase tilt (Glazer's tilt notation $a^{-}b^{+}a^{-}$). These three (pseudo-)cubic phases are collectively called the "black" phase as they typically exhibit similar optical properties. They are metastable at room temperature and transform into the thermodynamically more stable δ -phase through the first-order transition. The δ-phase shows a NH₄CdCl₃-type structure,

comprising 1-dimensional double chains of edge-sharing $[PbI_6]^{4-}$ octahedra and the Cs⁺ cation in the interchain space.

Møller in 1958,43 diffracted yellow crystals of CsPbI3 at room temperature on a Geiger-Muller X-ray diffractometer and proposed an orthorhombic chain structure with the Pnma space group (a = 4.795 Å, b = 10.45 Å, and c = 17.76 Å). A color change from yellow to black was reported on heating to 578-581 K. The black phase was reported to have a monoclinically distorted 3D structure (a = b = 6.15 Å and c = 6.23 Å; $\beta = 88.15^{\circ}$). Later studies by Trots et al.58 have shown that CsPbI3 changes from orthorhombic *Pnma* [a = 10.4581(3) Å, b = 4.80171(13) Å, and c= 17.7761(5) Å] to cubic $Pm\bar{3}m$ [a = 6.28940(19) Å] symmetry upon heating at 634 K, but there was no information on the cooling step. Marronnier et al.,12 studied the phase transition of CsPbI₃ by using high-resolution *in situ* synchrotron X-ray diffraction. They synthesized \delta-CsPbI3 at room temperature and converted it to α -CsPbI₃ [cubic $Pm\bar{3}m$; a = 6.29658(2) Å] by heating at around 600 K. On further heating, the α-CsPbI₃ phase remained structurally unchanged before melting at 753 K.58 The powder XRD patterns (Fig. 3b) show that δ -CsPbI₃ changes to α -CsPbI₃ in one step, but upon cooling, α-CsPbI₃ slowly returns to δ-CsPbI₃ form through the β-phase [tetragonal P4/mbm; a =8.82690(3) Å, b = 8.82690(3) Å, and c = 6.29902(3) Å] at 533 K and the γ -phase [orthorhombic *Pbnm*; a = 8.61979(6) Å, b =8.85176(6) Å, and c = 12.50128(7) Å] at 448 K. The γ -phase transforms into the δ -phase within \sim 48 h on further cooling to room temperature,12 or on exposure to humid air.68 The room temperature formation energy of α -CsPbI₃ is much larger (~14.1 kJ mol⁻¹) than that for the δ -phase as confirmed by calorimetric measurements.13 Therefore, the α-phase (even if it is successfully synthesized at ambient temperature) is difficult to store under ambient conditions. Sutton et al.65 calculated the formation energies of the CsPbI₃ polymorphs and found that the δ-phase has the largest formation enthalpy (Fig. 3c). The authors have further reported that rapid cooling leads to the yphase, while slow cooling (or on storing samples under ambient conditions, even in a vacuum) leads to the formation of the δ phase. This indicates that the former compound is a kinetic product, while the latter compound is a thermodynamic product.68 Although the perovskite phase of CsPbI₃ is metastable at room temperature, it has a small energy barrier before transitioning to the δ -phase and can therefore be realized at room temperature.⁶⁹⁻⁷² By using the energy barrier, Steele et al.⁷³ kinetically trapped the black phase CsPbI₃ at room temperature through rapid cooling its strain induced thin film. Straus et al.68 showed that bulk γ -CsPbI₃ crystals can be obtained at room temperature by rapid quenching the reaction in an ice-water bath and stored for up to 39 days under dry conditions in an evacuated ampule but exposure to humid air leads to the formation of the δ -phase within a few minutes. Their singlecrystal X-ray diffraction and X-ray pair distribution function (PDF) studies suggested that γ -CsPbI₃ crystals remain structurally unchanged in the 100-295 K range.74 However, the Cs cation was found disordered at two sites.74 Furthermore, water molecules lower the transition energy barrier and accelerate the conversion to the δ -phase.⁷¹ Also pure γ -CsPbI₃ is reported to show PCE above 18% and retain 95% of the initial efficiency

Published on 03 2023. Downloaded by Fail Open on 23/07/25 10:34:24.



Fig. 3 (a) Temperature-driven phases of CsPbl₃. (b) Temperature-dependent PXRD patterns, showing that the transition pathways are different for heating and cooling steps. Reproduced with permission from ref. 12 copyright 2018 American Chemical Society. (c) Calculated formation energies (E_i) for CsPbl₃ polymorphs with respect to the precursor's energies. The red circles and the blue dots correspond to DFT/PBE and DFT/ LDA calculations, respectively. Reproduced with permission from ref. 65 copyright 2018 American Chemical Society. (d) Optical absorption spectra of the black and the yellow phases of CsPbl₃. Adapted with permission from ref. 66 copyright the Royal Society of Chemistry 2015. (e) Calculated band structures for the black phases of CsPbl₃ along the high-symmetry points in the Brillouin zone by using the GGA-PBE function. Reproduced with permission from ref. 67 copyright 2020 Elsevier B.V.

after 30 days.⁷⁵ Ke *et al.*⁶⁹ carried out *in situ* synchrotron XRD and Raman scattering measurements to show that γ -CsPbI₃ can be synthesized and preserved at room temperature by heating and cooling under pressure. The authors first compressed δ -CsPbI₃ to pressures of 0.1–0.6 GPa at room temperature followed by heating up to 130 °C and cooling back to room temperature at a cooling rate of 25–35 °C h⁻¹. Eperon *et al.*⁶⁶ trapped α -CsPbI₃ at room temperature by carrying out device fabrication in an inert atmosphere, which showed 10% higher PCE compared with the other perovskite polymorphs.

Due to reduced structural dimensionality, the optoelectronic properties of δ -CsPbI₃ differ completely from those of the α -, β -

or γ-phases (collectively called the black phase). The optical absorption spectra (Fig. 3d) show that the black CsPbI₃ can absorb below ~715 nm (~1.73 eV), while the yellow CsPbI₃ can only absorb below ~440 nm (~2.82 eV).⁶⁶ This essentially indicates that the yellow phase is photoinactive from a solar cell point of view. Within the black phases, the bandgap marginally increases from α-CsPbI₃ ($E_g = 1.63 \text{ eV}$),⁶⁸ β-CsPbI₃ ($E_g = 1.68 \text{ eV}$),⁷⁶ and γ-CsPbI₃ ($E_g = 1.72 \text{ eV}$)⁶⁸ depending on the degree of PbI₆ octahedral tilting compared with the highest symmetry α-CsPbI₃. Furthermore, the bandgaps of these phases remain direct (Fig. 3e).⁶⁷

5.2. Methylammonium lead triiodide (MAPbI₃)

Just like CsPbI₃, MAPbI₃ also exhibits symmetry lowering phase transitions upon decreasing temperature (Fig. 4a): the α -phase (cubic $Pm\bar{3}m$) to the β -phase (tetragonal *I4/mbm*) to the γ -phase (orthorhombic *Pnma*) in which the MA⁺ cation is fully

disordered, partially disorder and fully ordered, respectively. The α - to β -phase transition is a consequence of one out-ofphase octahedral tilt (Glazer's tilt notation $a^0a^0c^-$), which is opposite in direction compared to the corresponding transition of CsPbI₃. All Pb–I–Pb bond angles in the α -phase are necessarily 180°. Owing to C_{3V} symmetry, the MA⁺ cation is expected to be orientationally disordered in cubic symmetry, with its C-N bond axis oriented along any of the three (100) directions.^{18,77} Different studies have pointed out that the I···H hydrogen bonds play an important role in the MA⁺ dynamics, thereby in the phase transition of MAPbI3.78,79 Therefore, the orderdisorder dynamics of the MA⁺ cation has been widely studied using a range of techniques such as solid-state NMR spectroscopy,10,80 dielectric spectroscopy,81 and quasi-elastic neutron scattering measurements.^{82–84} The β - to γ -phase transition is a consequence of two simultaneous in-phase and out-of-phase tilts (Glazer's tilt notation $a^-b^+a^-$) with respect to the α -phase.



Fig. 4 (a) Temperature-driven phases of MAPbI₃. (b) Molar heat capacity at constant pressure, showing two phase transitions. Reproduced with permission from ref. 87 copyright 1990 Pergamon Press plc. (c) Variable temperature ²H NMR spectra of CH₃ND₃PbI₃, acquired at 11.75 T. Reproduced with permission from ref. 88 copyright 2018 American Chemical Society. (d) DSC plots with the vertical line (blue) showing the room temperature and the hachured region (red) showing the solar cell working temperature range. (e) Temperature dependent molar heat capacity of MAPbI₃ pellets annealed at 523 K for different times. Adapted with permission from ref. 91 copyright the Royal Society of Chemistry 2021. (f and g) DSC plots of as-grown and thermally annealed polycrystalline samples of MAPbI₃ and CsPbI₃, respectively. Reproduced with permission from ref. 92 copyright 2023 American Chemical Society.

View Article Online Journal of Materials Chemistry A

During the phase transition, although the symmetry of $[PbI_6]^{4-}$ octahedra reduces from a perfect octahedral, it remains centrosymmetric with the Pb^{2+} metal ion at the centre of the octahedra. This indicates that the $6s^2$ lone pair electrons of the Pb^{2+} ion are stereochemically hidden which has been the subject of intense discussion, particularly in connection with the ferroelectric phase transition.^{20,56,85}

Weber in 1978 modelled the ambient temperature and pressure phase of MAPbI₃ in the cubic $Pm\bar{3}m$ space group (a = 6.27 Å) by using powder X-ray diffraction.44 Its temperature driven phase transition was studied by Poglitsch and Weber in 1987 using Xray diffraction data.¹⁶ They found that it crystallizes in the αphase [cubic $Pm\bar{3}m$; a = 6.3285(5)Å] above 327.4 K, in the β -phase [tetragonal *I*4/*mcm*; a = b = 8.855(6) Å and c = 12.659(8) Å] in the 327.4–162.2 K temperature range and in the γ -phase [orthorhombic $Pna2_1$; a = 8.861(2) Å, b = 8.581(2) Å and c = 12.620(3)Å] below 162.2 K.¹⁶ Later, the γ-phase was reassigned as the orthorhombic Pnma space group due to the absence of (or weak) ferroelectric behaviour.86 The orthorhombic Pnma structure was further confirmed by Baikie et al.17 from the single crystal X-ray diffraction data. No further phase transition was observed down to a temperature of 10 K.18 These two phase transitions were subsequently supported by calorimetric (Fig. 4b)⁸⁷ and NMR spectroscopic studies (Fig. 4c).⁸⁸ The β- to γ-phase transition follows first-order thermodynamics based on heat capacity measurements,87,89 as well as time-of-flight neutron and synchrotron diffraction studies.¹⁸ The α - to β -phase transition is close to second order as the volume change is continuous.77,90 However, a thermal hysteresis observed during the heating step suggests that the nature of the transition could be of firstorder.90,91 Furthermore, the whole phase transition process was found to be reversible during cooling and heating cycles.18

The α - $\rightarrow \beta$ -phase transition of MAPbI₃ falls within the typical device working temperature range (\sim 327 K), which often leads to adverse changes in the photophysical properties, affecting functioning of the whole device. Thus, stabilizing α -MAPbI₃ at room temperature by either suppressing the transition or by shifting the transition to a lower temperature has been an active research direction. Bonadio et al.91 showed that 40 h long thermal annealing of a MAPbI₃ film at 523 K lowers the α - $\rightarrow \beta$ -phase transition temperature by 36 K, stabilizing the α -phase at room temperature. The DSC data (Fig. 4d) and the heat capacity data (Fig. 4e) clearly show the lowering of the transition temperature as a function of annealing time. In their recent thermal annealing of the polycrystalline MAPbI₃ sample,⁹² the α -phase has been stabilized even at below room temperature by a shift of 62 K in the transition temperature (Fig. 4f). Furthermore, it was observed that thermal annealing has no influence on the transition temperature for the CsPbI₃ sample (Fig. 4g), indicating that the ordered-disordered orientation of MA⁺ has a substantial role in the phase transition.⁹²

Most of the studies suggest that the tetragonal *I*4/*mcm* space group is the commonly accepted structure of MAPbI₃ at room temperature. This is a non-polar space group, thereby excluding the possibility of ferroelectric behaviour. However, MAPbI₃ has been hotly debated as a promising ferroelectric material, for which the structure needs to be polar.^{59,93-95} Indeed, there have

been many reports, suggesting polar space groups for the room temperature β -phase: tetragonal *I4cm* [a = b = 8.8438(3) Å and c= 12.6215(5) Å],^{21,59,96,97} I422 [a = b = 8.88374(18) Å and c =12.7010(3) Å].^{60,98} These polar groups exhibit group/subgroup relationships with the non-polar I4/mcm space group.⁵⁹ In these groups, the [PbI₆]⁴⁻ octahedra exhibit second order Jahn-Teller distortion [Pb–I(1) distance = 3.125 Å \times 2; Pb–I(2) distance = 3.161 Å \times 4].²¹ The orientation of the MA⁺ cation in the [PbI₃]⁻ sublattice is believed to be responsible for the polarization.93 Despite many structural studies, the symmetry assignment of MAPbI₃ to a polar or non-polar space group is still debated, and so is its ferroic behaviour.20,59,93,99,100 The challenges that crystallographers encounter are often attributed to the disordered nature of the MA⁺ cation and the extremely complicated hydrogen bonding pattern.^{101,102} Crystal twinning further makes the process challenging.26,103

Symmetry lowering of the $[PbI_3]^-$ framework (or the deviation of the Pb–I–Pb bond angle from linear 180°) influences the electronic structure and the optical bandgap which eventually influences the optoelectronic properties. First-principles calculations have suggested that the valence band maximum (VBM) of MAPbI₃ is mainly composed of Pb 6s–I 5p* antibonding MOs, while the conduction band minimum (CMB) is mainly composed of Pb 6p–I 5p MOs with less antibonding character and more nonbonding character.^{79,104} Any deviation in the Pb–I–Pb bond angle from 180° reduces the Pb–I orbital overlap and stabilizes the VBM, lowering its energy, whilst due to nonbonding character the effect of lattice distortions is less on the CBM. This leads to an increase in the energy gap between the VBM and CBM.

We note that unlike δ -CsPbI₃ (vide supra) and δ -FAPbI₃ (vide infra), non-perovskite phases of MAPbI₃ are not experimentally reported but some theoretical calculations have predicted them as stable polymorphs.¹⁰⁵⁻¹⁰⁷ Although not strictly a phase transition, there have been reports on the solvent driven reconstructive transition of MAPbI₃. Huang et al.¹⁰⁸ reported the 1D δphase to 3D α -phase transition, driven by ethylamine (CH₃NH₂) intercalation. In the δ-phase, the CH₃NH₂ molecules form weak hydrogen bonds with the MA^+ cations to form $CH_3NH_2\cdots$ CH₃NH₃⁺ dimers in the space between face-sharing [PbI₃]⁻ chains. The increased ionic radius of the dimeric cation (4.45 Å), compared to the MA⁺ cation (2.17 Å), exceeds the tolerance factor beyond 1.0, which helps stabilizing the 1D structure. The yellow δ-phase (CH₃NH₂···CH₃NH₃⁺)PbI₃ irreversibly changes to the black (CH₃NH₃⁺)PbI₃ on exposure to air for few minutes. In the presence of solvents with hydrogen bonding capability (such as water, dimethyl sulfoxide, dimethylformamide etc.), MAPbI₃ can form needle-shaped yellow 1D chain compounds MAPbI₃ \cdot S (here, S = solvent molecule).¹⁰⁹ They spontaneously convert to the black perovskite phase on removal of solvent. The transitions due to water are discussed separately in Section 7.

5.3. Formamidinium lead triiodide (FAPbI₃)

FAPbI₃ exhibits four temperature driven structures: the α -phase (cubic $Pm\bar{3}m$) at high temperature, the β -phase and the γ -phase at low temperature (both crystallize in the tetragonal P4/mbm space group) and a yellow δ -phase (hexagonal $P6_3mc$) at room



Fig. 5 (a) Temperature-driven phases of FAPbl₃. (b) Diffuse reflectance spectra of α -FAPbl₃ and δ -FAPbl₃ with the inset showing photographs of their single crystals. Adapted from ref. 111 copyright 2021 the authors under CC-BY-NC-ND 4.0. (c) Temperature dependent heat capacity of FAPbl₃ at constant pressure. Reproduced with permission from ref. 119. Copyright 2019 American Chemical Society.

temperature (Fig. 5a). The α -phase has an un-tilted structure (Glazer's notation $a^0a^0a^0$) with all Pb–I–Pb angles of 180° and a fully disordered FA⁺ cation at the A-site. The β -phase and the γ -phase exhibit one in-phase octahedral tilt (Glazer's notation $a^0a^0c^+$). The δ -phase has a one-dimensional chain structure of

face-sharing $[PbI_6]^-$ octahedra along the (001) direction and the ordered FA⁺ cation located in the interchain space.¹¹⁰ Note that the structure of δ -FAPbI₃ is strikingly different than that of the δ -CsPbI₃ phase. As shown in Fig. 5b the bandgap of δ -FAPbI₃ (E_g = 2.17 eV) is wider compared to that of α -FAPbI₃ (E_g = 1.53

eV),^{111,112} making it unsuitable for most of the optoelectronic properties, including solar cells. The α -FAPbI₃ film usually contains a δ -phase impurity, which reduces light absorption and retards device efficiency.¹¹³ It is worthwhile to note that α -FAPbI₃ is more attractive than α -MAPbI₃ and α -CsPbI₃ because its bandgap is closer to the theoretically optimal bandgap (1.4 eV) required for solar cells.

Clearly, the symmetries of α -FAPbI₃ and δ -FAPbI₃ are not corelated by a simple group/subgroup relation, meaning that the δ - $\leftrightarrow \alpha$ -phase transition is a complex process of structural reconstruction. The density functional theory (DFT) calculations have suggested that the formation enthalpy (ΔH) of δ -FAPbI₃ is \sim 70 meV per formula unit lower than that of α -FAPbI₃.²⁸ This means that the Gibbs free energy ($\Delta G = \Delta H T\Delta S$) of α -FAPbI₃ is higher. The difference between the free energies of the two phases is largely attributed to the isotropic rotation of the FA⁺ cation. At temperature higher than the δ - \rightarrow α -phase transition temperature, the increased entropy of the disordered FA⁺ cation lowers the free energy of the system, thereby stabilizing α -FAPbI₃. At low temperature, the FA⁺ cation adopts preferred orientation with lower entropy and high Gibbs free energy, favouring the δ-phase. With mixed A-cation variants, such as $MA_{1-x}FA_xPbI_3$,^{114,115} and $FA_{1-x}Cs_xPbI_3$,¹¹⁶ the α - \rightarrow δ -phase transition can be inhibited, wherein the tolerance factor and the entropic factor are in favour of α-FAPbI₃.

There have been several reports on the structural studies of the α -FAPbI₃ $\rightarrow \delta$ -FAPbI₃ transition. Stoumpos *et al.*²¹ isolated both α -FAPbI₃ and δ -FAPbI₃ phases at room temperature from solution synthesis. They modelled the former phase in the trigonal P3m1 space group [a = b = 8.9817(13) Å and c =11.006(2) Å, measured at 293 K] with a 3-fold disordered FA⁺ cation at the A-site and the latter phase in the hexagonal P63mc space group [a = b = 8.6603(14) Å and c = 7.9022(6) Å, measuredat 293 K]. The authors observed that α -FAPbI₃ forms at above 333 K and transforms slowly into δ-FAPbI₃. Weller et al.⁶¹ revisited the X-ray structure of the high temperature phase and modelled it in the cubic $Pm\bar{3}m$ space group [a = 6.3620(8) Å]which is isostructural to α -CsPbI₃ [a = 6.1769(3) Å] and α -MAPbI₃ [a = 6.315(3) Å]. They used neutron powder diffraction which is a more sensitive tool compared to X-ray diffraction for locating light atoms of the FA⁺ cation. The FA⁺ cation was found to be orientationally disordered within the [PbI₃]⁻ framework of α -FAPbI₃ and fully ordered between the [PbI₃]⁻ chains of δ -FAPbI₃. The variable temperature X-ray diffraction studies have found that δ -FAPbI₃ undergoes a sharp first-order transition to α -FAPbI₃ above 350 K, which is reversible on cooling but at a lower temperature of 290 K.28 This large thermal hysteresis has been further supported by the first-principles calculations which suggest that there is a kinetic energy barrier ($\sim 0.84 \text{ eV}$) between these two phases.^{28,110} By making use of the barrier, α -FAPbI₃ has been trapped by thermal quenching at 8.2 K,²⁸ which changes to δ -FAPbI₃ on exposure to humid air. Furthermore, applied pressure is found to accelerate the α -FAPbI₃ to δ -FAPbI₃ transition by reducing the kinetic energy barrier.³³ Reports have also suggested that water vapor can trigger the α -FAPbI₃ to δ -FAPbI₃ transition.¹¹⁷ Keshavarz et al.¹¹⁸ reported that δ-FAPbI₃ also undergoes phase transitions at 173 K and 54 K which show

negative thermal expansion (NTE) with a large NTE coefficient (α) of $-1.3 \times 10^{-4} \text{ K}^{-1}$ and $-2.5 \times 10^{-4} \text{ K}^{-1}$, respectively.

Compared to the α - $\rightarrow \delta$ -FAPbI₃ transition, there have been fewer studies on the α -FAPbI₃ $\rightarrow \beta$ -FAPbI₃ $\rightarrow \gamma$ -FAPbI₃ phase transition. Fabini *et al.*³⁰ reported that α -FAPbI₃ (cubic *Pm*3*m*; a = 6.35788 Å, measured at 299 K) undergoes a second-order transition to the β -phase [tetragonal P4/mbm; a = b =8.9227(1) Å and c = 6.3263(1) Å, measured at 200 K] below 285 K, followed by a first-order re-entrant transition to the γ -phase [tetragonal P4/mbm; a = b = 8.8752(3) Å and c = 6.2792(4) Å, measured at 100 K] below 140 K. In this study, the atomic positions of the FA⁺ cation were not found in the diffraction due to its dynamic disorder and limited sensitivity of the X-ray diffractometer. It was modelled as a single pseudo-atom with an equivalent scattering factor. For this experiment, δ -FAPbI₃ was synthesized at room temperature, and packed and sealed in a Kapton capillary, and converted to α -FAPbI₃ by *in situ* heating up to 490 K during the X-ray data acquisition. Although the phase transition is continuous, there were discontinuous changes observed in the dielectric and the optical properties. Subsequent heat capacity measurements (Fig. 5c) on FAPbI₃ single crystals have suggested that there exists another phase, named I-phase,¹¹⁹ between β -FAPbI₃ and γ -FAPbI₃ in the temperature range of 141.2-130.2 K, but no further characterization of the I-phase was reported.

6. Pressure induced phase transition

6.1. Cesium lead triiodide (CsPbI₃)

Pressure has emerged as an effective way of tuning structures and revealing novel properties of various materials.120-122 In the recent past, tremendous research efforts have been devoted to pressure driven behaviour of lead iodide perovskites. To our knowledge, the studies on CsPbI3 are less prolific, compared to those on MAPbI3 and FAPbI3 perovskites. Ke et al.69 showed that metastable γ -CsPbI₃ can be stored at room temperature by heating yellow δ -CsPbI₃ followed by rapid quenching under an applied pressure in the 0.1 GPa to 0.6 GPa range. Yuan et al.14 reported reversible transformation of the orthorhombic δ-CsPbI₃ to the monoclinic $P2_1/m$ structure [a = 9.832(2) Å, b =4.543(1) Å, and c = 16.919(2) Å; $\beta = 90.34(3)^{\circ}$, measured at ambient temperature and 3.9 GPa pressure]. The Raman scattering studies, carried out up to ~15 GPa pressure, have suggested that the transition is reversible on decompression. Liang et al.¹⁵ reported that the room temperature δ -CsPbI₃ shows a first-order transition to the monoclinic C2/m structure [a =15.423(2) Å, b = 3.511(3) Å, c = 18.461(1) Å, and $\beta = 132.829(2)^{\circ}$ at 6.9 GPa, with complete conversion to the monoclinic phase at 18 GPa. The C2/m phase exhibits highly distorted $[PbI_6]^{4-}$ octahedra with the shortest Pb-I bond length among the known phases of CsPbI₃. The shortening of Pb-I bond length under pressure gives rise to strong hybridization of Pb 6s-I 5p* orbitals, which pushes up the energy of the VBM and decreases the bandgap energy.²⁴ The bandgap redshifts from 2.5 to 1.2 eV as pressure reaches 15 GPa. At high pressure, the bandgap closes and the insulator to metal transition occurs at 39.3 GPa.15

6.2. Methylammonium lead triiodide (MAPbI₃)

Upon compression, Ou *et al.*¹²³ observed discontinuous changes in the electric transport properties of MAPbI₃ nanorods. They found a mixed tetragonal and orthorhombic phase in the pressure range of 0.5–0.7 GPa and amorphous phase above 11.5 GPa from *in situ* synchrotron powder X-ray diffraction data, but lattice parameters and structural transition pathways were not reported as a function of applied pressure. A more comprehensive study has been carried out by Szafrański *et al.*²² from *in situ* single-crystal X-ray diffraction. The room temperature phase is fully consistent with the tetragonal *I4/mcm* symmetry which is stable up to 0.35 GPa. At higher pressure, it transforms into a ReO₃-type cubic *Im*3̄ phase [a = 12.4076(8) Å, measured at 293 K and 0.45(2) GPa] (Fig. 6a). A complete transformation into the $Im\bar{3}$ phase has been observed when the pressure is above 0.40 GPa.²³ The $Im\bar{3}$ unit cell is a bilayer 2 × 2 × 2 supercell of the high temperature α -phase.²³ The octahedra show three equal component tilts along all three axes (Glazer's tilt notation: $a^+a^+a^+$), generating two inequivalent A-sites. The space group assignment is consistent with the assignments in ref. 26 and 23 but is inconsistent with the orthorhombic symmetry reported in ref. 123. On increasing pressure, an isostructural phase transition occurs above 2.5 GPa, with discontinuous changes in the unit cell constants [a = 11.7936(13) Å, measured at 293 K and 2.71(2) GPa]. The authors measured the absorption spectra of MAPbI₃ as a function of applied pressure (Fig. 6b). The bandgap reduces upto 0.35 GPa followed by an abrupt widening and then is almost unchanged up to 2.5 GPa



Fig. 6 (a) Schematic illustration of pressure-induced phases of MAPbI₃. (b) Absorption spectra of MAPbI₃ as a function of pressure and (c) optical bandgaps *versus* pressure estimated from the spectra in (b), with the insets showing change in the colour of the crystal with pressure. Reproduced with permission from ref. 22 copyright 2016 American Chemical Society. (d) Synchrotron powder XRD patterns collected at room temperature under compression with diamond anvil cells (DACs) and (e) XRD patterns under decompression. (f) Room temperature photoluminescence spectra (right panel) and estimated bandgaps (left panel) as a function of pressure. Reproduced with permission from ref. 23 copyright 2016 Wiley-VCH Verlag GmbH & Co. KGaA, Weinheim.

(Fig. 6c). Above 3.5 GPa, the bandgap abruptly widens again, and the crystal turns black to transparent red. The bandgap variation was ascribed to the pressure-induced Pb–I bond contraction and hindered Pb–I–Pb angle bending in the compressed perovskite framework. By using a confocal μ -spectroscope, Liu *et al.*¹²⁴ measured the absorption and the emission spectra of single-crystal MAPbI₃ up to a pressure of 20.43 GPa. The bandgap first decreases from ambient pressure to 0.23 GPa due to contraction of the Pb–I bonds in the tetragonal *I4/mcm* phase. At 0.46 GPa, the bandgap suddenly increases due to the tetragonal to cubic phase transition and then increases slightly up to 2.72 GPa. After the isostructural phase transition at above 3.02 to 3.57 GPa, the bandgap further increases. This trend in the electronic structure change is broadly consistent with other reports.²²

By using in situ synchrotron X-ray diffraction data, Jiang et al.²³ showed that the room temperature tetragonal I4/mcm phase changes to cubic $Im\bar{3}$ [a = 12.4076(8) Å, measured at 0.41 GPa and room temperature] in the pressure range of 0.40-2.7 GPa, followed by a lower symmetry orthorhombic Immm phase [a = 12.0137(3) Å; b = 12.1984(8) Å; c = 11.9623(3) Å,measured at 2.75 GPa and room temperature] above 2.7 GPa. The authors found that the crystals gradually turn amorphous at a pressure above 4.0 GPa, reaching complete amorphization at 4.7 GPa and no further changes were noticed in the powder diffraction pattern up to 6.4 GPa (Fig. 6d).²³ The unit cell lengths of the Immm phase are marginally different than that of the Im $\overline{3}$ phase. During decompression, the structural changes are reversible as shown in the PXRD patterns (Fig. 6e). The sample remains partially amorphous until above 0.58 GPa and the $Im\bar{3}$ and I4/mcm phases were restored at 0.58 GPa and 0.55 GPa, respectively. The in situ photoluminescence spectra (Fig. 6f) suggest narrowing of the bandgap with increasing pressure up to ≈ 0.3 GPa, followed by bandgap widening up to 2.7 GPa. Beyond 2.7 GPa, the PL band weakens and slowly vanishes. A high pressure study on the MAPbI₃ single-crystal by Jaffe et al.¹²⁵ has suggested that the bandgap widening happens up to 60 GPa, followed by a semiconductor to metal transition due to high distortion in the structure. The structure as well as the bandgap are reversible between ambient pressure and 66 GPa.125 The bandgap widening is attributed to pressure induced lattice strain in the perovskite framework, triggered by disordering of the MA⁺ ion in the squeezed A-site cavity.

Despite substantial work on pressure-driven structural phenomena of MAPbI₃, there are considerable inconsistencies in the structures and the structural transition sequence. Kong *et al.*²⁴ carried out a high pressure single crystal X-ray diffraction study and suggested that the ambient pressure *I4/mcm* phase changes to an orthorhombic *Imm2* space group [a = 12.408(3) Å; b = 12.419(3) Å; c = 12.25(5) Å] upon increasing pressure to 0.4 GPa.²⁴ Capitani *et al.*²⁵ also reported the *Imm2* phase but at a lower pressure of 0.26 GPa which remains unchanged up to amorphization in the 2.3–3.0 GPa pressure range. They observed that the bandgap of MAPbI₃ steadily reduces up to 0.32 GPa before phase change, which is consistent with the other studies.^{22,124} The bandgap decreases due to compression of Pb–I bonds, leading to enhanced overlap between Pb 6s and I

5p orbitals. This pushes up the VBM, while the CBM, being mostly nonbonding, is insensitive to bond length or pressure. This effect reverses when the tetragonal phase changes to the orthorhombic Imm2 or cubic Im3 in which the Pb-I-Pb bond angle deviates from 180°. The deviation in the angle reduces overlap between Pb 6s and I 5p orbitals, which narrows down the width of the VBM and increases the energy gap between the VBM and CBM. Jaffe et al.²⁶ even modelled the ambient pressure phase in an orthorhombic *Fmmm* space group [a = 12.326 Å; b =12.346 Å; c = 12.897 Å; $\alpha = 88.828^{\circ}$; $\beta = 90.199^{\circ}$; $\gamma = 89.895^{\circ}$] by using high pressure synchrotron single crystal and powder X-ray diffraction. This contradicts reports about the tetragonal I4/ mcm space group reported at ambient pressure and temperature. Upon compression up to 0.6 GPa, the Fmmm structure changes to a cubic $Im\bar{3}$ structure which is consistent with other high pressure studies.^{22,23} Though the material undergoes amorphization at above ca. 2.7 GPa, some degree of crystallinity is retained up to ca. 50 GPa.26 Furthermore, the amorphous and the crystalline phases co-exist at all pressures above 4.0 GPa.^{125,126} Upon decompression, the original orthorhombic phase was recovered with a small hysteresis.

6.3. Formamidinium lead triiodide (FAPbI₃)

Under compression, the response of the FA⁺ cation is very different than that of the MA⁺ cation due to differences in their symmetries and H-bonding patterns. Liu et al.127 studied optoelectronic properties of α -FAPbI₃ under pressure up to 2.1 GPa. The authors reported shrinking of the bandgap from 1.489 eV to 1.337 eV, but they did not report any structural phase transition up to a pressure of 7.0 GPa. This finding is in sharp contrast to another study based on in situ synchrotron X-ray diffraction by Wang et al.32 who identified two new low symmetry phases of FAPbI3 under applied pressure: an orthorhombic Imm2 phase [reduced cell parameters, a = 6.2475 Å, b = 6.2450 Å, and c =6.2714 Å] at 0.3 GPa and an orthorhombic Immm phase [reduced cell parameters of a = 6.1132 Å, b = 6.0920 Å, and c = 6.1792 Å] at 1.7 GPa. The compound becomes amorphous at a pressure above 4.0 GPa. Sun et al.34 carried out high pressure single crystal X-ray diffraction study under an applied pressure of up to 0.74 GPa and suggested that cubic α-FAPbI3 changes to the tetragonal P4/mbm phase at 0.49 GPa. Importantly, they could locate the orientation of the FA⁺ cation in the tetragonal lattice with its N-N axis in the ab-plane. By using high pressure powder XRD, Jiang et al.³³ observed two phase transitions of α-FAPbI₃ (Fig. 7a–c): cubic $Pm\bar{3}m$ [a = 6.3651(1) Å, measured at ambient temperature and pressure] to tetragonal P4/mbm [a = b =8.8388(8) Å, c = 6.2732(2) Å, measured at ambient temperature and 0.46 GPa pressure] at 0.46 GPa, followed by cubic $Im\bar{3}$ phase [a = 12.0654(6) Å, measured at ambient temperature and 3.13 GPa pressure] at 3.13 GPa. These results contrast with those of an earlier study,³² suggesting $Pm\bar{3}m \rightarrow Imm2 \rightarrow Immm$ \rightarrow partially amorphous. Upon decompression, the crystalline phase was completely restored (Fig. 7d), showing reversibility of the transition process. The authors also studied that δ-FAPbI₃ shows a pressure-induced reversible transition from hexagonal $P6_3mc$ to an orthorhombic $Cmc2_1$ phase between 1.26 and



Fig. 7 (a) Schematic illustration of pressure-induced phases of α -FAPbI₃. (b and c) Powder XRD patterns of α -FAPbI₃ under applied pressure. (d) Powder XRD patterns measured during decompression. Reproduced with permission from ref. 33 copyright 2018 American Chemical Society.

1.73 GPa pressure. Note that the transition pressures for $FAPbI_3$ are slightly higher than those of $MAPbI_3$ (Table 1), indicating that the former compound is structurally more stable. This arises from better packing of the FA^+ cation in the A-site cavity as indicated by the tolerance factor as well as its enhanced H-bonding capability.¹²⁸

7. Decomposition under humid conditions

Irrespective of the phase, lead iodide perovskites decompose into constituents or transform into low dimensional structures in moist air which is detrimental to their applications. Many studies have shown that water present in humid air reacts with perovskites and accelerates their decomposition.^{129–133} A general belief is that decomposition by humidity occurs within a condensed liquid phase on the grain/crystal surface. The studies are more prolific in the case of MAPbI₃ and various pathways of decomposition have been proposed in different studies.¹³³⁻¹³⁵ In this section we shall focus on those studies which report the structural aspect. From single crystal X-ray diffraction data, Leguy *et al.*¹³⁶ showed that MAPbI₃ forms two hydrated phases upon exposure to water vapour at room temperature MAPbI₃·H₂O (pale yellow and needle-shaped crystals) and MA₄PbI₆·2H₂O (Fig. 8). With liquid water, decomposition into PbI₂ and MAI takes place. The monohydrate, structurally reported first by Hao *et al.*¹⁰⁹ consists of 1-D [PbI₃]⁻ chains separated by zig-zag chains of hydrogen bonded MA⁺ cations and water molecules. The [PbI₃]⁻ chain adopts a "double chain" structure and is reminiscent of δ -CsPbI₃. It



Fig. 8 X-ray structures of MAPbl₃ and its hydrated products, with the hydrogen bonded (green dashed bonds) MA⁺ cation and water molecules.

crystallizes in the monoclinic $P2_1/m$ space group [a = 4.5760(2)]Å, b = 25.4930(17) Å, and c = 12.2186(6) Å; $\beta = 96.065(4)$ °]. The dihydrate has a 0-D structure with isolated [PbI₆]⁴⁻ octahedra separated by a hydrogen bonded $[(MA \cdots H_2 O \cdots MA)_2]^{4+}$ dimeric cation.¹³⁷ It crystallizes in the monoclinic $P2_1/n$ space group [a =10.3937(3) Å, b = b 11.3055(2) Å, and c = 10.5519(2) Å; $\beta =$ 91.2981(13)°].138 Although monohydrates and dihydrates coexist on prolonged exposure to humid air, they essentially form via a two-step sequential process as shown in Scheme 1.136 The second step leads to irreversible formation of PbI₂ with the loss of MAI and H₂O. The monohydrate is an intermediate which either changes back to black MAPbI₃ by losing the crystalline water upon exposure to dry nitrogen/applying vacuum or further allows water infiltration to form MA₄PbI₆·2H₂O. Computational studies have shown that PbI₂ is thermodynamically more stable than MAPbI₃ but it has an activation energy barrier which is lowered upon interaction with water molecules.139 On exposure to low humidity, the perovskite phase of MAPbI₃ remains stable for several weeks and shows only a slight decrease in the device performance but on long exposure to high humidity or liquid water leads to irreversible decomposition of MAPI₃ into MAI and PbI₂ species, which then dissolve in excess water. Unlike MAPbI₃, FAPbI₃ is found to be more stable under ambient conditions. At low relative humidity (<50%), the δ-phase shows no signs of decomposition, but on dipping in liquid water it decomposes into yellow PbI2 and other constituents.140



Scheme 1 Chemistry of MAPbl₃ perovskite with water.

MAPbI₃ can also decompose into PbI₂ and other constituents on exposure to oxygen and light.¹⁴¹ The decomposition is induced by oxygen insertion at the iodide vacancies followed by the formation of superoxide (O_2^-) species in the presence of light. The photogenerated superoxide can deprotonate the methylammonium cation (CH₃NH₃⁺) of photo-excited MAPbI₃ and decompose it into PbI₂, water, methylamine and iodine.

8. Concluding remarks

Lead halide perovskites have come to the fore very recently and have opened an entirely new avenue of phase transition materials. In this article, we have reviewed temperature and pressure induced structures of three conventional lead iodide perovskites, CsPbI₃, MAPbI₃ and FAPbI₃. They show transitions to higher symmetry on heating and lower symmetry on compressing. There are diverse views about the number of phases, the phase transition pathways and important physical properties including electronic and ferroelectric. Tables 1-3 list several of such intricacies in the lattice symmetries. The complexity is more in the hybrid systems primarily due to dynamic disorder in the molecular lattice of organic amine cations. Such local disorders are challenging to find through Bragg scattering. Thus, phase transitions in these perovskites deserve continued studies. X-ray diffraction needs to be employed in combination with other highly sensitive techniques such as neutron scattering, pair distribution function (PDF), time-resolved spectroscopies (such as NMR, infrared and Raman), and highresolution electron microscope imaging to fully unravel structures.

In addition to revisiting the phase transition pathways and the phase structures, there are plenty of new directions waiting to be explored. Firstly, it is reasonable to ask whether we already know all polymorphs of these perovskites that are possible? To answer this, investigations in a wider range of temperature and pressure need to be carried out. Furthermore, studies need to be carried out to explore simultaneous effects of temperature and pressure. Under combined P-T effect, new structures with novel phenomena can be expected. Computation tools can help in this aspect by predicting unknown phases. Secondly, the phase transitions often show large entropy change at transition temperatures, thus leading to a large caloric effect which might make them potential candidates for heat storage. Study of the caloric effect presents an important avenue for future studies. Thirdly, we note that the conventional iodide perovskites that have been reported until now are all monometallic, meaning that they contain only one type of metal ion at the B-site (such as Pb²⁺ or Sn²⁺). Incorporation of other metal ions offers a new direction for chemical tunability in addition to structures. A very recent report on the iodide double perovskite Cs₂NaLnI₆ $(Ln = Ce, Nd, Gd, Tb, Dy)^{142}$ family is one such example, presenting an opportunity to develop non-toxic (or less toxic) alternatives to lead iodide perovskites for magnetic and optoelectronic applications.

Conflicts of interest

There are no conflicts to declare.

Acknowledgements

PV acknowledges the Science & Engineering Research Board (SERB) of the Govt. of India for a Ramanujan Fellowship (Award No. RJF/2020/000106).

References

- 1 C. N. R. Rao, Annu. Rev. Phys. Chem., 1989, 40, 291-326.
- 2 C. N. R. Rao and A. K. Ganguli, *Chem. Soc. Rev.*, 1995, **24**, 1–7.
- 3 G. E. Eperon, S. D. Stranks, C. Menelaou, M. B. Johnston, L. M. Herz and H. J. Snaith, *Energy Environ. Sci.*, 2014, 7, 982–988.
- 4 A. Kojima, K. Teshima, Y. Shirai and T. Miyasaka, J. Am. Chem. Soc., 2009, **131**, 6050–6051.
- 5 A. K. Jena, A. Kulkarni and T. Miyasaka, *Chem. Rev.*, 2019, **119**, 3036–3103.
- 6 C. C. Stoumpos and M. G. Kanatzidis, *Acc. Chem. Res.*, 2015, **48**, 2791–2802.
- 7 J. Even, M. Carignano and C. Katan, *Nanoscale*, 2016, 8, 6222–6236.
- 8 D. H. Fabini, T. A. Siaw, C. C. Stoumpos, G. Laurita, D. Olds, K. Page, J. G. Hu, M. G. Kanatzidis, S. Han and R. Seshadri, *J. Am. Chem. Soc.*, 2017, **139**, 16875–16884.
- 9 E. M. Mozur and J. R. Neilson, *Annu. Rev. Mater. Res.*, 2021, **51**, 269–291.
- 10 A. Mishra, M. A. Hope, M. Grätzel and L. Emsley, *J. Am. Chem. Soc.*, 2023, 145, 978–990.
- 11 Y. An, J. Hidalgo, C. A. R. Perini, A.-F. Castro-Méndez, J. N. Vagott, K. Bairley, S. Wang, X. Li and J.-P. Correa-Baena, ACS Energy Lett., 2021, 6, 1942–1969.

- 12 A. Marronnier, G. Roma, S. Boyer-Richard, L. Pedesseau, J.-M. Jancu, Y. Bonnassieux, C. Katan, C. C. Stoumpos, M. G. Kanatzidis and J. Even, *ACS Nano*, 2018, **12**, 3477– 3486.
- 13 B. Wang, N. Novendra and A. Navrotsky, J. Am. Chem. Soc., 2019, 141, 14501–14504.
- 14 G. Yuan, S. Qin, X. Wu, H. Ding and A. Lu, *Phase Transitions*, 2018, **91**, 38–47.
- 15 Y. Liang, X. Huang, Y. Huang, X. Wang, F. Li, Y. Wang, F. Tian, B. Liu, Z. X. Shen and T. Cui, *Adv. Sci.*, 2019, 6, 1900399.
- 16 A. Poglitsch and D. Weber, *J. Chem. Phys.*, 1987, **87**, 6373–6378.
- 17 T. Baikie, Y. Fang, J. M. Kadro, M. Schreyer, F. Wei, S. G. Mhaisalkar, M. Graetzel and T. J. White, *J. Mater. Chem. A*, 2013, 1, 5628–5641.
- 18 P. S. Whitfield, N. Herron, W. E. Guise, K. Page, Y. Q. Cheng, I. Milas and M. K. Crawford, *Sci. Rep.*, 2016, 6, 35685.
- 19 M. T. Weller, O. J. Weber, P. F. Henry, A. M. Di Pumpo and T. C. Hansen, *Chem. Commun.*, 2015, **51**, 4180–4183.
- 20 G. Sharada, P. Mahale, B. P. Kore, S. Mukherjee, M. S. Pavan, C. De, S. Ghara, A. Sundaresan, A. Pandey, T. N. G. Row and D. D. Sarma, *J. Phys. Chem. Lett.*, 2016, 7, 2412–2419.
- 21 C. C. Stoumpos, C. D. Malliakas and M. G. Kanatzidis, *Inorg. Chem.*, 2013, **52**, 9019–9038.
- 22 M. Szafrański and A. Katrusiak, J. Phys. Chem. Lett., 2016, 7, 3458–3466.
- 23 S. Jiang, Y. Fang, R. Li, H. Xiao, J. Crowley, C. Wang, T. J. White, W. A. Goddard III, Z. Wang, T. Baikie and J. Fang, Angew. Chem., Int. Ed., 2016, 55, 6540–6544.
- 24 L. Kong, G. Liu, J. Gong, Q. Hu, R. D. Schaller, P. Dera,
 D. Zhang, Z. Liu, W. Yang, K. Zhu, Y. Tang, C. Wang,
 S.-H. Wei, T. Xu and H. Mao, *Proc. Natl. Acad. Sci. U. S. A.*,
 2016, **113**, 8910–8915.
- 25 F. Capitani, C. Marini, S. Caramazza, P. Postorino, G. Garbarino, M. Hanfland, A. Pisanu, P. Quadrelli and L. Malavasi, *J. Appl. Phys.*, 2016, **119**, 185901.
- 26 A. Jaffe, Y. Lin, C. M. Beavers, J. Voss, W. L. Mao and H. I. Karunadasa, *ACS Cent. Sci.*, 2016, **2**, 201–209.
- 27 Q. Han, S.-H. Bae, P. Sun, Y.-T. Hsieh, Y. M. Yang, Y. S. Rim,
 H. Zhao, Q. Chen, W. Shi, G. Li and Y. Yang, *Adv. Mater.*,
 2016, 28, 2253–2258.
- 28 T. Chen, B. J. Foley, C. Park, C. M. Brown, L. W. Harriger, J. Lee, J. Ruff, M. Yoon, J. J. Choi and S.-H. Lee, *Sci. Adv.*, 2022, 2, e1601650.
- 29 T. Chen, W.-L. Chen, B. J. Foley, J. Lee, J. P. C. Ruff, J. Y. P. Ko, C. M. Brown, L. W. Harriger, D. Zhang, C. Park, M. Yoon, Y.-M. Chang, J. J. Choi and S.-H. Lee, *Proc. Natl. Acad. Sci. U. S. A.*, 2017, **114**, 7519–7524.
- 30 D. H. Fabini, C. C. Stoumpos, G. Laurita, A. Kaltzoglou, A. G. Kontos, P. Falaras, M. G. Kanatzidis and R. Seshadri, *Angew. Chem., Int. Ed.*, 2016, 55, 15392–15396.
- 31 O. J. Weber, D. Ghosh, S. Gaines, P. F. Henry, A. B. Walker, M. S. Islam and M. T. Weller, *Chem. Mater.*, 2018, **30**, 3768– 3778.

Published on 03 2023. Downloaded by Fail Open on 23/07/25 10:34:24.

- 32 P. Wang, J. Guan, D. T. K. Galeschuk, Y. Yao, C. F. He, S. Jiang, S. Zhang, Y. Liu, M. Jin, C. Jin and Y. Song, *J. Phys. Chem. Lett.*, 2017, 8, 2119–2125.
- 33 S. Jiang, Y. Luan, J. I. Jang, T. Baikie, X. Huang, R. Li,
 F. O. Saouma, Z. Wang, T. J. White and J. Fang, *J. Am. Chem. Soc.*, 2018, 140, 13952–13957.
- 34 S. Sun, Z. Deng, Y. Wu, F. Wei, F. H. Isikgor, F. Brivio, M. W. Gaultois, J. Ouyang, P. D. Bristowe, A. K. Cheetham and G. Kieslich, *Chem. Commun.*, 2017, 53, 7537–7540.
- 35 C. N. R. Rao, Acc. Chem. Res., 1984, 17, 83-89.
- 36 C. N. R. Rao and K. J. Rao, *Phase Transitions in Solids: An* Approach to the Study of the Chemistry and Physics of Solids, McGraw-Hill, New York, 1978.
- 37 P. Ehrenfest, Proc. R. Acad. Sci. Amsterdam, 1933, 36, 153– 157.
- 38 M. Stavinoha, J. A. Cooley, S. G. Minasian, T. M. McQueen, S. M. Kauzlarich, C.-L. Huang and E. Morosan, *Phys. Rev. B*, 2018, 97, 195146.
- 39 F. Guillou, A. K. Pathak, D. Paudyal, Y. Mudryk, F. Wilhelm, A. Rogalev and V. K. Pecharsky, *Nat. Commun.*, 2018, 9, 2925.
- 40 W. Aftab, A. Usman, J. Shi, K. Yuan, M. Qin and R. Zou, *Energy Environ. Sci.*, 2021, **14**, 4268–4291.
- 41 H. L. Wells, Z. Anorg. Allg. Chem., 1893, 3, 195-210.
- 42 C. K. Møller, Nature, 1957, 180, 981-982.
- 43 C. K. Møller, Nature, 1958, 182, 1436.
- 44 D. Weber, Z. Naturforsch., B: Chem. Sci., 1978, 33, 1443-1445.
- 45 A. Maalej, Y. Abid, A. Kallel, A. Daoud, A. Lautié and F. Romain, *Solid State Commun.*, 1997, **103**, 279–284.
- 46 R. J. Worhatch, H. Kim, I. P. Swainson, A. L. Yonkeu and S. J. L. Billinge, *Chem. Mater.*, 2008, 20, 1272–1277.
- 47 I. B. Koutselas, L. Ducasse and G. C. Papavassiliou, *J. Phys.:* Condens.Matter, 1996, **8**, 1217–1227.
- 48 K. Yamada, H. Kawaguchi, T. Matsui, T. Okuda and S. Ichiba, *Bull. Chem. Soc. Jpn.*, 1990, **63**, 2521–2525.
- 49 K. V Saurav and P. Vishnoi, in *Energy Materials*, World Scientific, 2022, pp. 141–174.
- 50 H. Min, D. Y. Lee, J. Kim, G. Kim, K. S. Lee, J. Kim, M. J. Paik, Y. K. Kim, K. S. Kim, M. G. Kim, T. J. Shin and S. Il Seok, *Nature*, 2021, **598**, 444–450.
- 51 J. Jeong, M. Kim, J. Seo, H. Lu, P. Ahlawat, A. Mishra, Y. Yang, M. A. Hope, F. T. Eickemeyer, M. Kim, Y. J. Yoon, I. W. Choi, B. P. Darwich, S. J. Choi, Y. Jo, J. H. Lee, B. Walker, S. M. Zakeeruddin, L. Emsley, U. Rothlisberger, A. Hagfeldt, D. S. Kim, M. Grätzel and J. Y. Kim, *Nature*, 2021, **592**, 381–385.
- 52 J. Y. Kim, J.-W. Lee, H. S. Jung, H. Shin and N.-G. Park, *Chem. Rev.*, 2020, **120**, 7867–7918.
- 53 D. V. S. Muthu, P. Teredesai, S. Saha, Suchitra, U. V. Waghmare, A. K. Sood and C. N. R. Rao, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 2015, 91, 224308.
- 54 V. M. Goldschmidt, Naturwissenschaften, 1926, 14, 477-485.
- 55 L. Pauling, J. Am. Chem. Soc., 1927, 49, 765-790.
- 56 Y. Fu, S. Jin and X.-Y. Zhu, Nat. Rev. Chem, 2021, 5, 838-852.
- 57 S. Bera, A. Saha, S. Mondal, A. Biswas, S. Mallick, R. Chatterjee and S. Roy, *Mater. Adv.*, 2022, 3, 5234–5247.

- 58 D. M. Trots and S. V Myagkota, J. Phys. Chem. Solids, 2008, 69, 2520–2526.
- 59 J. Breternitz, F. Lehmann, S. A. Barnett, H. Nowell and S. Schorr, *Angew. Chem., Int. Ed.*, 2020, **59**, 424–428.
- 60 A. Arakcheeva, D. Chernyshov, M. Spina, L. Forró and E. Horváth, *Acta Crystallogr., Sect. B: Struct. Sci., Cryst. Eng. Mater.*, 2016, **72**, 716–722.
- 61 M. T. Weller, O. J. Weber, J. M. Frost and A. Walsh, *J. Phys. Chem. Lett.*, 2015, **6**, 3209–3212.
- 62 P. Vishnoi, R. Seshadri and A. K. Cheetham, J. Phys. Chem. C, 2021, 125, 11756–11764.
- 63 X. Ding, M. Cai, X. Liu, Y. Ding, X. Liu, Y. Wu, T. Hayat, A. Alsaedi and S. Dai, ACS Appl. Mater. Interfaces, 2019, 11, 37720–37725.
- 64 J. Hieulle, X. Wang, C. Stecker, D.-Y. Son, L. Qiu, R. Ohmann, L. K. Ono, A. Mugarza, Y. Yan and Y. Qi, *J. Am. Chem. Soc.*, 2019, **141**, 3515–3523.
- 65 R. J. Sutton, M. R. Filip, A. A. Haghighirad, N. Sakai,
 B. Wenger, F. Giustino and H. J. Snaith, ACS Energy Lett., 2018, 3, 1787–1794.
- 66 G. E. Eperon, G. M. Paternò, R. J. Sutton, A. Zampetti, A. A. Haghighirad, F. Cacialli and H. J. Snaith, *J. Mater. Chem. A*, 2015, 3, 19688–19695.
- 67 M. A. Fadla, B. Bentria, T. Dahame and A. Benghia, *Phys. B*, 2020, **585**, 412118.
- 68 D. B. Straus, S. Guo and R. J. Cava, J. Am. Chem. Soc., 2019, 141, 11435–11439.
- 69 F. Ke, C. Wang, C. Jia, N. R. Wolf, J. Yan, S. Niu, T. P. Devereaux, H. I. Karunadasa, W. L. Mao and Y. Lin, *Nat. Commun.*, 2021, **12**, 461.
- 70 X. Kong, K. Shayan, S. Hua, S. Strauf and S. S. Lee, *ACS Appl. Energy Mater.*, 2019, **2**, 2948–2955.
- 71 J. A. Steele, V. Prakasam, H. Huang, E. Solano, D. Chernyshov, J. Hofkens and M. B. J. Roeffaers, *J. Am. Chem. Soc.*, 2021, 143, 10500–10508.
- 72 Z.-G. Li, M. Zacharias, Y. Zhang, F. Wei, Y. Qin, Y.-Q. Yang, L.-C. An, F.-F. Gao, W. Li, J. Even and X.-H. Bu, ACS Energy Lett., 2023, 8, 3016–3024.
- 73 J. A. Steele, H. Jin, I. Dovgaliuk, R. F. Berger, T. Braeckevelt, H. Yuan, C. Martin, E. Solano, K. Lejaeghere, S. M. J. Rogge, C. Notebaert, W. Vandezande, K. P. F. Janssen, B. Goderis, E. Debroye, Y.-K. Wang, Y. Dong, D. Ma, M. Saidaminov, H. Tan, Z. Lu, V. Dyadkin, D. Chernyshov, V. Van Speybroeck, E. H. Sargent, J. Hofkens and M. B. J. Roeffaers, *Science*, 2019, 365, 679–684.
- 74 D. B. Straus, S. Guo, A. M. M. Abeykoon and R. J. Cava, *Adv. Mater.*, 2020, **32**, 2001069.
- 75 L. Duan, H. Zhang, M. Liu, M. Grätzel and J. Luo, ACS Energy Lett., 2022, 7, 2911–2918.
- 76 W. Yong, D. M. Ibrahim, L. K. Ono, Z. Taiyang, K. Miao,
 L. Yawen, Z. Lijun, W. Xingtao, Y. Yingguo, G. Xingyu,
 Q. Yabing, G. Michael and Z. Yixin, *Science*, 2019, 365, 591–595.
- 77 Y. Kawamura, H. Mashiyama and K. Hasebe, *J. Phys. Soc. Jpn.*, 2002, **71**, 1694–1697.
- 78 K. P. Ong, T. W. Goh, Q. Xu and A. Huan, J. Phys. Chem. Lett., 2015, 6, 681–685.

- 79 J.-H. Lee, N. C. Bristowe, J. H. Lee, S.-H. Lee, P. D. Bristowe,
 A. K. Cheetham and H. M. Jang, *Chem. Mater.*, 2016, 28, 4259–4266.
- 80 D. J. Kubicki, D. Prochowicz, A. Hofstetter, P. Péchy, S. M. Zakeeruddin, M. Grätzel and L. Emsley, *J. Am. Chem. Soc.*, 2017, **139**, 10055–10061.
- 81 I. Anusca, S. Balčiūnas, P. Gemeiner, Š. Svirskas, M. Sanlialp, G. Lackner, C. Fettkenhauer, J. Belovickis, V. Samulionis, M. Ivanov, B. Dkhil, J. Banys, V. V Shvartsman and D. C. Lupascu, *Adv. Energy Mater.*, 2017, 7, 1700600.
- 82 A. M. A. Leguy, J. M. Frost, A. P. McMahon, V. G. Sakai, W. Kockelmann, C. Law, X. Li, F. Foglia, A. Walsh, B. C. O'Regan, J. Nelson, J. T. Cabral and P. R. F. Barnes, *Nat. Commun.*, 2015, 6, 7124.
- 83 T. Chen, B. J. Foley, B. Ipek, M. Tyagi, J. R. D. Copley,
 C. M. Brown, J. J. Choi and S.-H. Lee, *Phys. Chem. Chem. Phys.*, 2015, 17, 31278–31286.
- 84 B. Li, Y. Kawakita, Y. Liu, M. Wang, M. Matsuura,
 K. Shibata, S. Ohira-Kawamura, T. Yamada, S. Lin,
 K. Nakajima and S. F. Liu, *Nat. Commun.*, 2017, 8, 16086.
- 85 A. Gómez, Q. Wang, A. R. Goñi, M. Campoy-Quiles and A. Abate, *Energy Environ. Sci.*, 2019, **12**, 2537–2547.
- 86 I. P. Swainson, R. P. Hammond, C. Soullière, O. Knop and W. Massa, J. Solid State Chem., 2003, 176, 97–104.
- 87 N. Onoda-Yamamuro, T. Matsuo and H. Suga, J. Phys. Chem. Solids, 1990, 51, 1383–1395.
- 88 G. M. Bernard, R. E. Wasylishen, C. I. Ratcliffe, V. Terskikh, Q. Wu, J. M. Buriak and T. Hauger, *J. Phys. Chem. A*, 2018, 122, 1560–1573.
- 89 N. Onoda-Yamamuro, O. Yamamuro, T. Matsuo and H. Suga, *J. Phys. Chem. Solids*, 1992, **53**, 277–281.
- 90 G. Sombrio, Z. Zhang, A. Bonadio, L. S. de Oliveira, T. B. de Queiroz, F. F. Ferreira, A. Janotti and J. A. Souza, *J. Phys. Chem. C*, 2020, **124**, 10793–10803.
- 91 A. Bonadio, C. A. Escanhoela, F. P. Sabino, G. Sombrio, V. G. de Paula, F. F. Ferreira, A. Janotti, G. M. Dalpian and J. A. Souza, *J. Mater. Chem. A*, 2021, 9, 1089–1099.
- 92 A. Bonadio, F. P. Sabino, A. L. M. Freitas, M. R. Felez, G. M. Dalpian and J. A. Souza, *Inorg. Chem.*, 2023, 62, 7533–7544.
- 93 H. Röhm, T. Leonhard, M. J. Hoffmann and A. Colsmann, Energy Environ. Sci., 2017, 10, 950–955.
- 94 X.-K. Wei, N. Domingo, Y. Sun, N. Balke, R. E. Dunin-Borkowski and J. Mayer, *Adv. Energy Mater.*, 2022, **12**, 2201199.
- 95 L. M. Garten, D. T. Moore, S. U. Nanayakkara, S. Dwaraknath, P. Schulz, J. Wands, A. Rockett, B. Newell, K. A. Persson, S. Trolier-McKinstry and D. S. Ginley, *Sci. Adv.*, 2023, 5, eaas9311.
- 96 Y. Dang, Y. Liu, Y. Sun, D. Yuan, X. Liu, W. Lu, G. Liu, H. Xia and X. Tao, *CrystEngComm*, 2015, 17, 665–670.
- 97 Y. Rakita, O. Bar-Elli, E. Meirzadeh, H. Kaslasi, Y. Peleg, G. Hodes, I. Lubomirsky, D. Oron, D. Ehre and D. Cahen, *Proc. Natl. Acad. Sci. U. S. A.*, 2017, **114**, E5504–E5512.

- 98 A. Glushkova, K. Mantulnikovs, G. Giriat, K. Semeniuk, L. Forró, E. Horváth and A. Arakcheeva, *Sol. RRL*, 2019, 3, 1900044.
- 99 M. Manzi, G. Pica, M. De Bastiani, S. Kundu, G. Grancini and M. I. Saidaminov, *J. Phys. Chem. Lett.*, 2023, 14, 3535– 3552.
- 100 F. Ambrosio, F. De Angelis and A. R. Goñi, *J. Phys. Chem. Lett.*, 2022, **13**, 7731–7740.
- 101 Y. Ren, I. W. H. Oswald, X. Wang, G. T. McCandless and J. Y. Chan, *Cryst. Growth Des.*, 2016, 16, 2945–2951.
- 102 J.-H. Lee, N. C. Bristowe, P. D. Bristowe and A. K. Cheetham, *Chem. Commun.*, 2015, **51**, 6434–6437.
- 103 M. Szafrański and A. Katrusiak, *J. Phys. Chem. Lett.*, 2017, **8**, 2496–2506.
- 104 R. Prasanna, A. Gold-Parker, T. Leijtens, B. Conings, A. Babayigit, H.-G. Boyen, M. F. Toney and M. D. McGehee, *J. Am. Chem. Soc.*, 2017, **139**, 11117–11124.
- 105 A. S. Thind, X. Huang, J. Sun and R. Mishra, *Chem. Mater.*, 2017, **29**, 6003–6011.
- 106 J. A. Finkler and S. Goedecker, *Mater. Adv.*, 2023, 4, 184–194.
- 107 J. A. Flores-Livas, D. Tomerini, M. Amsler, A. Boziki, U. Rothlisberger and S. Goedecker, *Phys. Rev. Mater.*, 2018, 2, 85201.
- 108 X. Huang, R. Chen, G. Deng, F. Han, P. Ruan, F. Cheng, J. Yin, B. Wu and N. Zheng, *J. Am. Chem. Soc.*, 2020, 142, 6149–6157.
- 109 F. Hao, C. C. Stoumpos, Z. Liu, R. P. H. Chang and M. G. Kanatzidis, J. Am. Chem. Soc., 2014, 136, 16411– 16419.
- 110 M. Lai, T. Lei, Y. Zhang, J. Jin, J. A. Steele and P. Yang, *MRS Bull.*, 2021, **46**, 310–316.
- 111 M. Aebli, N. Porenta, N. Aregger and M. V Kovalenko, *Chem. Mater.*, 2021, 33, 6965–6973.
- 112 X. Chen, Y. Xia, Z. Zheng, X. Xiao, C. Ling, M. Xia, Y. Hu, A. Mei, R. Cheacharoen, Y. Rong and H. Han, *Chem. Mater.*, 2022, 34, 728–735.
- 113 S. Li, J. Xia, Z. Wen, H. Gu, J. Guo, C. Liang, H. Pan, X. Wang and S. Chen, *Adv. Sci.*, 2023, **10**, 2300056.
- 114 O. J. Weber, B. Charles and M. T. Weller, *J. Mater. Chem. A*, 2016, 4, 15375–15382.
- 115 B. Charles, J. Dillon, O. J. Weber, M. S. Islam and M. T. Weller, *J. Mater. Chem. A*, 2017, 5, 22495–22499.
- 116 Z. Li, M. Yang, J.-S. Park, S.-H. Wei, J. J. Berry and K. Zhu, *Chem. Mater.*, 2016, **28**, 284–292.
- 117 P. Raval, R. M. Kennard, E. S. Vasileiadou, C. J. Dahlman, I. Spanopoulos, M. L. Chabinyc, M. Kanatzidis and G. N. Manjunatha Reddy, ACS Energy Lett., 2022, 7, 1534– 1543.
- 118 M. Keshavarz, M. Ottesen, S. Wiedmann, M. Wharmby, R. Küchler, H. Yuan, E. Debroye, J. A. Steele, J. Martens, N. E. Hussey, M. Bremholm, M. B. J. Roeffaers and J. Hofkens, *Adv. Mater.*, 2019, **31**, 1900521.
- 119 S. Kawachi, M. Atsumi, N. Saito, N. Ohashi, Y. Murakami and J. Yamaura, *J. Phys. Chem. Lett.*, 2019, **10**, 6967–6972.
- 120 C. Zhu, X. Niu, Y. Fu, N. Li, C. Hu, Y. Chen, X. He, G. Na, P. Liu, H. Zai, Y. Ge, Y. Lu, X. Ke, Y. Bai, S. Yang,

P. Chen, Y. Li, M. Sui, L. Zhang, H. Zhou and Q. Chen, *Nat. Commun.*, 2019, **10**, 815.

- 121 D. Ghosh, A. Aziz, J. A. Dawson, A. B. Walker and M. S. Islam, *Chem. Mater.*, 2019, **31**, 4063–4071.
- 122 P. V Teredesai, A. K. Sood, D. V. S. Muthu, R. Sen,
 A. Govindaraj and C. N. R. Rao, *Chem. Phys. Lett.*, 2000,
 319, 296–302.
- 123 T. Ou, J. Yan, C. Xiao, W. Shen, C. Liu, X. Liu, Y. Han, Y. Ma and C. Gao, *Nanoscale*, 2016, **8**, 11426–11431.
- 124 C. Liu, Z. Li, L. Yang, X. Yao, H. Li, X. Liu, Y. Zhao, P. Zhu,
 T. Cui, C. Sun and Y. Bao, *J. Phys. Chem. C*, 2019, 123, 30221–30227.
- 125 A. Jaffe, Y. Lin, W. L. Mao and H. I. Karunadasa, J. Am. Chem. Soc., 2017, **139**, 4330–4333.
- 126 A. Arakcheeva, V. Svitlyk, E. Polini, L. Henry, D. Chernyshov, A. Sienkiewicz, G. Giriat, A. Glushkova, M. Kollar, B. Náfrádi, L. Forro and E. Horváth, Acta Crystallogr., Sect. B: Struct. Sci., Cryst. Eng. Mater., 2019, 75, 361–370.
- 127 G. Liu, L. Kong, J. Gong, W. Yang, H. Mao, Q. Hu, Z. Liu,
 R. D. Schaller, D. Zhang and T. Xu, *Adv. Funct. Mater.*, 2017, 27, 361–370.
- 128 A. Amat, E. Mosconi, E. Ronca, C. Quarti, P. Umari, M. K. Nazeeruddin, M. Grätzel and F. De Angelis, *Nano Lett.*, 2014, 14, 3608–3616.
- 129 J. M. Frost, K. T. Butler, F. Brivio, C. H. Hendon, M. van Schilfgaarde and A. Walsh, *Nano Lett.*, 2014, **14**, 2584–2590.
- 130 G. E. Eperon, S. N. Habisreutinger, T. Leijtens, B. J. Bruijnaers, J. J. van Franeker, D. W. deQuilettes, S. Pathak, R. J. Sutton, G. Grancini, D. S. Ginger, R. A. J. Janssen, A. Petrozza and H. J. Snaith, ACS Nano, 2015, 9, 9380–9393.
- 131 C. Müller, T. Glaser, M. Plogmeyer, M. Sendner, S. Döring, A. A. Bakulin, C. Brzuska, R. Scheer, M. S. Pshenichnikov,

W. Kowalsky, A. Pucci and R. Lovrinčić, *Chem. Mater.*, 2015, **27**, 7835–7841.

- 132 J. Yang, B. D. Siempelkamp, D. Liu and T. L. Kelly, ACS Nano, 2015, 9, 1955–1963.
- 133 T. D. Siegler, W. A. Dunlap-Shohl, Y. Meng, Y. Yang, W. F. Kau, P. P. Sunkari, C. E. Tsai, Z. J. Armstrong, Y.-C. Chen, D. A. C. Beck, M. Meilă and H. W. Hillhouse, J. Am. Chem. Soc., 2022, 144, 5552–5561.
- 134 A. Kakekhani, R. N. Katti and A. M. Rappe, *APL Mater.*, 2019, 7, 41112.
- 135 P. Toloueinia, H. Khassaf, A. S. Amin, Z. M. Tobin, S. P. Alpay and S. L. Suib, ACS Appl. Energy Mater., 2020, 3, 8240–8248.
- 136 A. M. A. Leguy, Y. Hu, M. Campoy-Quiles, M. I. Alonso,
 O. J. Weber, P. Azarhoosh, M. van Schilfgaarde,
 M. T. Weller, T. Bein, J. Nelson, P. Docampo and
 P. R. F. Barnes, *Chem. Mater.*, 2015, 27, 3397–3407.
- 137 B. R. Vincent, K. N. Robertson, T. S. Cameron and O. Knop, *Can. J. Chem.*, 1987, **65**, 1042–1046.
- 138 A. Wakamiya, M. Endo, T. Sasamori, N. Tokitoh, Y. Ogomi, S. Hayase and Y. Murata, *Chem. Lett.*, 2014, **43**, 711–713.
- E. Smecca, Y. Numata, I. Deretzis, G. Pellegrino, S. Boninelli, T. Miyasaka, A. La Magna and A. Alberti, *Phys. Chem. Chem. Phys.*, 2016, 18, 13413–13422.
- 140 J. S. Yun, J. Kim, T. Young, R. J. Patterson, D. Kim, J. Seidel, S. Lim, M. A. Green, S. Huang and A. Ho-Baillie, *Adv. Funct. Mater.*, 2018, 28, 1705363.
- 141 N. Aristidou, C. Eames, I. Sanchez-Molina, X. Bu, J. Kosco, M. S. Islam and S. A. Haque, *Nat. Commun.*, 2017, 8, 15218.
- 142 G. Kent, E. Morgan, K. R. Albanese, A. Kallistova, A. Brumberg, L. Kautzsch, G. Wu, P. Vishnoi, R. Seshadri and A. K. Cheetham, *Angew. Chem., Int. Ed.*, 2023, e202306000.